



SYRTE

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CNRS - Université Pierre et Marie Curie

Systèmes de Référence Temps-Espace

# Variation des constantes Progrès des méthodes astrophysiques et en laboratoire



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# Tests de variation des constantes fondamentales utilisant les horloges atomiques



# Outline

- ✓ Sensibility of atomic transitions to varying constants.
- ✓ Principle of atomic clocks and review of some experimental methods.
- ✓ Comparisons of Rb and Cs hyperfine frequency at LNE-SYRTE. Recent results.
- ✓ Summary of clock results.
- ✓ Variation of constants with gravity.

## TRANSITIONS ATOMIQUES ET CONSTANTES FONDAMENTALES (1)

- ✓ Atomic transitions and fundamental constants

- ✓ Hyperfine transition

$$\nu_{\text{hfs}}^{(i)} \simeq R_{\infty} c \times \mathcal{A}_{\text{hfs}}^{(i)} \times g^{(i)} \left( \frac{m_e}{m_p} \right) \alpha^2 F_{\text{hfs}}^{(i)}(\alpha).$$

- ✓ Electronic transition

$$\nu_{\text{elec}}^{(i)} \simeq R_{\infty} c \times \mathcal{A}_{\text{elec}}^{(i)} \times F_{\text{elec}}^{(i)}(\alpha).$$

- ✓ See also, molecular vibrational and rotation  $\Rightarrow (m_e/m_p)^{1/2}, m_e/m_p$

- ✓ Actual measurements: ratio of frequencies

$$\begin{aligned} \frac{\nu_{\text{elec}}^{(ii)}}{\nu_{\text{elec}}^{(i)}} &\propto \frac{F_{\text{elec}}^{(ii)}(\alpha)}{F_{\text{elec}}^{(i)}(\alpha)} \\ \frac{\nu_{\text{hfs}}^{(ii)}}{\nu_{\text{elec}}^{(i)}} &\propto g^{(ii)} \frac{m_e}{m_p} \alpha^2 \frac{F_{\text{hfs}}^{(ii)}(\alpha)}{F_{\text{elec}}^{(i)}(\alpha)} \\ \frac{\nu_{\text{hfs}}^{(ii)}}{\nu_{\text{hfs}}^{(i)}} &\propto \frac{g^{(ii)} F_{\text{hfs}}^{(ii)}(\alpha)}{g^{(i)} F_{\text{hfs}}^{(i)}(\alpha)}. \end{aligned}$$

- ✓ Electronic transitions test  $\alpha$  alone (electroweak interaction)

- ✓ Hyperfine and molecular transitions bring sensitivity to the strong interaction

## TRANSITIONS ATOMIQUES ET CONSTANTES FONDAMENTALES (2)

- ✓  $m_p$  ,  $g^{(i)}$  are not fundamental parameters of the Standard Model
- ✓  $m_p$  ,  $g^{(i)}$ , can be related to fundamental parameters of the Standard Model ( $m_q/\Lambda_{\text{QCD}}$ ,  $m_s/\Lambda_{\text{QCD}}$ ,  $m_q=(m_u+m_d)/2$ )

It is often assumed that : 
$$\frac{\delta(m_s/\Lambda_{\text{QCD}})}{(m_s/\Lambda_{\text{QCD}})} = \frac{\delta(m_q/\Lambda_{\text{QCD}})}{(m_q/\Lambda_{\text{QCD}})}$$

*V. V. Flambaum et al., PR **D69**, 115006 (2004)*

- ✓ Recent accurate calculations have been done for some relevant transitions

*V. V. Flambaum and A. F. Tedesco, PR **C73**, 055501 (2006)*

- ✓ Any atomic transition (i) has a sensitivity to one particular combination of only 3 parameters ( $\alpha$ ,  $m_e/\Lambda_{\text{QCD}}$ ,  $m_q/\Lambda_{\text{QCD}}$ )

$$\delta \ln \left( \frac{\nu^{(i)}}{R_\infty c} \right) \simeq K_\alpha^{(i)} \times \frac{\delta \alpha}{\alpha} + K_q^{(i)} \times \frac{\delta(m_q/\Lambda_{\text{QCD}})}{(m_q/\Lambda_{\text{QCD}})} + K_e^{(i)} \times \frac{\delta(m_e/\Lambda_{\text{QCD}})}{(m_e/\Lambda_{\text{QCD}})}$$

# COEFFICIENT DE SENSIBILITE DE QUELQUES TRANSITIONS

	$K_\alpha$	$K_q$	$K_e$
<b>Rb hfs</b>	2.34	-0.064	1
<b>Cs hfs</b>	2.83	-0.039	1
<b>H opt</b>	0	0	0
<b>Yb<sup>+</sup> opt</b>	0.88	0	0
<b>Hg<sup>+</sup> opt</b>	-3.2	0	0
<b>Dy comb.</b>	$1.5 \cdot 10^7$	0	0

$K_\alpha, K_e$  : accuracy at the percent level or better

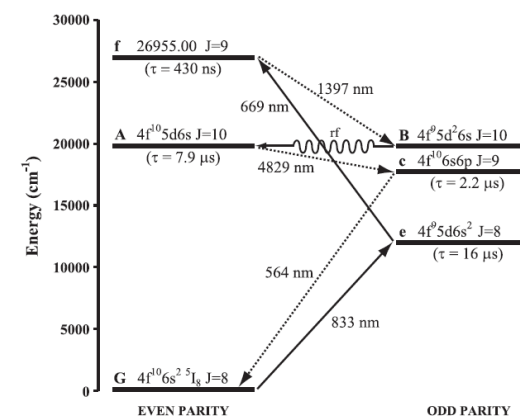
$K_q$  : accuracy ?

