

# Challenges in Interpreting Ly $\alpha$ Absorption at High-z (impacts of an inhomogeneous UVB)

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# Reionization is Patchy!

DexM public release available at <http://www.astro.princeton.edu/~mesinger>



250 Mpc



QuickTime™ and a  
mpeg4 decompressor  
are needed to see this picture.

# Patchy Reionization--> LAEs

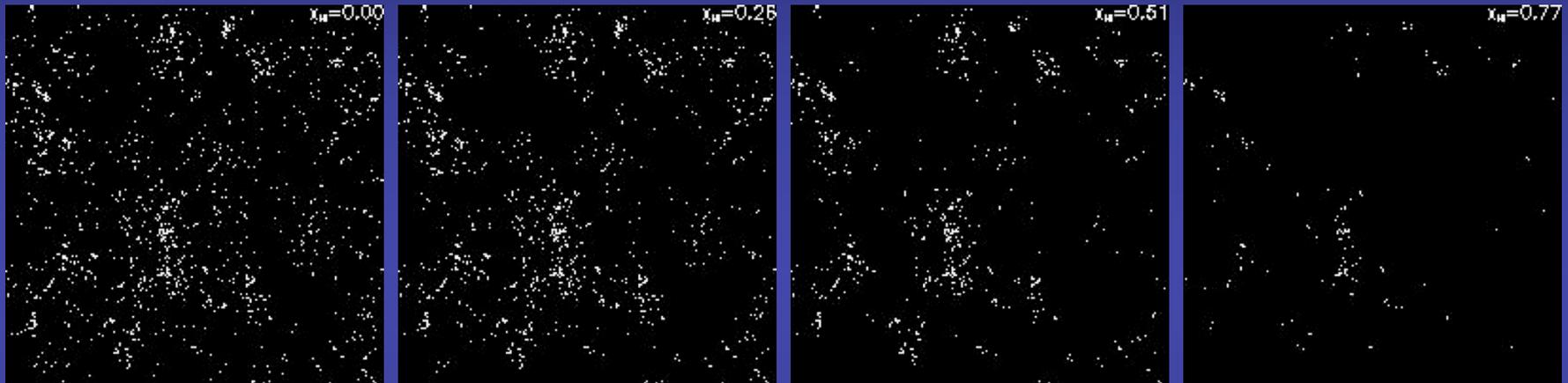
Reionization modulates the observed LAE maps

$x_{\text{HI}} \sim 0$

$x_{\text{HI}} = 0.26$

$x_{\text{HI}} = 0.51$

$x_{\text{HI}} = 0.77$



Mesinger & Furlanetto 2008b

$$M > 1.67 \times 10^{10} e^{-\tau} M_{\text{SUN}}$$

Impacts:

