



Credit: Gemini/AURA

# Characterizing Exoplanet Atmospheres with Gemini/GMOS: First Results

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University of Colorado, Boulder

P.I. J.-M. Désert

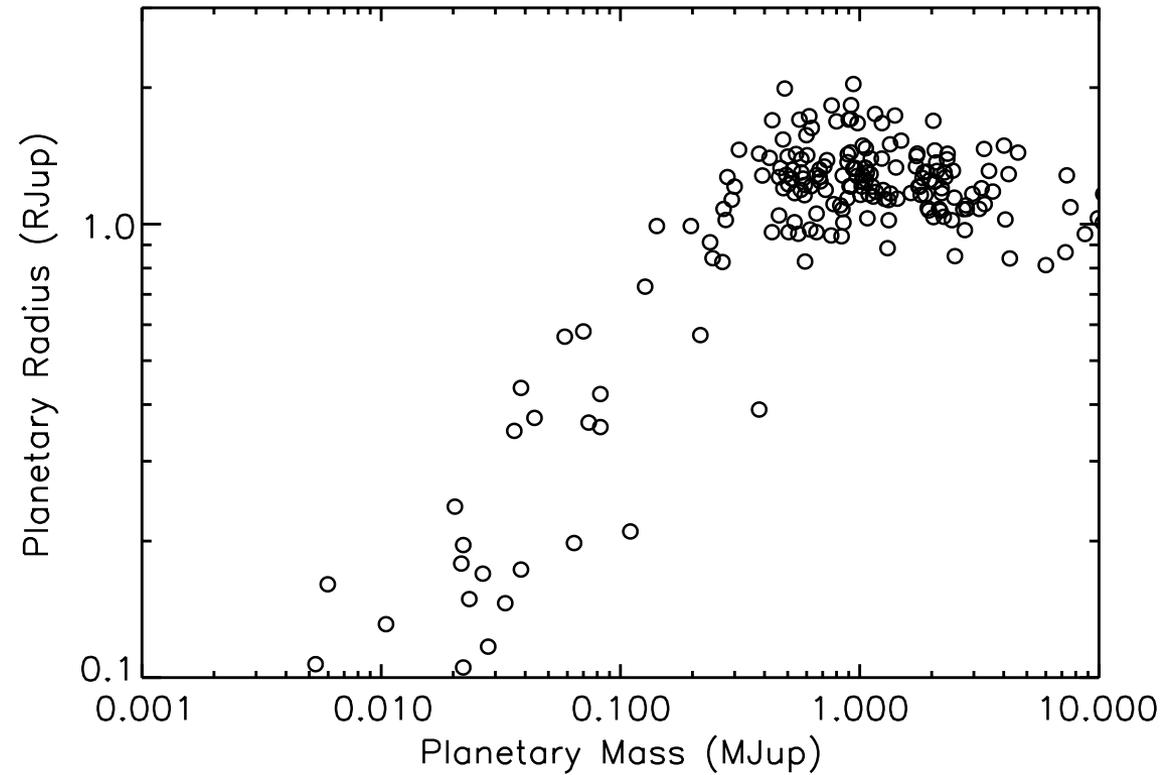
**Collaborators: J. Bean, J. Fortney, K. Stevenson, M. Bergmann, D. Deming, S. Seager**



University of Colorado

# Introduction to the GMOS Survey

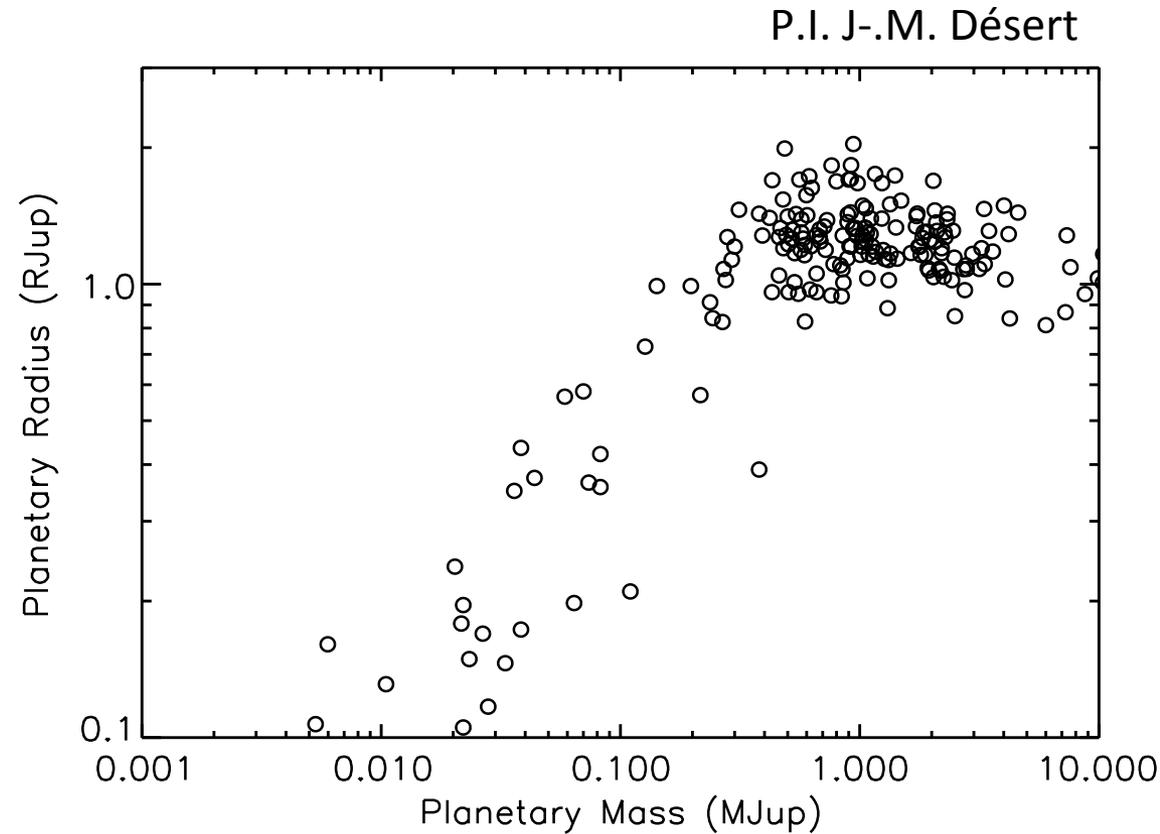
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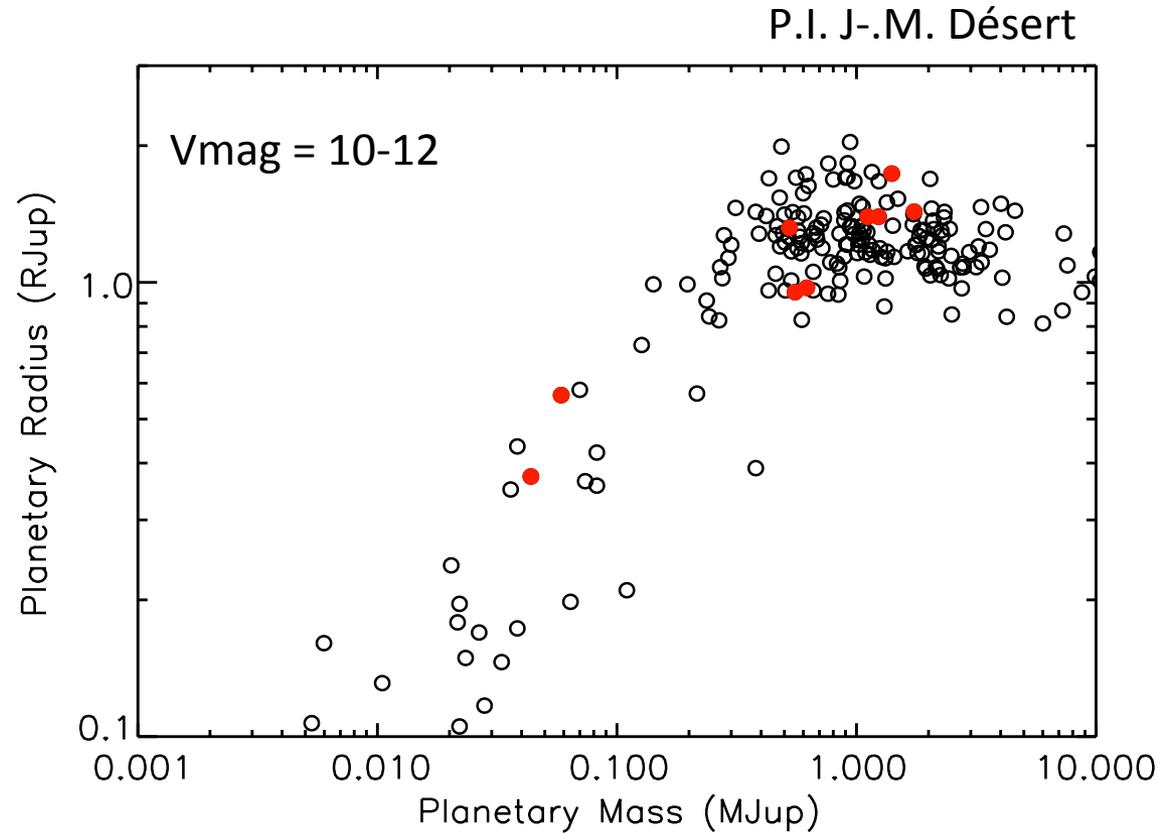
- **Aims**