Atom	<sup>87</sup> Rb	<sup>133</sup> Cs
Method A	-0.074	0.127
Method B	-0.056	0.044
Method C	-0.016	0.009

PR **C73**, 055501 (2006)

Dysprosium : RF transition between 2 accidentally degenerated electronic states

*Dzuba et al., PRL* **82**, 888 (1999)



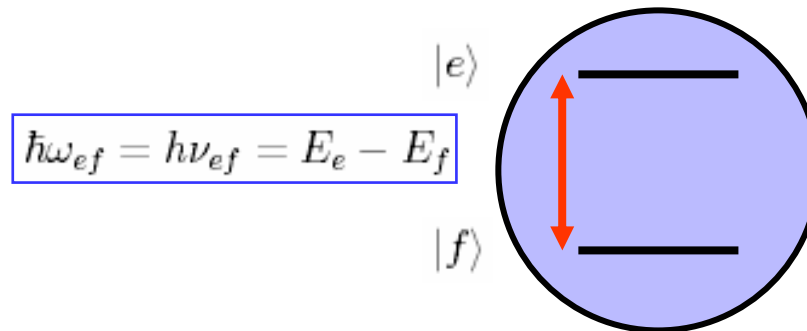
In some diatomic molecules: cancellation between hyperfine and rotational energies also leads to large (2-3 orders of magnitude enhancement)

*Flambaum, PRA* **73**, 034101 (2006)

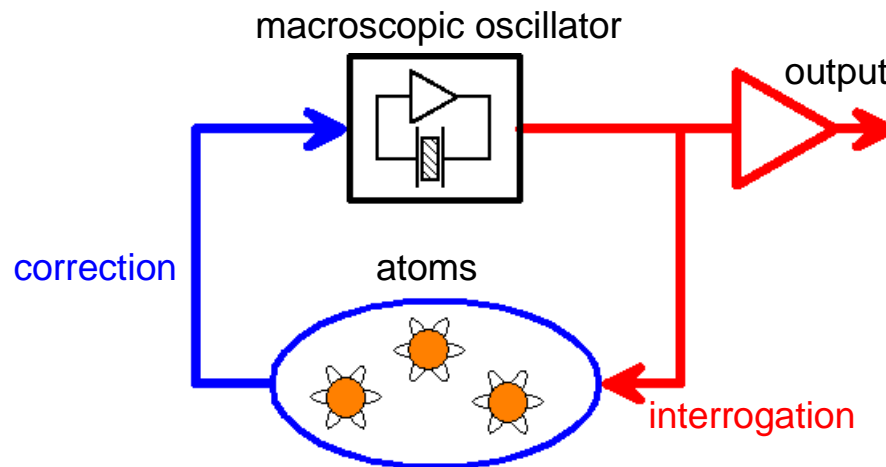
# Principle of atomic clocks (1)