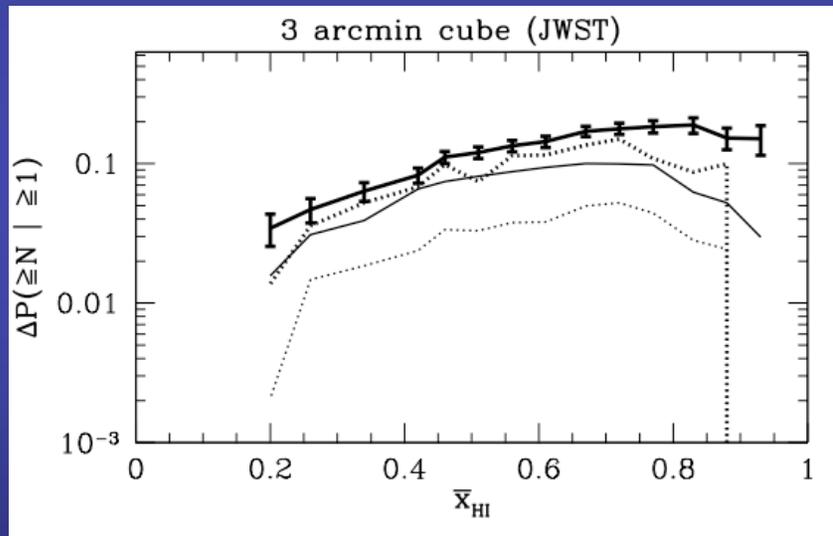
- Luminosity function
- Clustering

*Tough to do since we don't know*

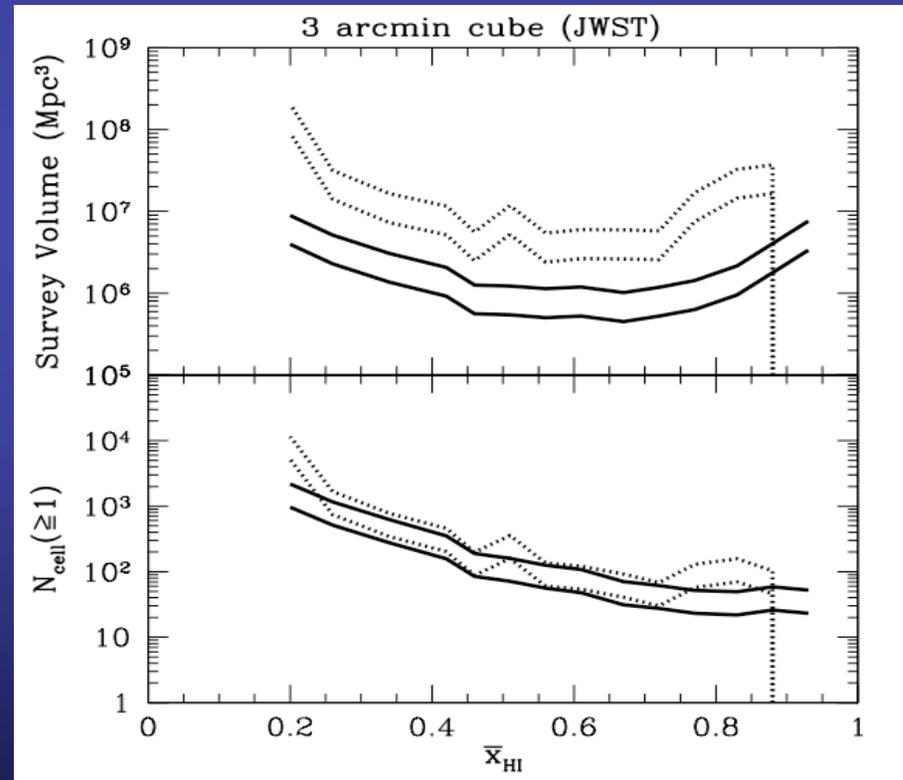
$$M_{\text{halo}} \leftrightarrow L$$

# Counts-in-Cell Statistics

- Includes higher-order, non-Gaussian corrections to clustering, unlike the commonly studied power spectrum (e.g. [McQuinn et al. 2007](#); [Iliev et al. 2008](#))
- Not very model dependent; reionization signal is separable from the evolution in structure, especially in higher-order (see [Mesinger & Furlanetto 2008b](#) for details)
- Few constraints on survey geometry; useful for follow-up



*Reionization detectable with  
< 100 fields*



# Patchy Reionization --> Ly $\alpha$ Damping Wing

QuickTime™ and a  
mpeg4 decompressor  
are needed to see this picture.

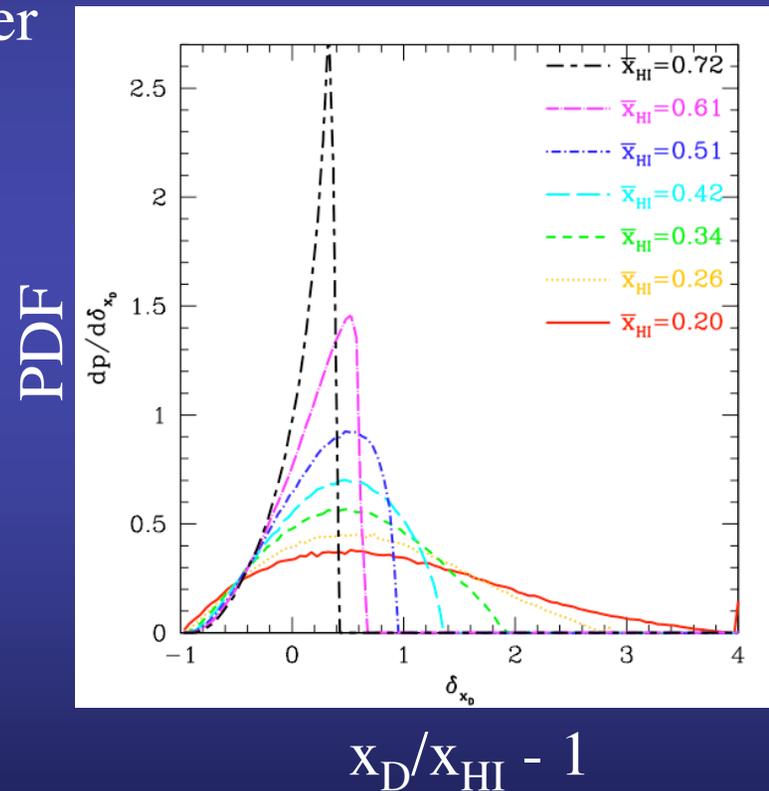
- How dangerous is it to assume a homogeneous  $x_{\text{HI}}$  or  $J_{\text{UV}}$  in damping wing studies:
  - QSOs proximity region (Mesinger & Haiman 2004; 2007)
  - GRB after disentangling DLA (Totani et al. 2006)