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  - Comparative study
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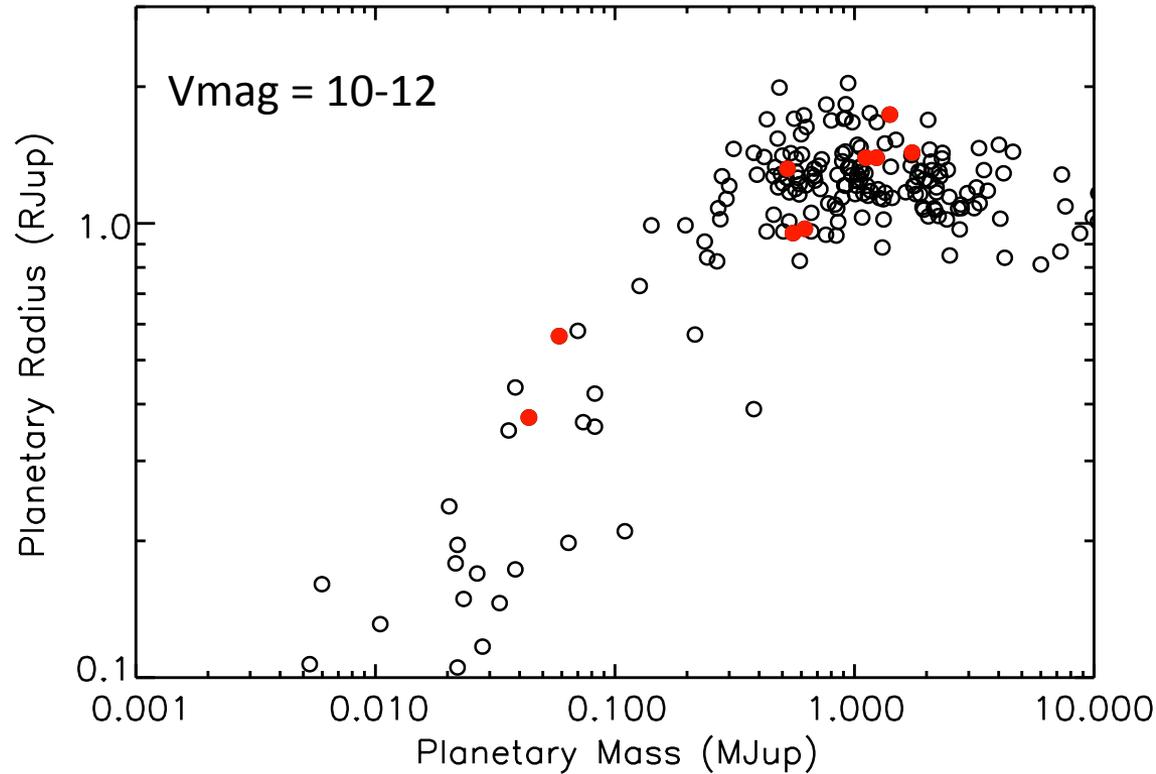
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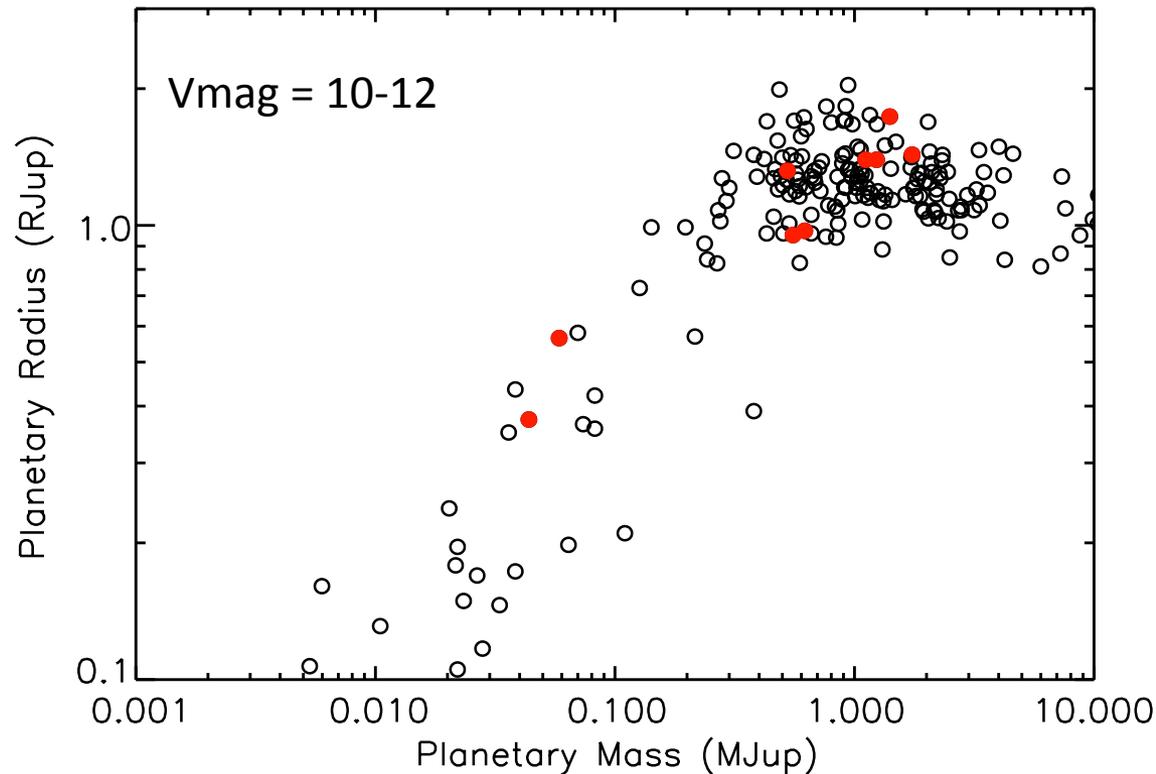
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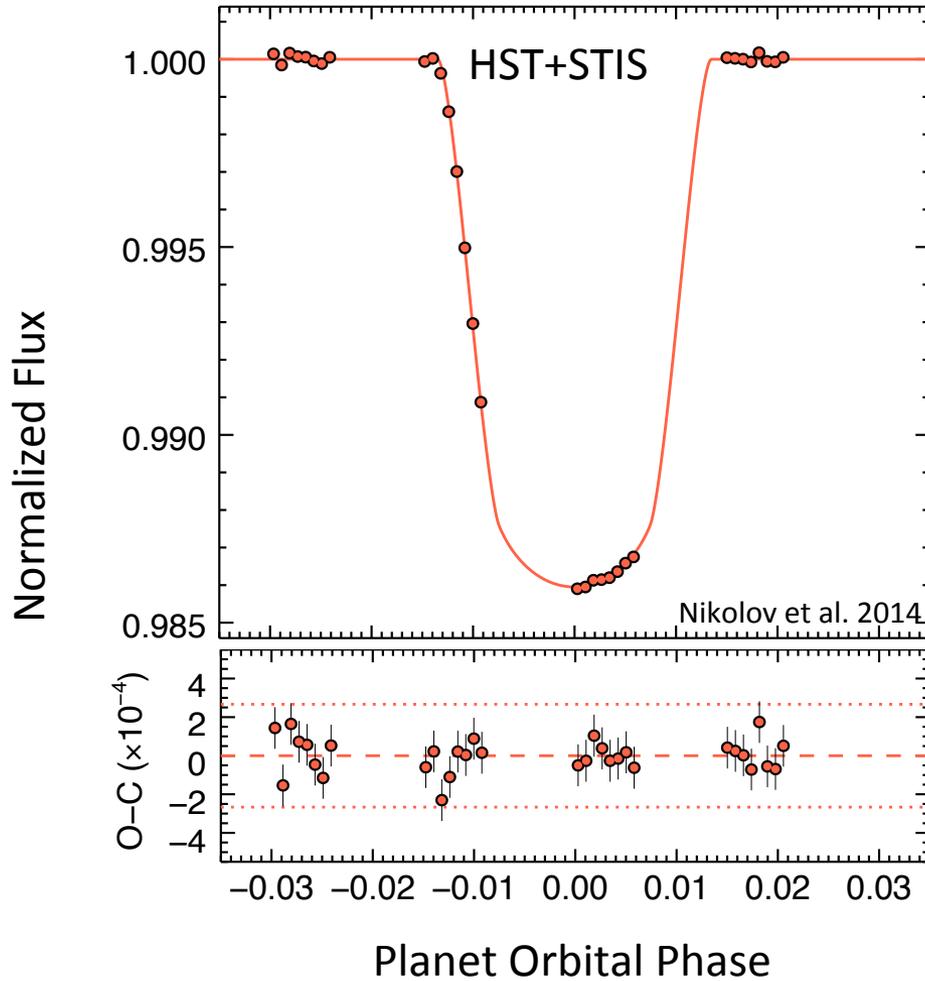
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Total = 29 transits, 193 hours