Goal: deliver a signal with stable and universal frequency

Bohr frequencies of unperturbed atoms are expected to be stable and universal



Building blocks of an atomic clock



$$\omega(t) = \omega_{ef} \times (1 + \varepsilon + y(t))$$

$\varepsilon$  : fractional frequency offset

Accuracy: overall uncertainty on  $\varepsilon$

$y(t)$  : fractional frequency fluctuations

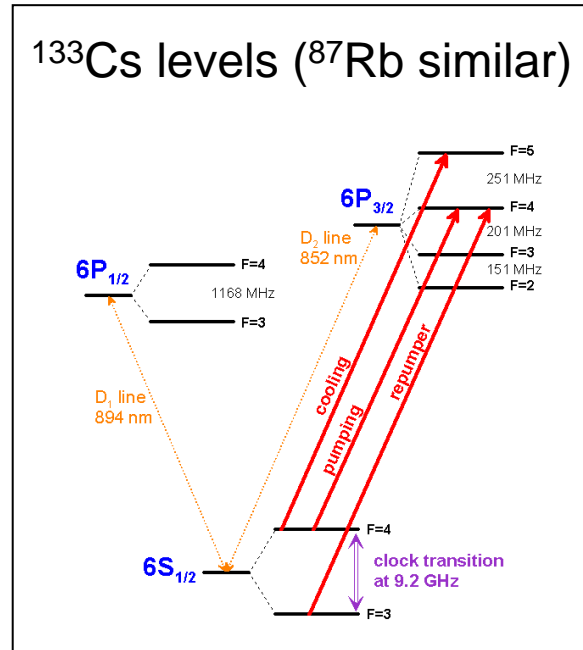
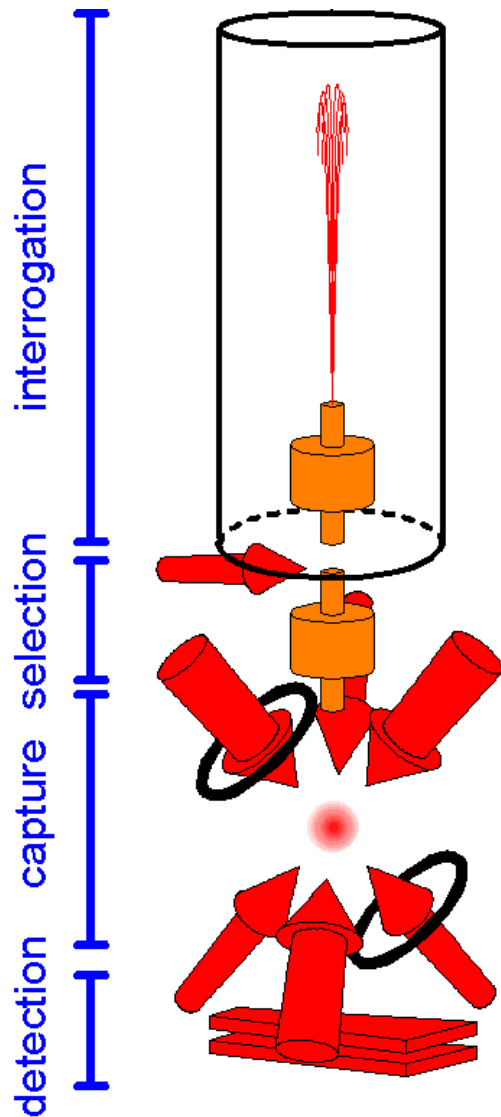
Stability: statistical properties of  $y(t)$ , characterized by the Allan variance

$$\sigma_y^2(\tau)$$

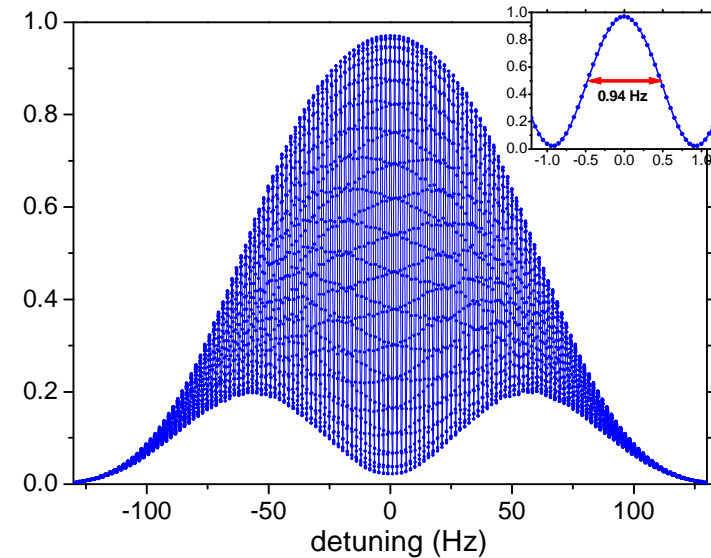
Can be done with microwave or optical frequencies, with neutral atoms, ions or molecules

=> THESE GOALS CLOSELY MATCH THE NEED OF FUNDAMENTAL TESTS

# Atomic fountain clocks



## Ramsey fringes



Atomic quality factor:

$$Q_{at} = \nu_{ef} / \Delta\nu \simeq 9.8 \times 10^9$$

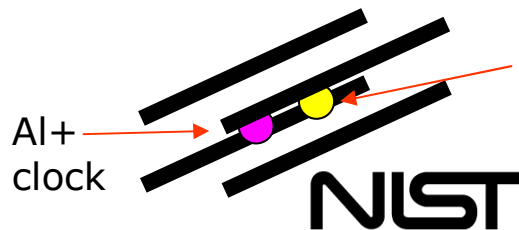
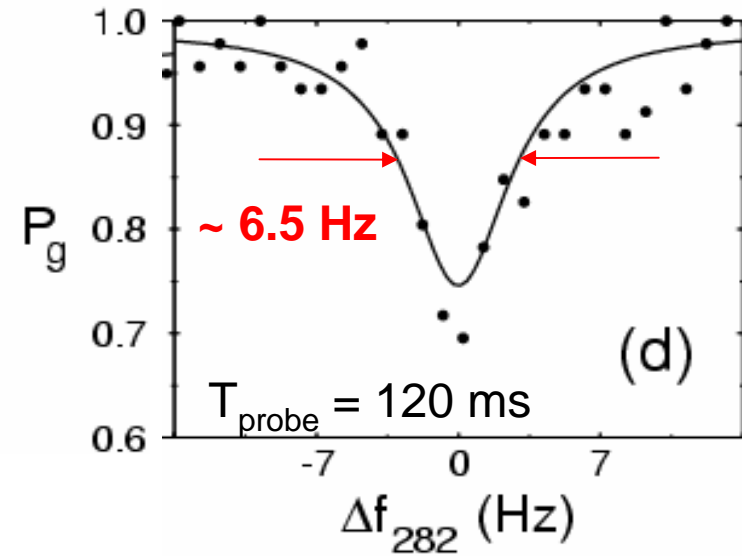
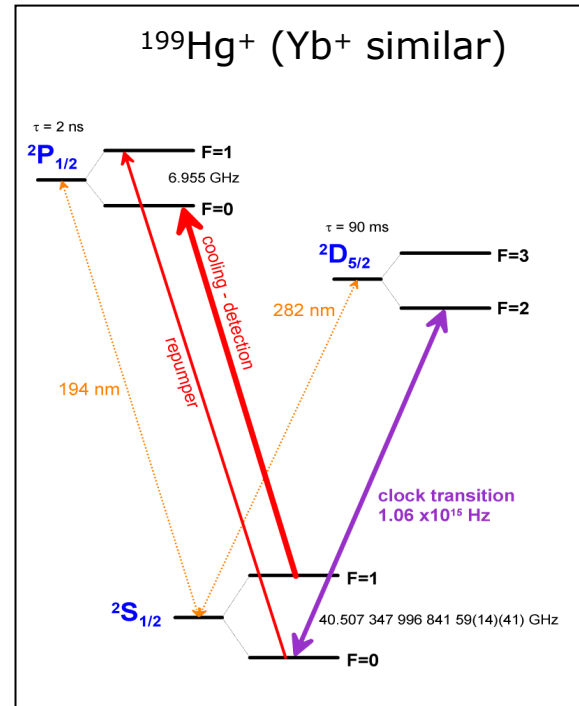
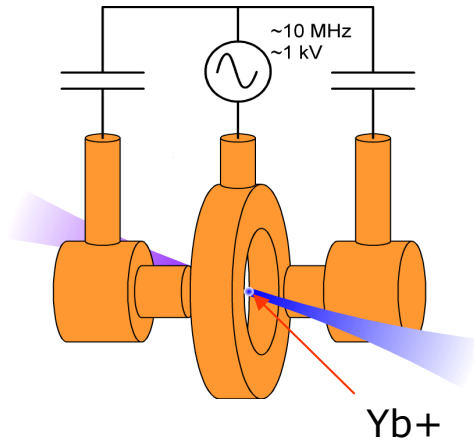
Best frequency stability:  $1.6 \times 10^{-14}$  @ 1s

Best accuracy:  $4 \times 10^{-16}$

~ 10 fountains in operation (LNE-SYRTE, PTB, NIST, USNO, ON, INRIM, NPL, USP,...) with an accuracy  $\sim 10^{-15}$  and  $< 10^{-15}$  for a few of them.



# Trapped ion clocks



Be+  
cooling  
detection

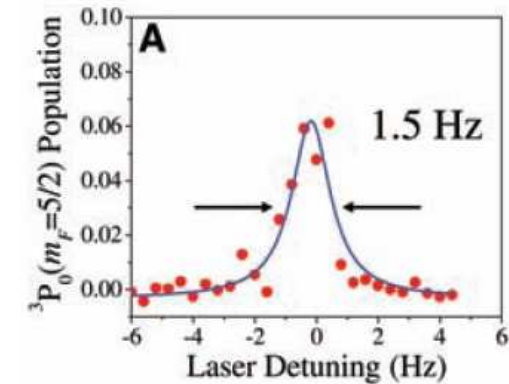
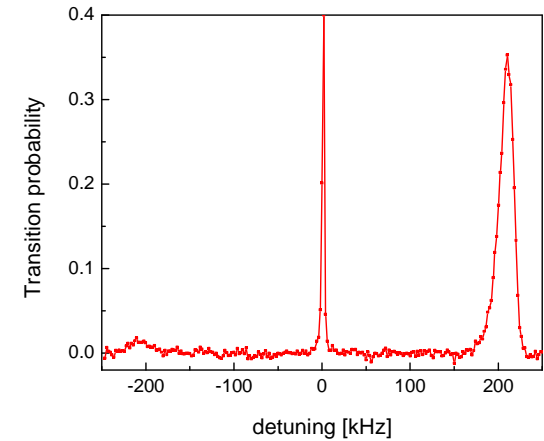
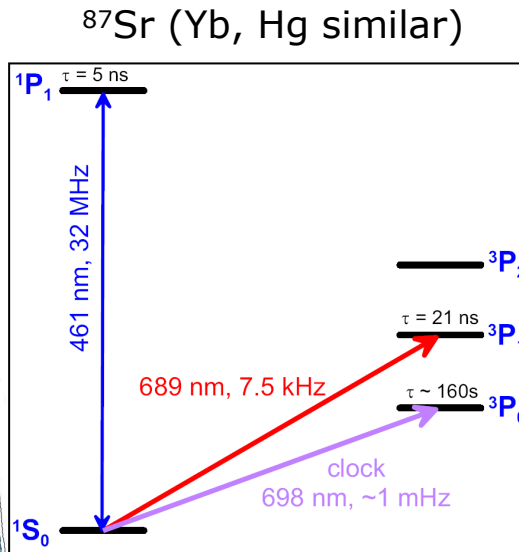
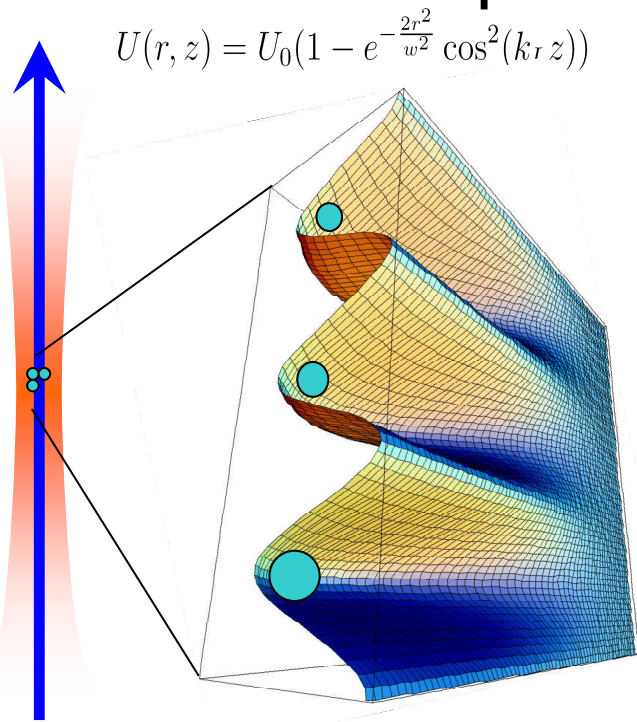
Atomic quality factor:  $\sim 1.5 \times 10^{14}$