# Bias

- Common reasoning: absorption cross-section is flat in the wings and so is sensitive to a large path length in the IGM, so ionization structure is averaged-over
- Not flat enough! -> bias + scatter

• *constrain  $x_{\text{HI}}$  with scatter?*  
*Noise --> Signal*

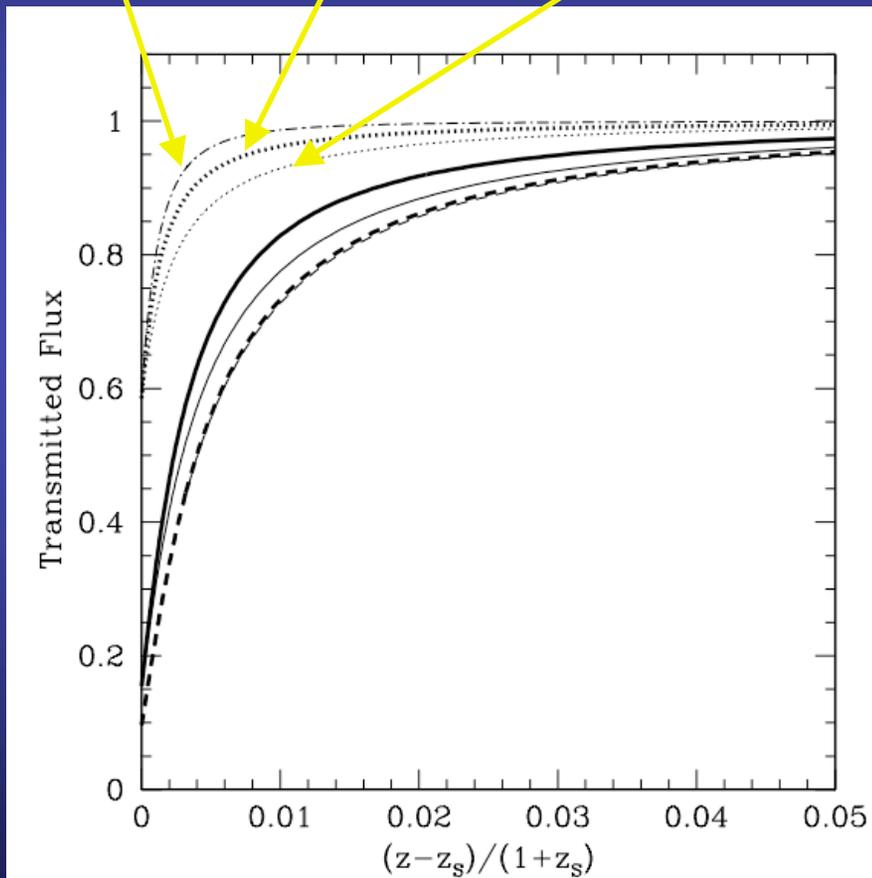
• *bias and scatter are reduced if one probes subset (e.g.  $R_s > 40$ )*