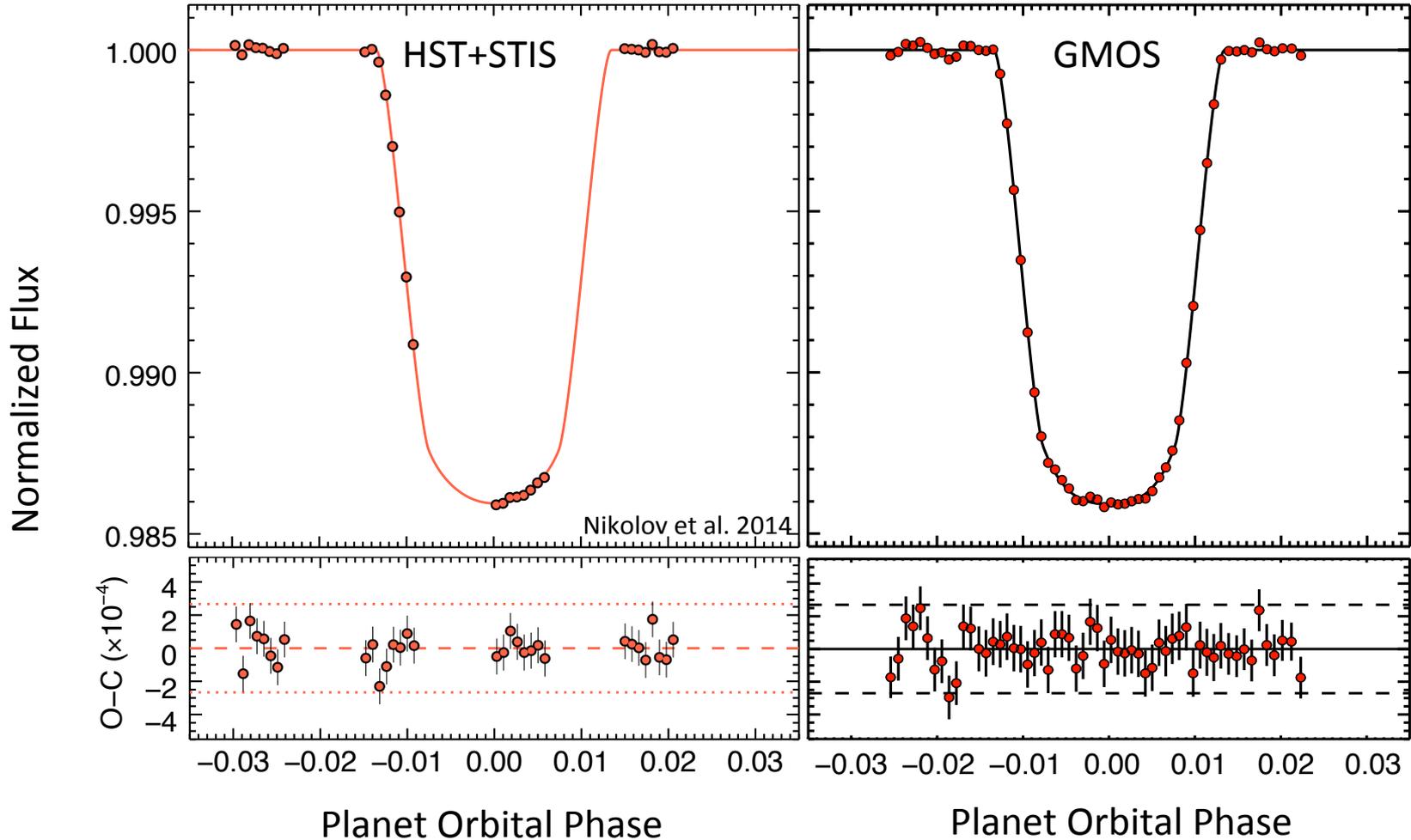
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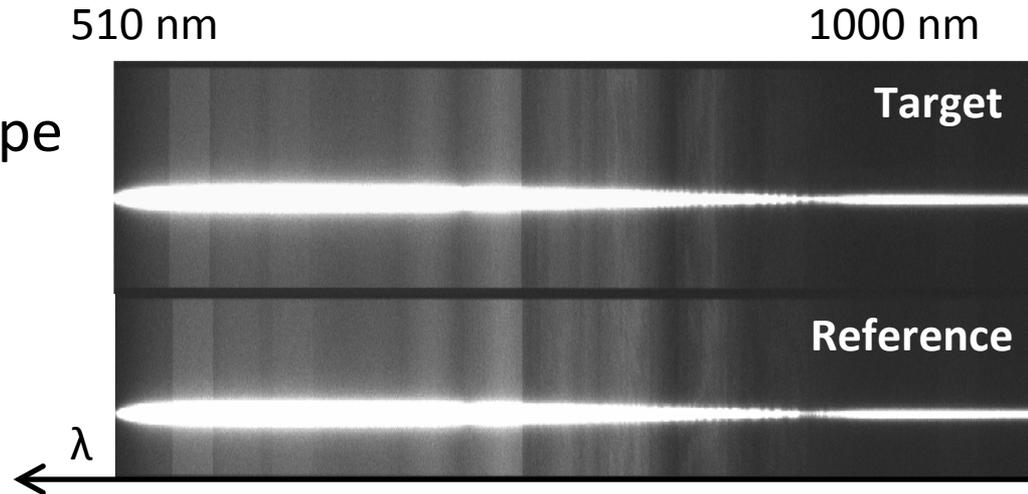


# Why Gemini/GMOS?

- Ground-based
- Multi-Object optical spectroscopy
- Fully-sky coverage with GMOS instruments
- Large 5 x 5 arcmin FOV
- Queue mode

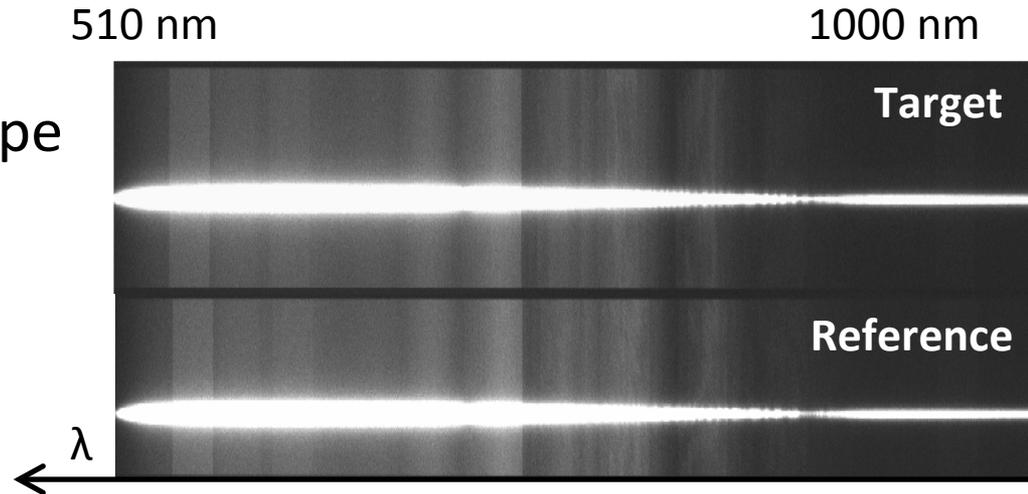
# Our Technique: MOS (GMOS)

- Target + reference stars:  
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- This technique allows us to  
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wavelength by wavelength
- Wide 10'' slit to improve  
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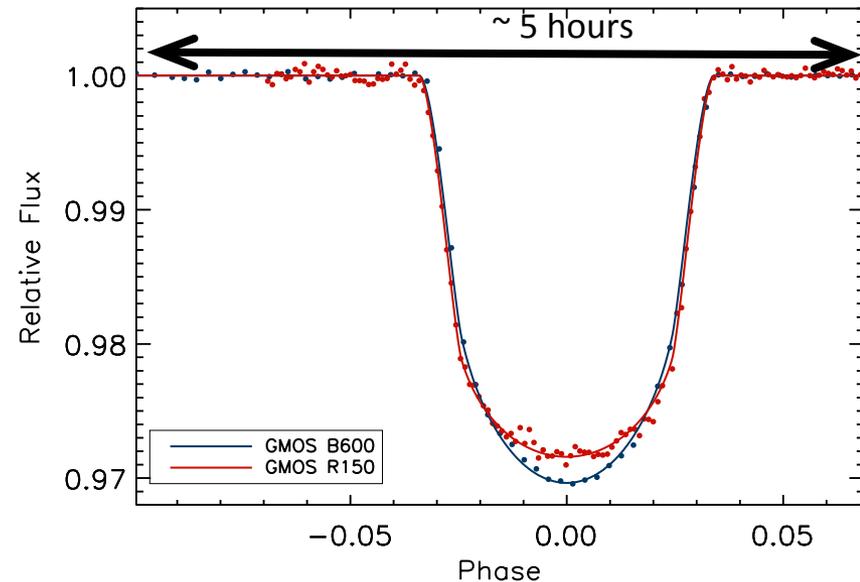
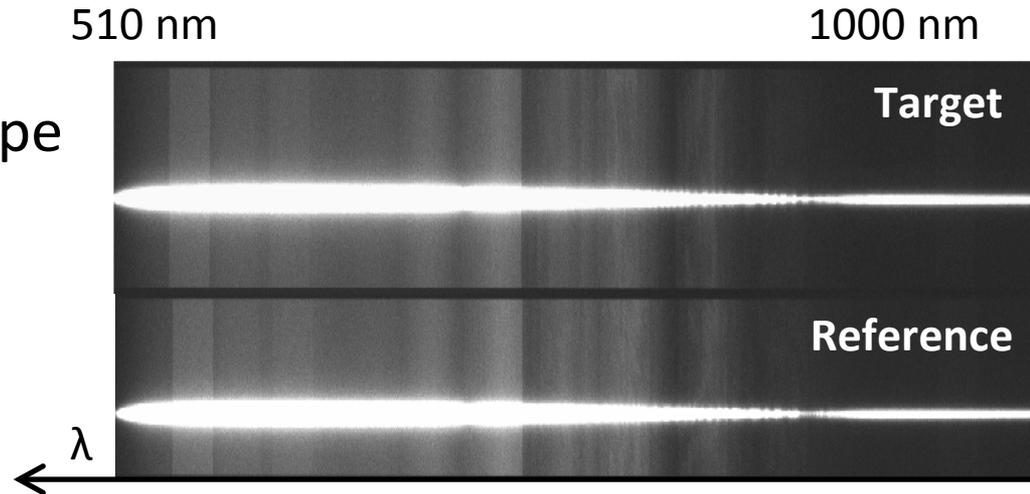
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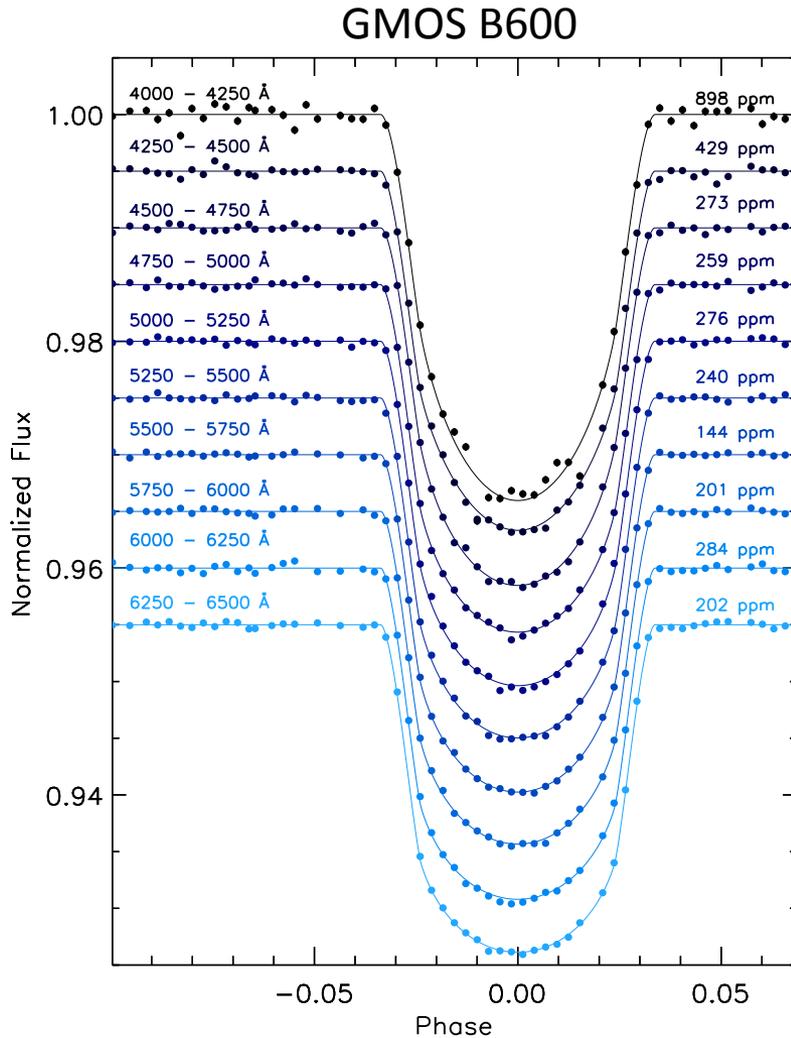