Best frequency stability:  $\sim 2 \times 10^{-15}$  @ 1s

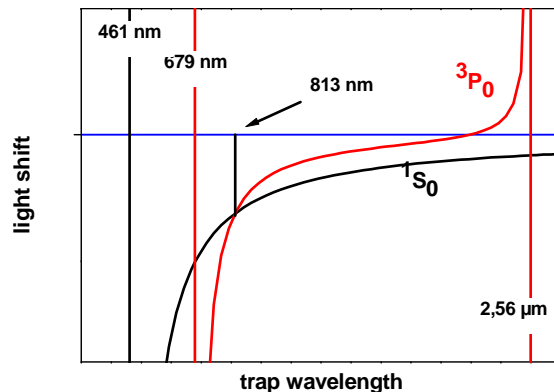
Best accuracy :  $\sim 2 \times 10^{-17}$

*Work on trapped ion clock at NIST, PTB, NPL, Innsbruck,...*

# Optical lattice clocks



“non-perturbing” dipole lattice trap



Atomic quality factor:  $\sim 2.8 \times 10^{14}$

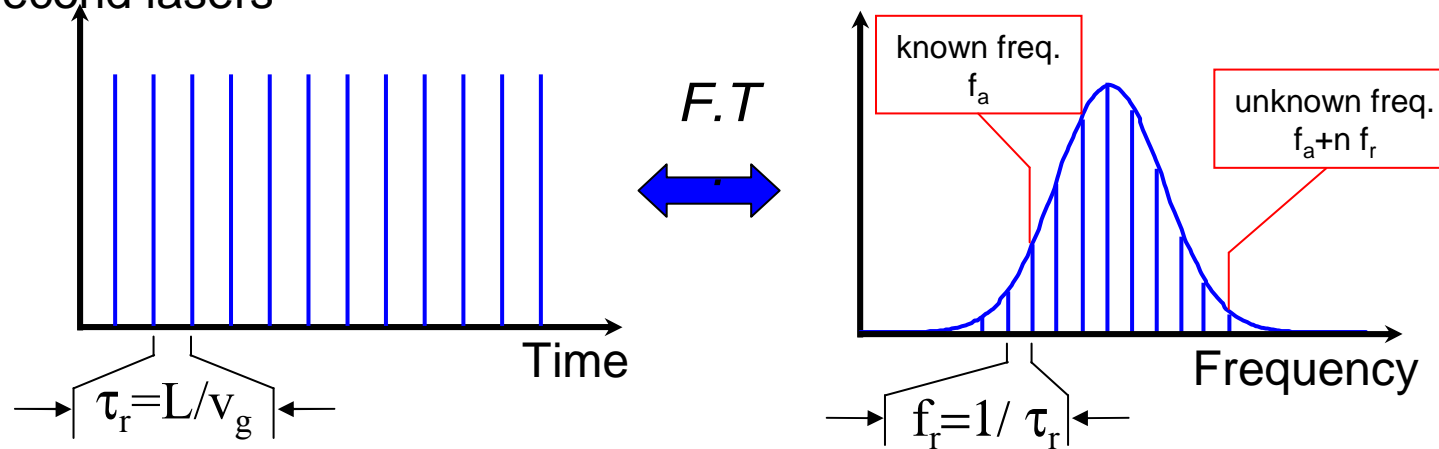
Best frequency stability:  $\sim 2 \times 10^{-15}$  @ 1s

Best accuracy :  $\sim 1 \times 10^{-16}$

Work on optical lattice clocks at Tokyo university, LNE-SYRTE, JILA, NIST, INRIM,...