# Absorption Profile

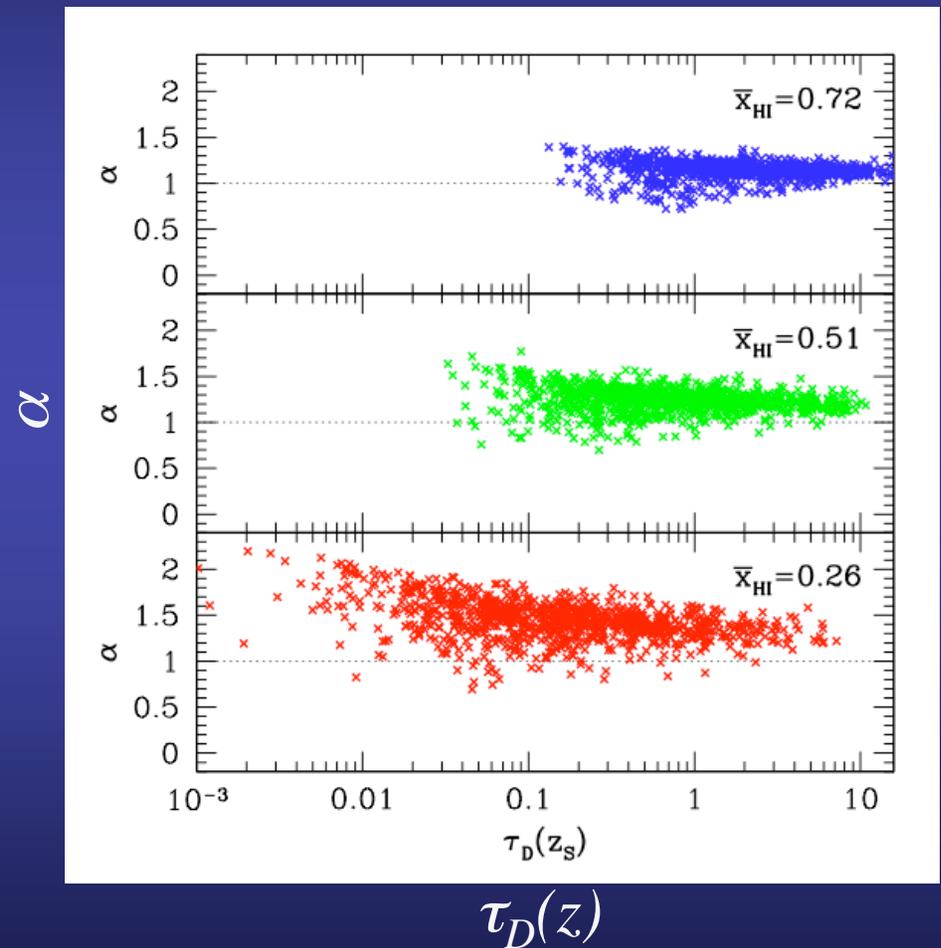
$$x_{\text{HI}}=0.1$$

DLA patchy homogeneous



Challenges in Interpreting Ly $\alpha$  at High-z

$$\tau_D(z) \propto R_{b1}^{-\alpha}$$

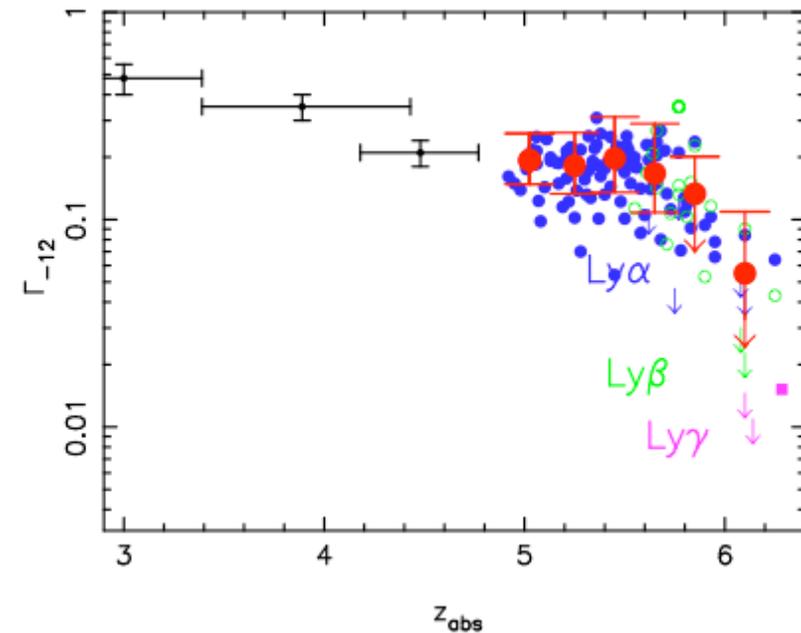
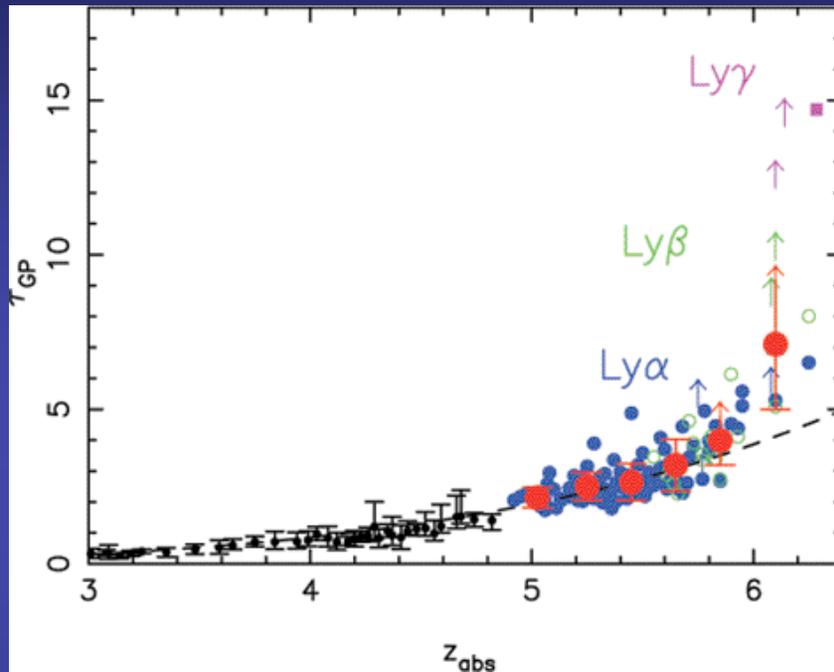