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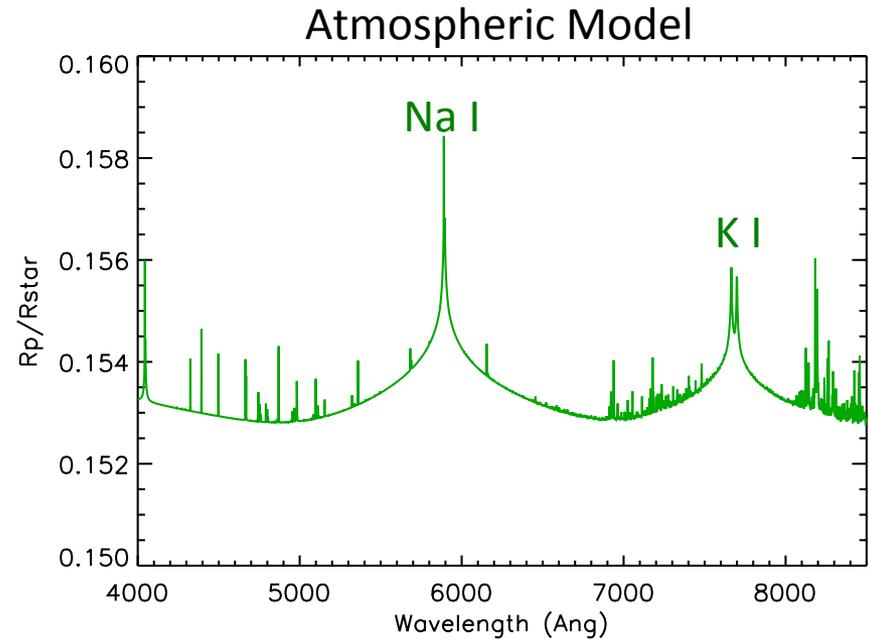
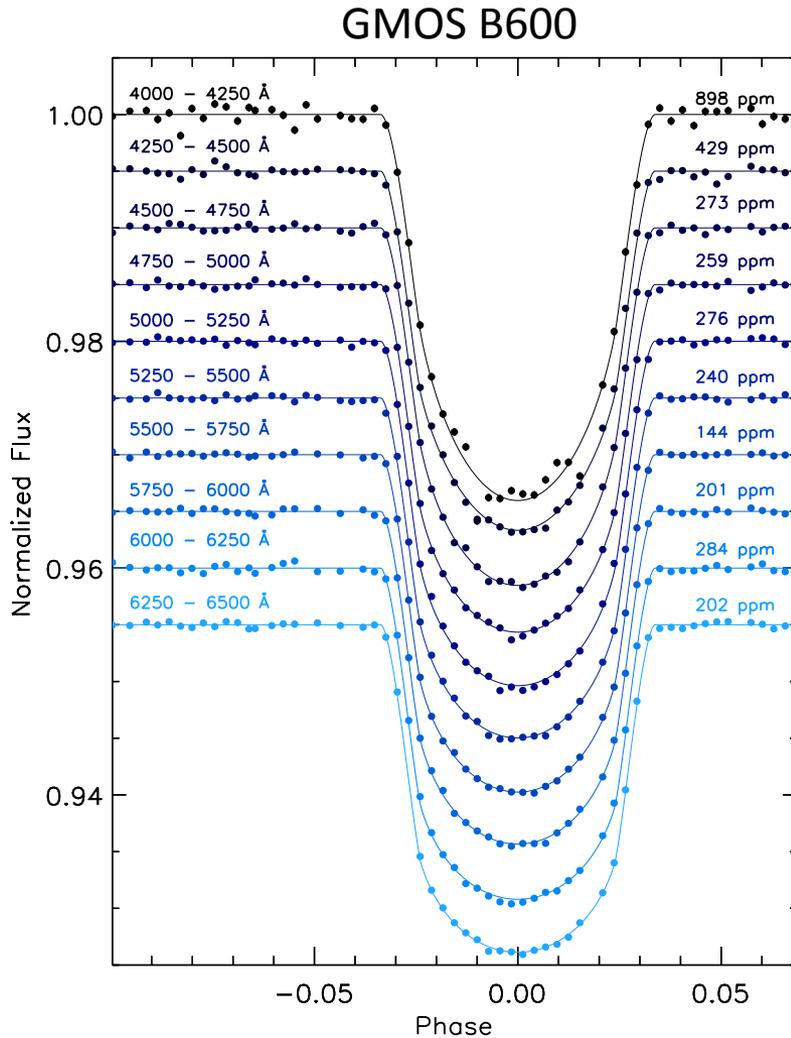
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# Transmission Spectra