# Comparison methods

Between several regions of electromagnetic spectrum (MHz to  $10^{15}$  Hz): ultra low noise RF and microwave synthesizers and optical frequency combs generated by femtosecond lasers



Between remote clocks:

Satellite systems : GPS phase, TWSTFT :  $\sim 10^{-15}$  @1d, PHARAO/ACES when available

Telecom fibers :

Dissemination of RF frequency reference : few  $10^{-15}$  @1s on <100 km scale

Dissemination of optical frequency reference : few  $10^{-15}$  @1s on <100 km scale

Dissemination of optical frequency reference on continental scale under study

# LNE-SYRTE ATOMIC CLOCK ENSEMBLE

Systèmes de Référence Temps-Espace



H-maser

H,  $\mu\text{W}$



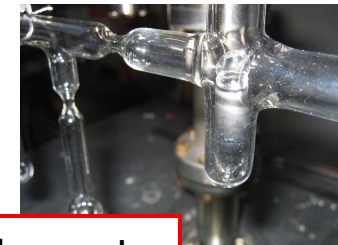
Cryogenic sapphire Osc.

Macroscopic oscillator

Phaselock loop  
 $\tau \sim 1000 \text{ s}$

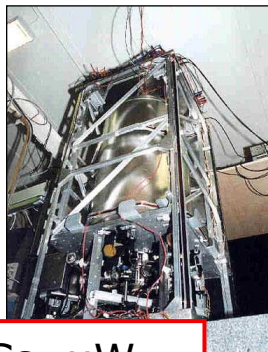


Optical lattice clock



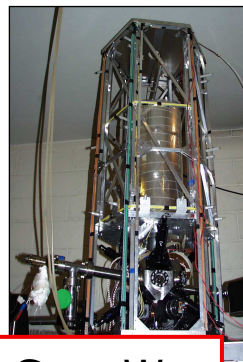
Hg, opt

FO1 fountain



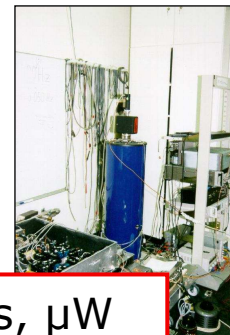
Cs,  $\mu\text{W}$

FO2 fountain



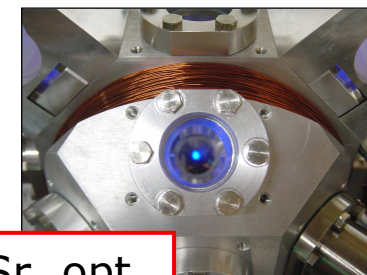
Rb, Cs,  $\mu\text{W}$

FOM transportable fountain



Cs,  $\mu\text{W}$

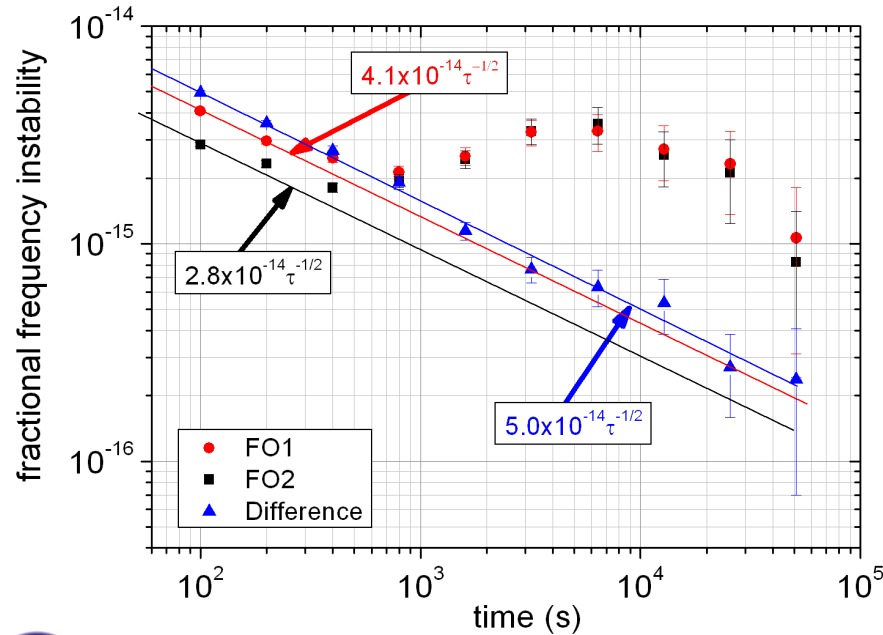
Optical lattice clock



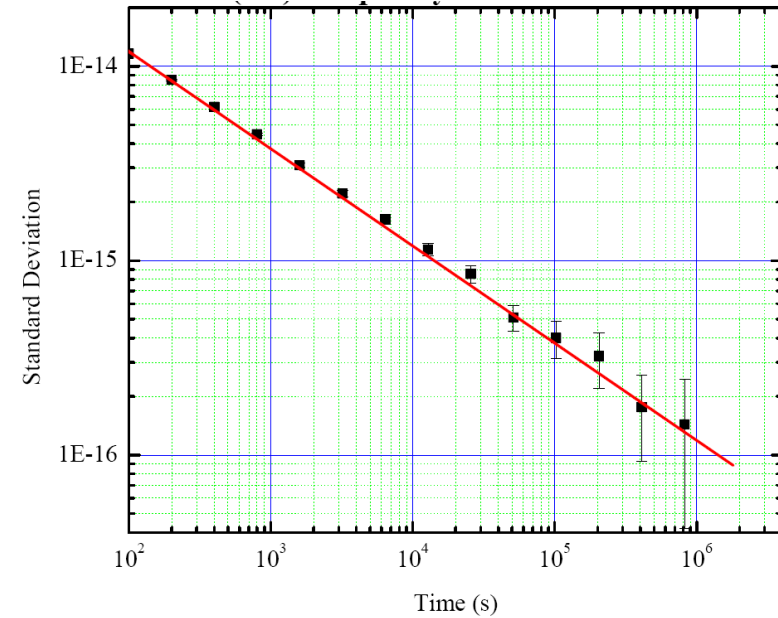
Sr, opt

# LNE-SYRTE FOUNTAINS: FREQUENCY STABILITY AND ACCURACY

FO1 vs FO2-Cs (2004)