# Impact on Present Damping Wing Studies

- Not clear, however profile is more important than bias: steeper profile -> harder detection
  - > weakens upper limit from [Totani et al. 2006](#)
  - > strengthens lower limits from [Mesinger & Haiman 2004, 2007](#)
- Scatter likely causes confidence contours to degrade for all studies
- Should be redone! More sources would be nice

# Are we witnessing the end of patchiness?

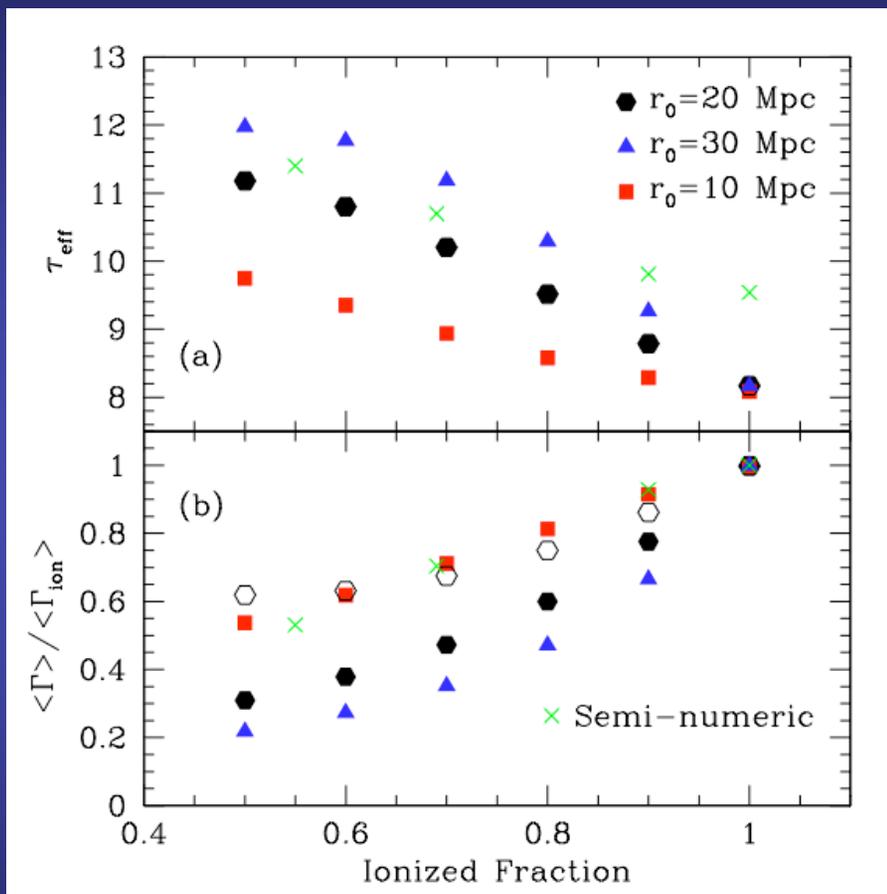
Fan+ (2006)



Note empirical objection by [Becker+ \(2007\)](#)

# Overlap does NOT directly result in a sharp UVB increase!

Furlanetto & Mesinger (2009)