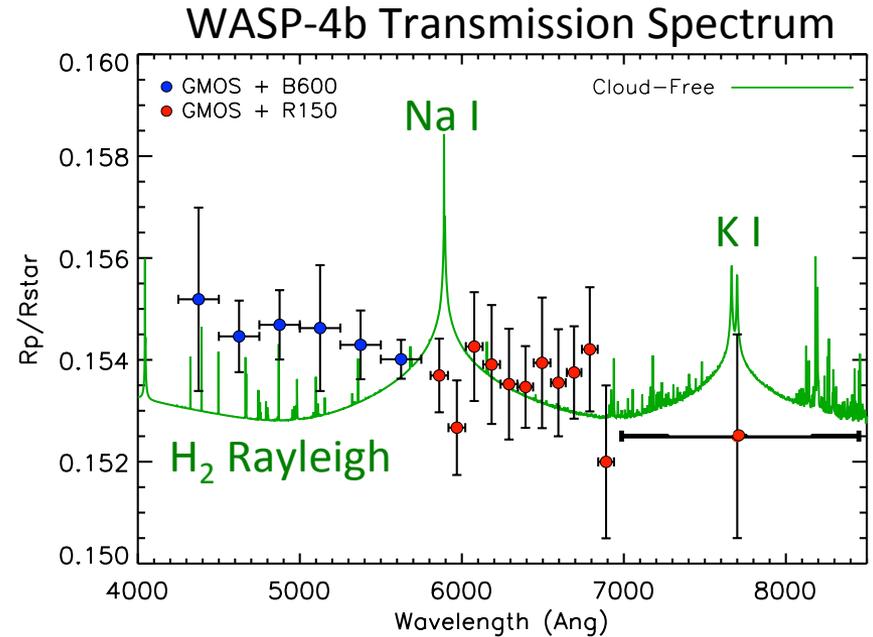
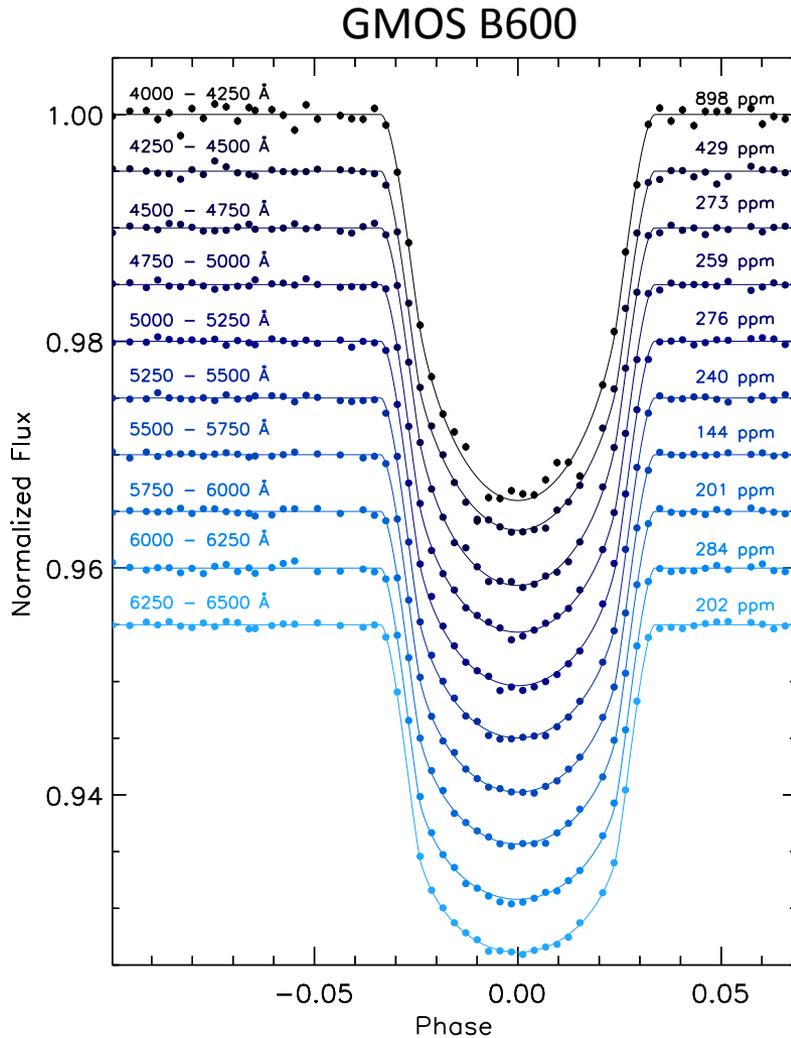


# Transmission Spectra: WASP-4b



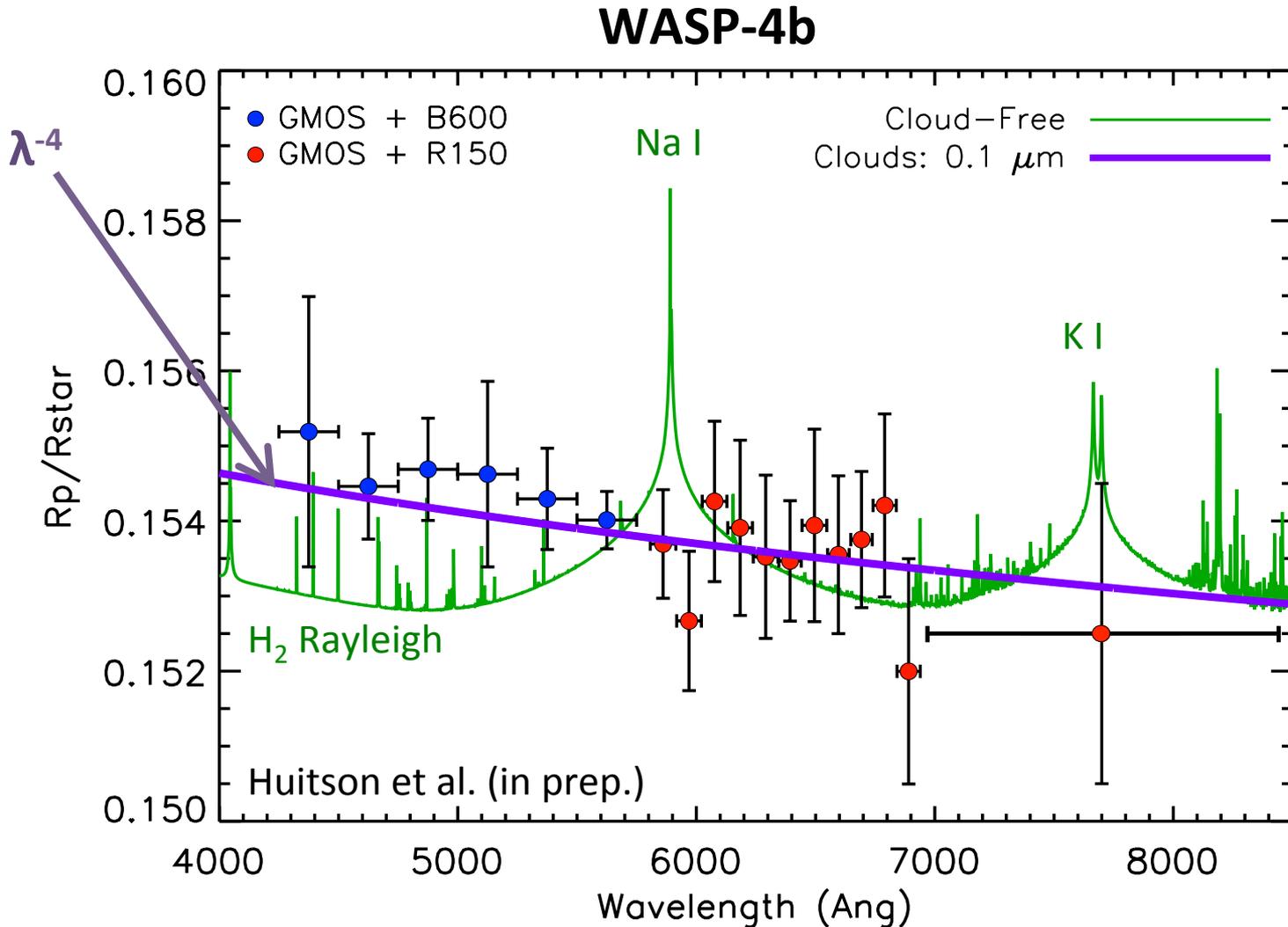
Model: Fortney et al. (2008,2010)