FOM vs FO2-Rb (Nov. 2007)

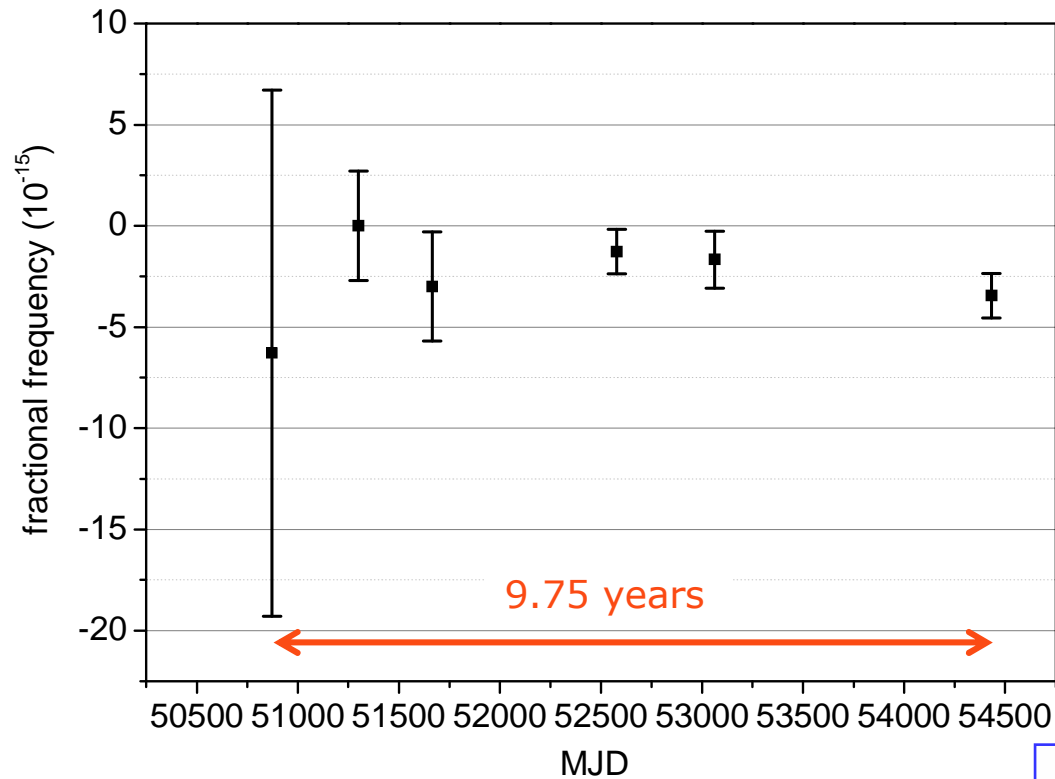


Systèmes de Référence Temps-Espace



	FO1	FO2	FOM
Quadratic Zeeman	1927.3 +/- 0.3	1927.3 +/- 0.3	210.2 +/- 1.1
Blackbody radiation	-165.0 +/- 1.0	-168.2 +/- 0.6	-160.45 +/- 0.6
Collision and cavity pulling	-201.4 +/- 2.4	-357.5 +/- 1.0	-39.5 +/- 6.7
Microwave spectral purity & leakage	0 +/- 0.6	0 +/- 0.5	0 +/- 10.0
First order Doppler	< 3.2	< 3.0	< 3.2
Ramsey & Rabi pulling	< 0.1	< 1	< 0.1
Quantized motion ("microwave recoil")	< 1.4	< 1.4	< 1.4
Background collisions	< 0.3	< 1	< 1
<b>TOTAL UNCERTAINTY</b>	<b>4</b>	<b>4</b>	<b>12</b>

## COMPARISON OF Rb and Cs HFS at LNE-SYRTE



one data point  $\Leftrightarrow$  ~1 to 2 months of measurements, with many checks of systematic shifts

Weighted least squares fit gives:

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Rb}}}{\nu_{\text{Cs}}} \right) = (-3.2 \pm 2.3) \times 10^{-16} \text{ yr}^{-1}$$

(improvement by ~2.3)



$$\frac{d}{dt} \ln \left( \frac{g_{\text{Rb}}}{g_{\text{Cs}}} \alpha^{-0.49} \right) = (-3.2 \pm 2.3) \times 10^{-16} \text{ yr}^{-1}$$

With further theory, nuclear g-factors can be related to  $m_q/\Lambda_{\text{QCD}}$ :



$$\frac{d}{dt} \ln \left( \alpha^{-0.49} [m_q/\Lambda_{\text{QCD}}]^{-0.025} \right) = (-3.2 \pm 2.3) \times 10^{-16} \text{ yr}^{-1}$$