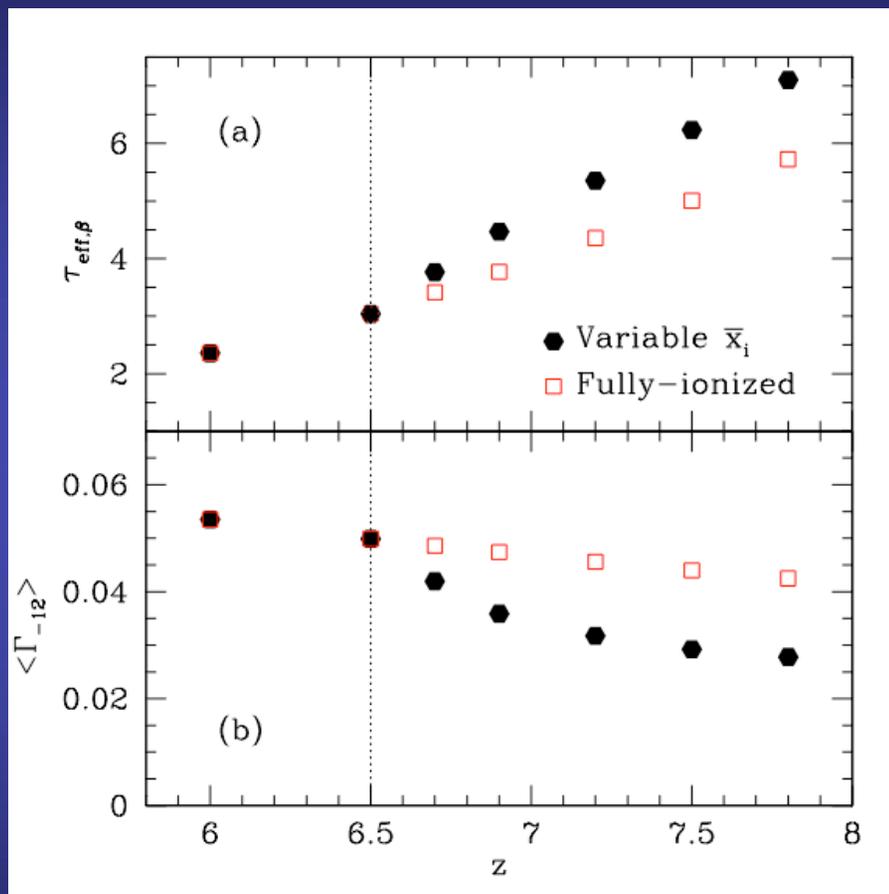


- mfp is fairly modest at these redshifts --> absorption systems regulate the UVB evolution late in reionization

constant  $z$  and mfp

# Overlap does NOT directly result in a sharp UVB increase!

Furlanetto & Mesinger (2009)



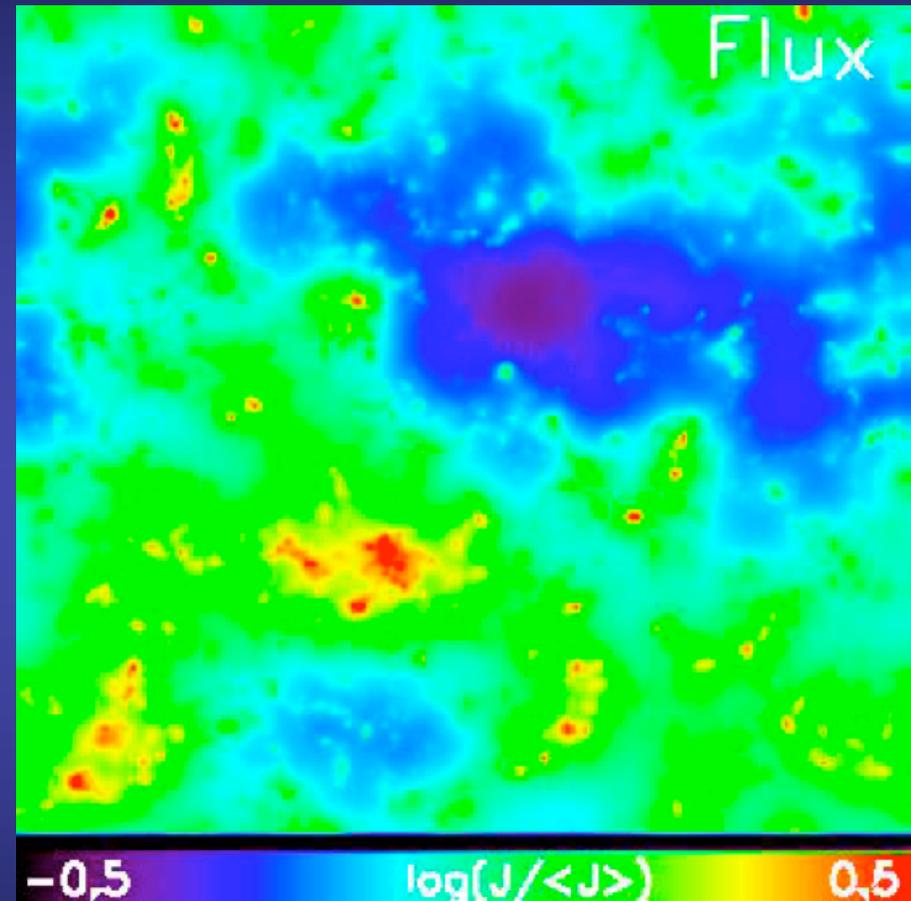
- mfp is fairly modest at these redshifts --> absorption systems regulate the UVB evolution late in reionization
- UVB increases as absorption systems are either photoevaporated or new sources appear in neutral regions --> hard to get sharp rise in UVB