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# WASP-4b: Cloud-Dominated



# Challenges

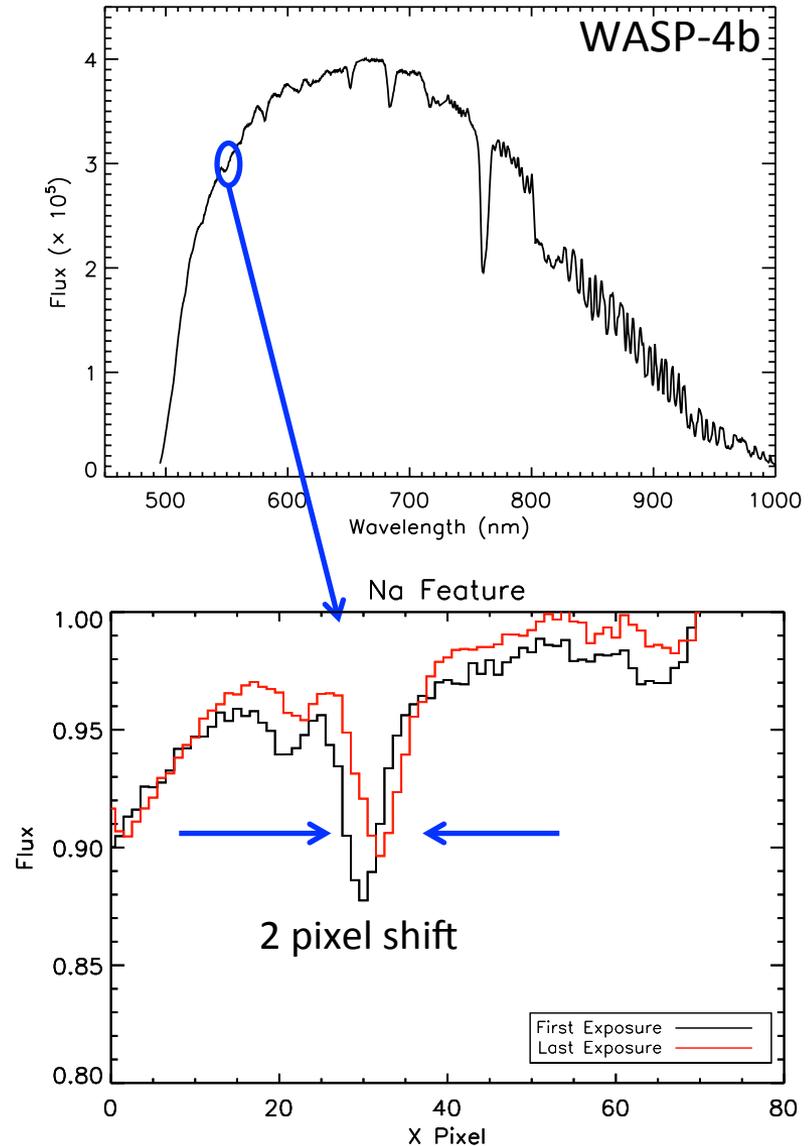
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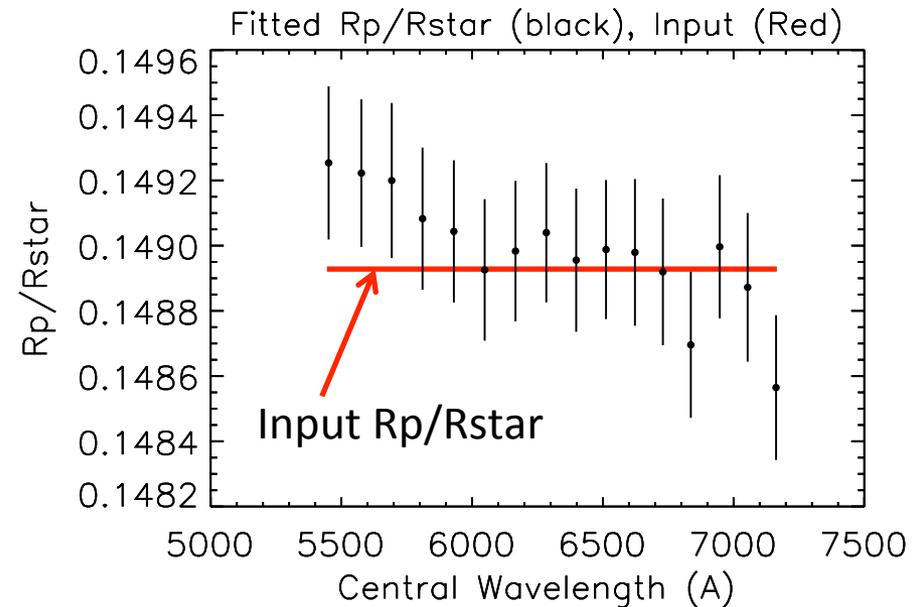
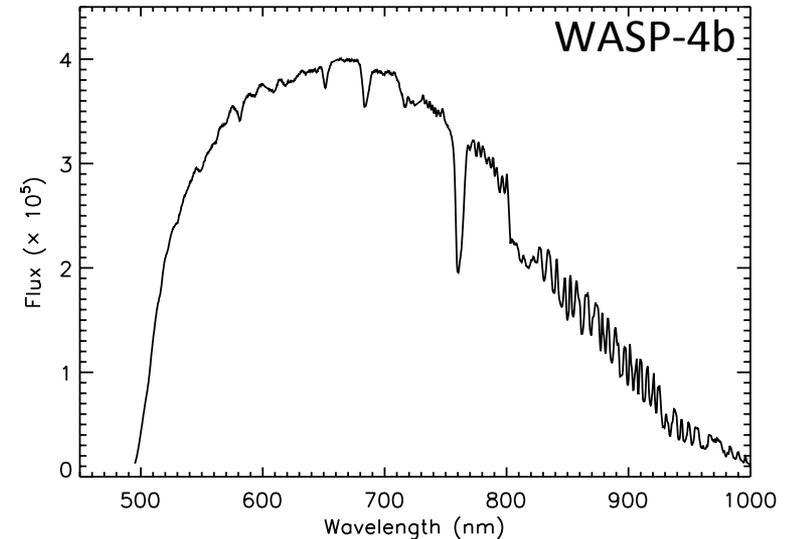
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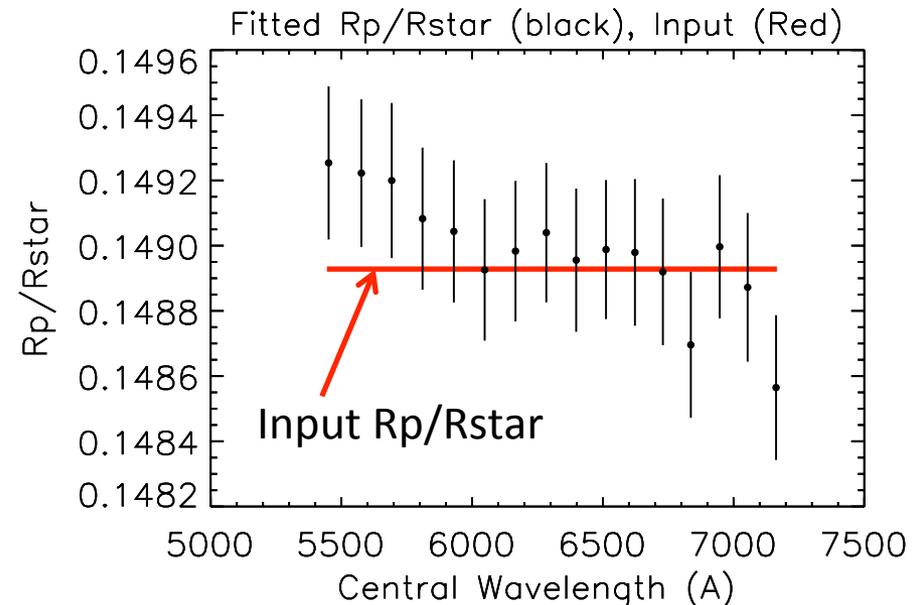
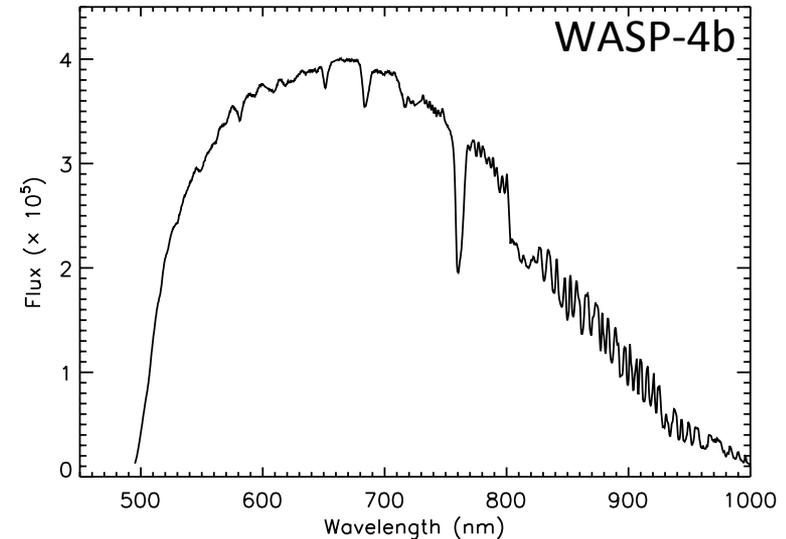
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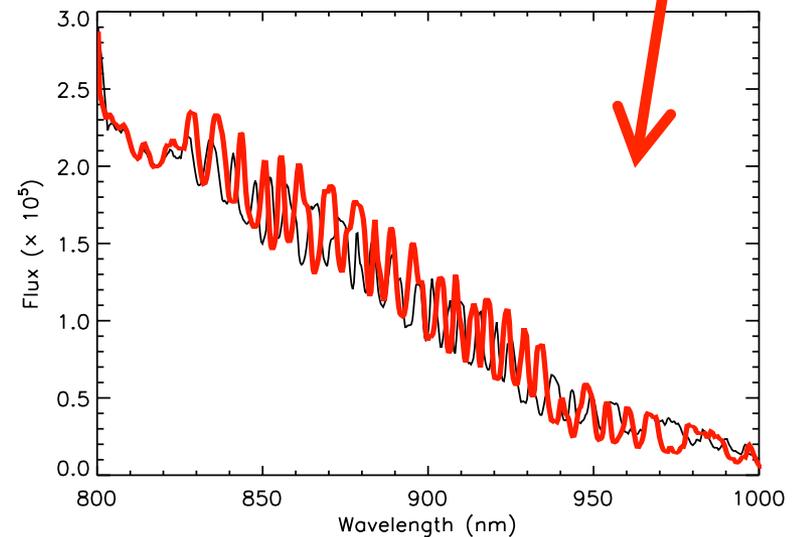
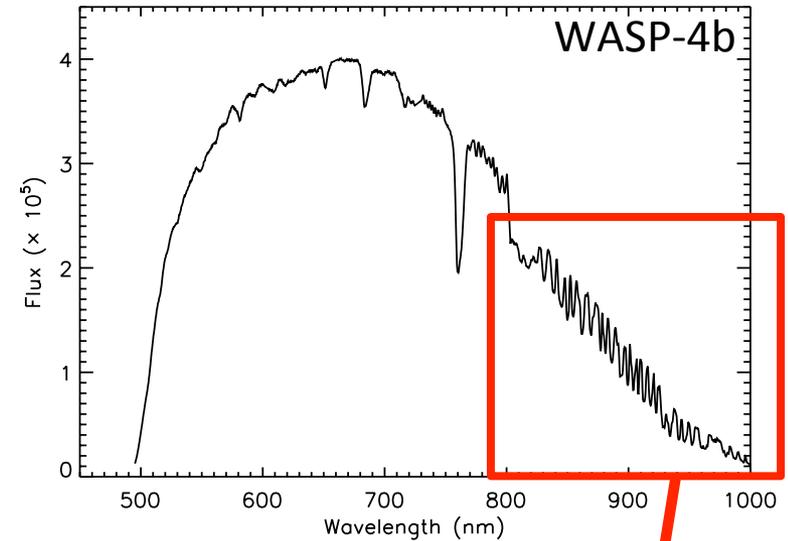
# Challenges

- Requirement:  
100 ppm/10 nm
- Problem: Wavelength solution changes in time
- Solution: procedures to measure and correct shifts in focal plane
  - Precision improved 2x



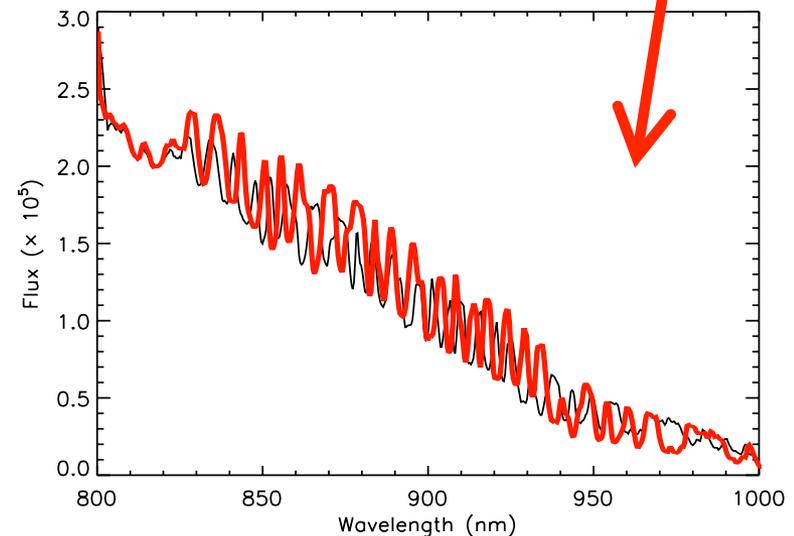
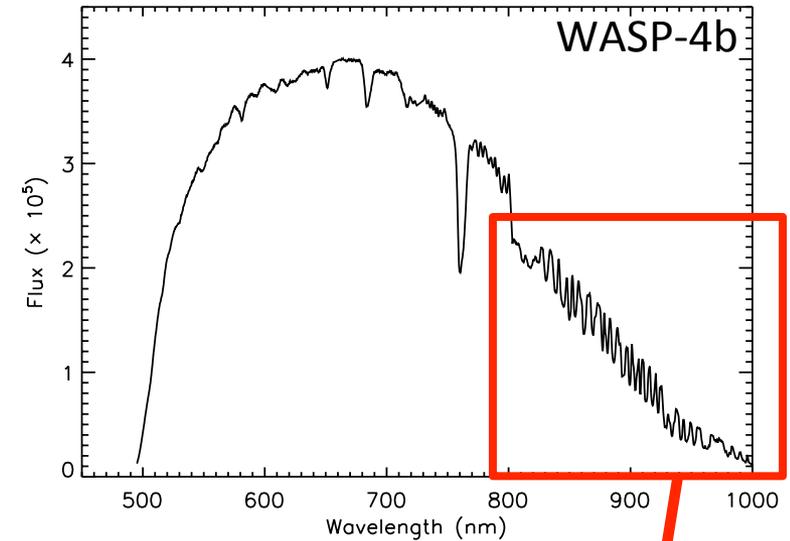
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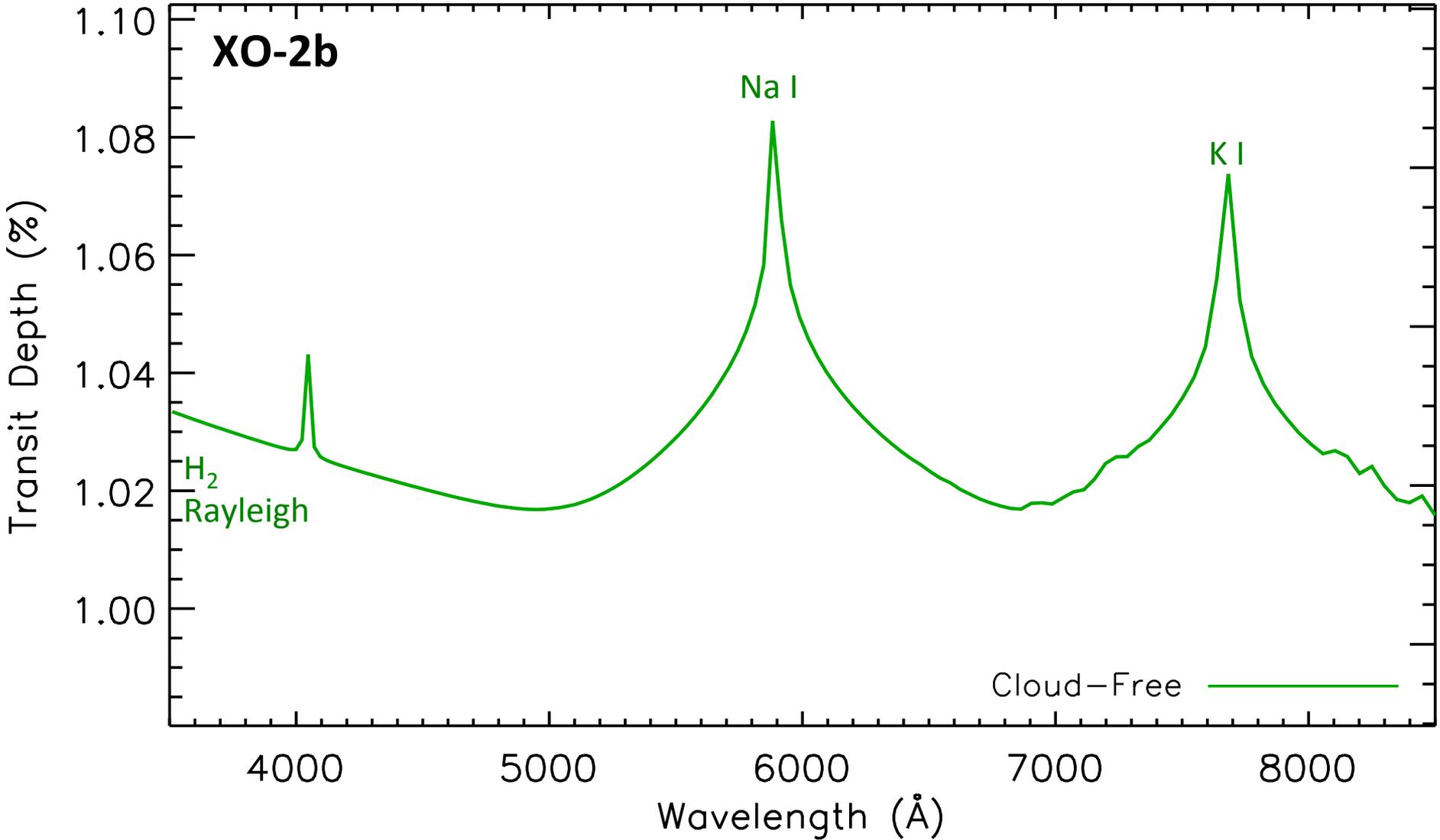


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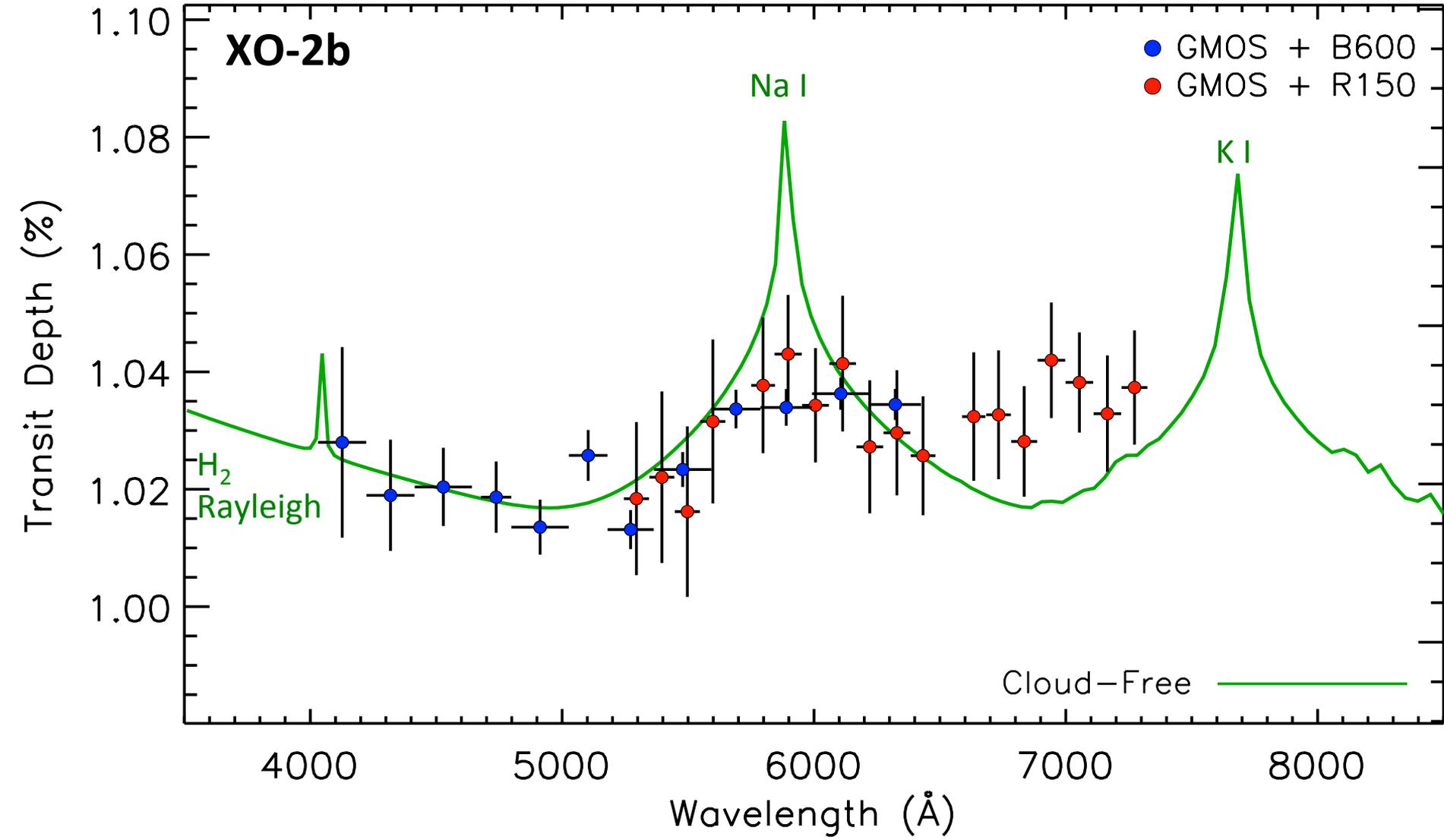
- Requirement:  
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- Problem: Fringing at 10% amplitude
- Solution: Use large photometric bin to measure continuum level