simple model of  $x_{\text{HI}}$  and mfp evolution

# Patchiness even after reionization!

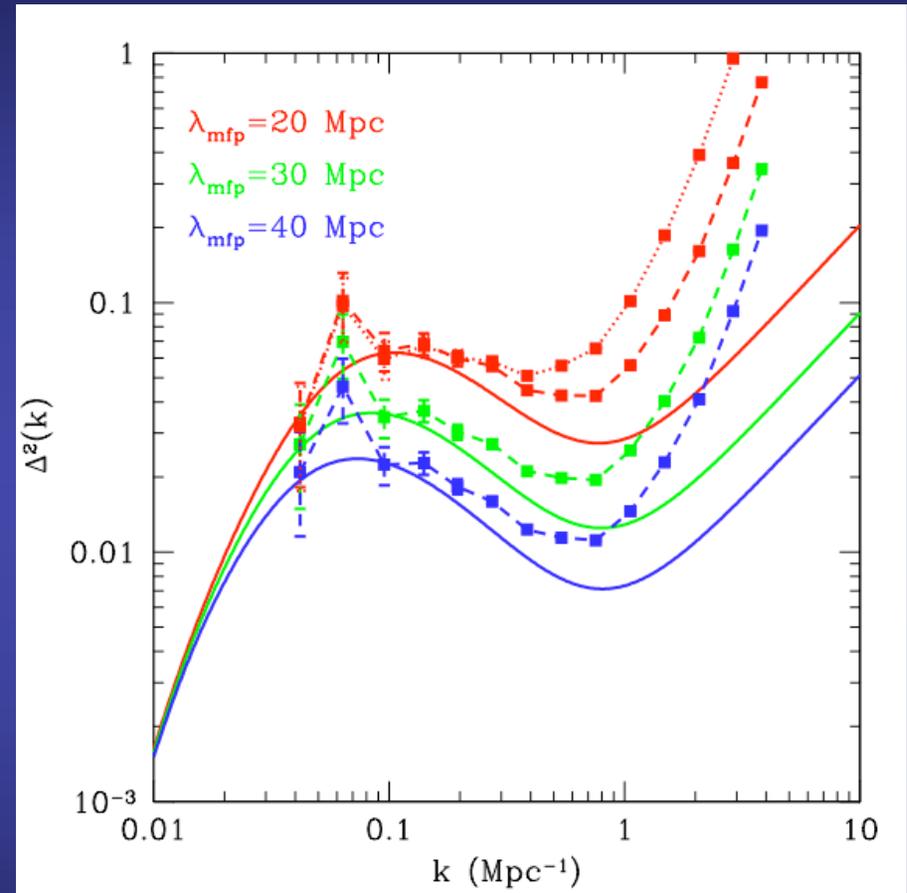
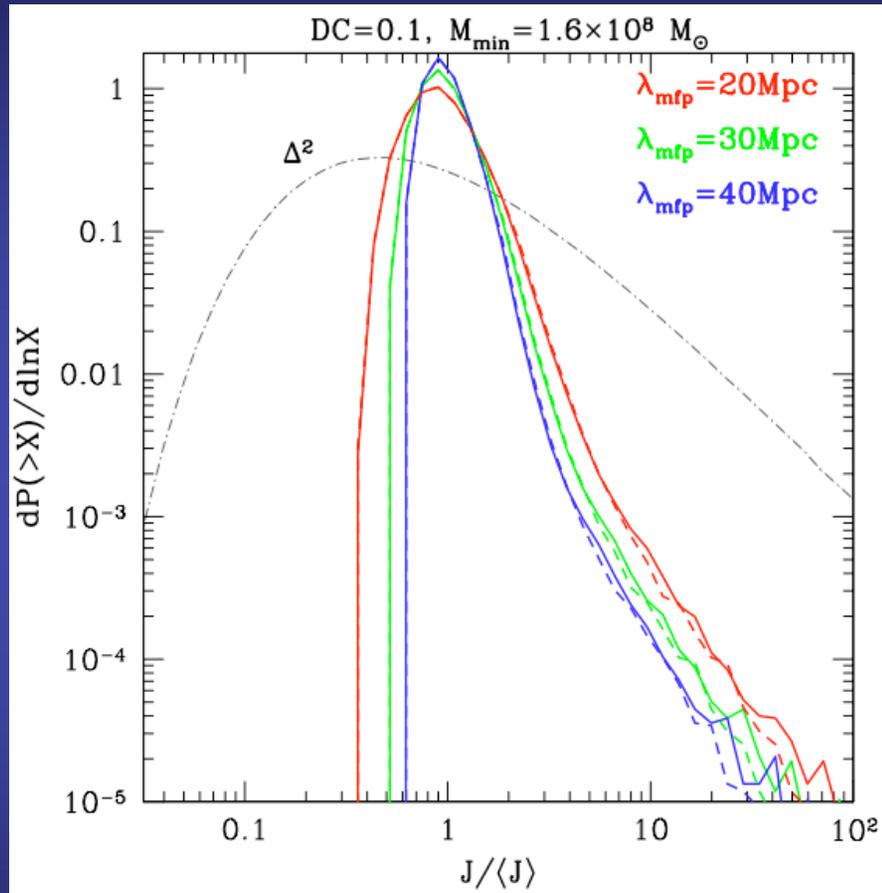
Mesinger & Furlanetto (2009)