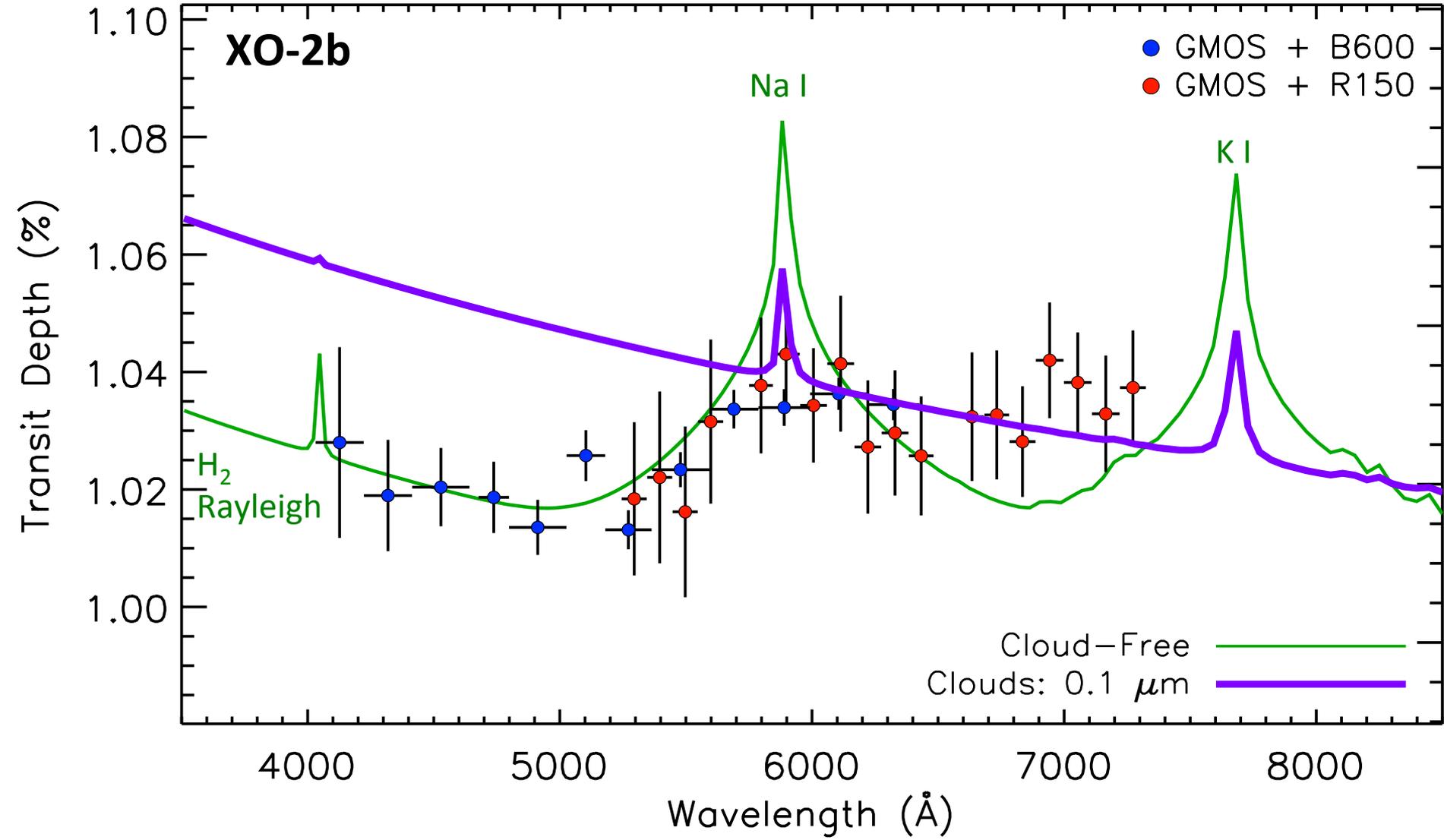
# Second Result: XO-2b



# XO-2b: Detection of Na

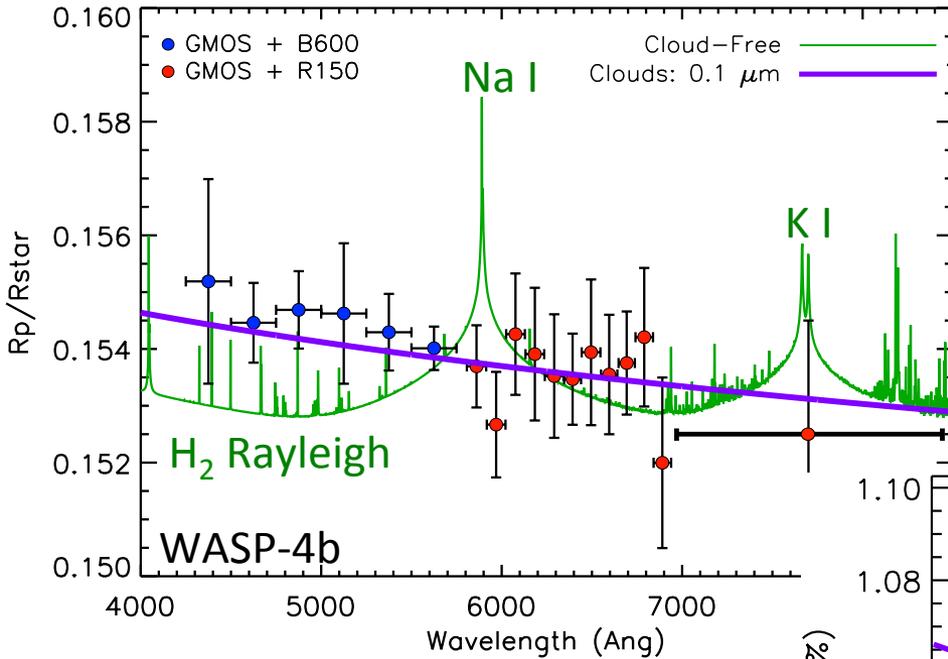


# XO-2b: Cloud-Free



# First Results from GMOS Survey

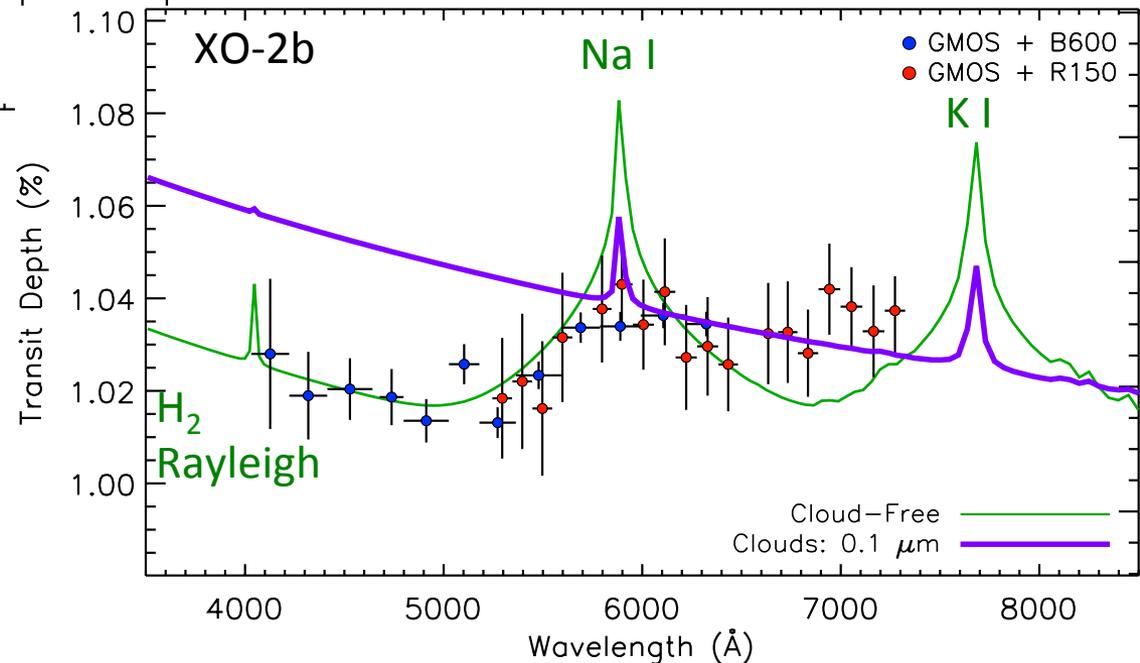
## Clouds detected



## First Important Result

Two planets similar, yet two different signatures => diversity

## No Clouds, detection of Na



# Summary

- Conducted a 9-planet survey
- Worked in-depth on 2 targets
- Preliminary results show diversity in planet atmospheres
- Pathfinder for MOS technique (e.g. Keck, VLT).
- Lessons:
  - Repeatability
  - MOS provides the precisions that we require
  - Scheduling
  - Shifting of spectra can introduce spurious signal and must be corrected
- Require dedicated instrument in order to improve on stability