- “modest” mfp at  $z \sim 5-6$   
+ clustering of sources  
+  $1/r^2$  flux  
= **spatial fluctuations in UVB**  
(see also sec. 5 in Bolton & Haehnelt 2007)



SFR taken from Trac+ (2008)  $z=5.71$  output

# Inhomogeneous UVB at $z \sim 5-6$

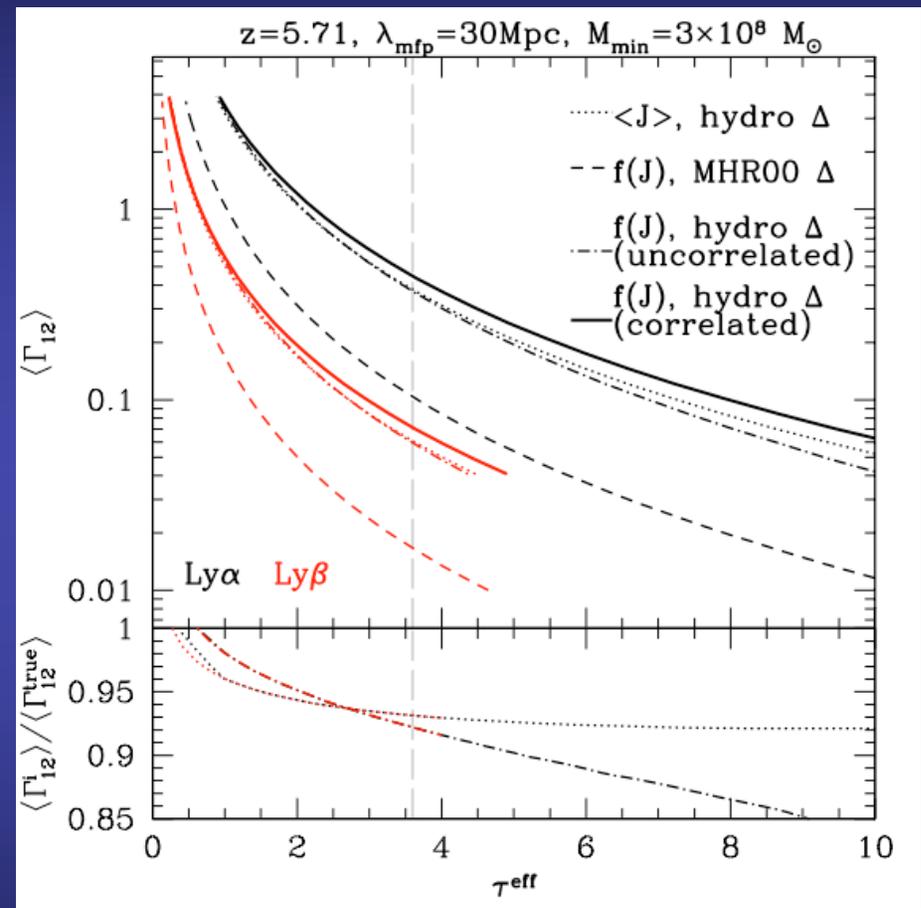
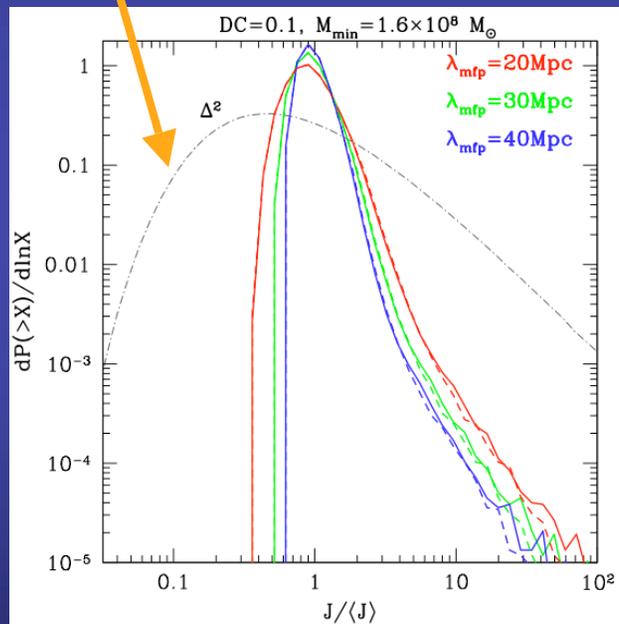


Mesinger & Furlanetto (2009)

# Not important for Ly $\alpha$ forest

*not real mock spectra, but...*

- $\Delta^2$  distribution is even wider and dominates  $\tau \sim \Delta^2/\Gamma$



SFR and density taken from Trac+ (2008)