## OVERVIEW OF MEASUREMENTS

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Rb}}}{\nu_{\text{Cs}}} \right) = (-3.2 \pm 2.3) \times 10^{-16} \text{ yr}^{-1} = -0.49 \frac{d}{dt} \ln(\alpha) - 0.025 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) \quad \text{LNE-SYRTE, JPB (2004)}$$

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Hg}^+}}{\nu_{\text{Cs}}} \right) = (3.7 \pm 3.9) \times 10^{-16} \text{ yr}^{-1} = -6.03 \frac{d}{dt} \ln(\alpha) + 0.039 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) - \frac{d}{dt} \ln(m_e / \Lambda_{\text{QCD}}) \quad \text{NIST, PRL (2007)}$$

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Yb}^+}}{\nu_{\text{Cs}}} \right) = (-7.8 \pm 14) \times 10^{-16} \text{ yr}^{-1} = -1.95 \frac{d}{dt} \ln(\alpha) + 0.039 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) - \frac{d}{dt} \ln(m_e / \Lambda_{\text{QCD}}) \quad \text{PTB, arXiv (2006)}$$

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{H}}}{\nu_{\text{Cs}}} \right) = (-32 \pm 63) \times 10^{-16} \text{ yr}^{-1} = -2.83 \frac{d}{dt} \ln(\alpha) + 0.039 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) - \frac{d}{dt} \ln(m_e / \Lambda_{\text{QCD}}) \quad \begin{array}{l} \text{MPQ} \\ + \text{LNE-SYRTE} \\ \text{PRL (2004)} \end{array}$$

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Dy}}}{\nu_{\text{Cs}}} \right) = (-4 \pm 3.9) \times 10^{-8} \text{ yr}^{-1} = 1.5 \times 10^7 \frac{d}{dt} \ln(\alpha) + 0.039 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) - \frac{d}{dt} \ln(m_e / \Lambda_{\text{QCD}}) \quad \text{Berkley, PRL (2007)}$$

$$\frac{d}{dt} \ln \left( \frac{\nu_{\text{Sr}}}{\nu_{\text{Cs}}} \right) = (-7 \pm 18) \times 10^{-16} \text{ yr}^{-1} = -2.77 \frac{d}{dt} \ln(\alpha) + 0.039 \frac{d}{dt} \ln(m_q / \Lambda_{\text{QCD}}) - \frac{d}{dt} \ln(m_e / \Lambda_{\text{QCD}}) \quad \begin{array}{l} \text{Tokyo} \\ \text{JILA} \\ \text{LNE-SYRTE} \\ \text{arXiv (2008)} \end{array}$$

- All optical frequency measurements are against Cs
- Only 2 hyperfine transitions Rb and Cs
- Direct optical vs optical measurements to come

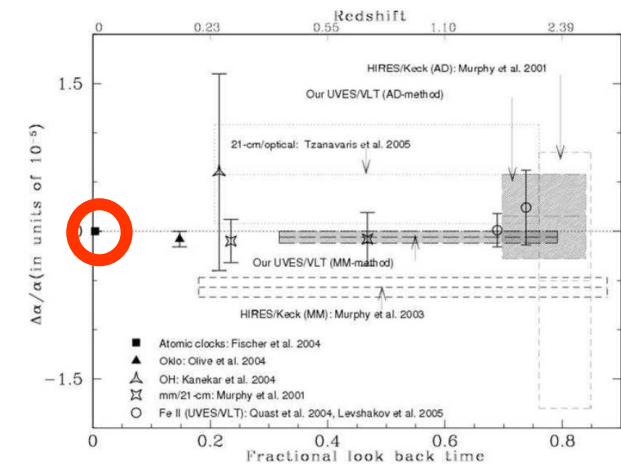
## LABORATORY TESTS: RESULTS

Using a weighted least squares fit to previous data:

$$\frac{d}{dt} \ln(\alpha) = (-3.5 \pm 3.0) \times 10^{-16} \text{ yr}^{-1}$$

$$\frac{d}{dt} \ln(m_q/\Lambda_{\text{QCD}}) = (195 \pm 110) \times 10^{-16} \text{ yr}^{-1}$$

$$\frac{d}{dt} \ln(m_e/\Lambda_{\text{QCD}}) = (24.6 \pm 20) \times 10^{-16} \text{ yr}^{-1}$$



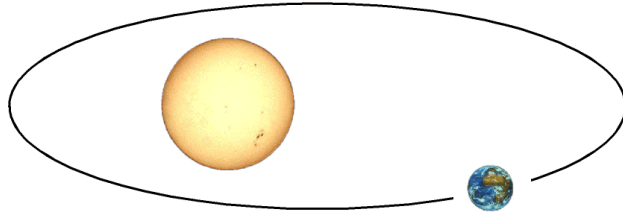
### INDEPENDENT OF COSMOLOGICAL MODELS

- Limit on  $\alpha$  var. is becoming competitive with Oklo ( $\sim 10^{-17} \text{ yr}^{-1}$ ) and Quasar limits ( $\sim 10^{-16} \text{ yr}^{-1}$ ) assuming linear change.
- Assuming linear change, limits on  $m_q$  and  $m_e$  do not exclude the positive result on  $m_e/m_p$  (Reinhold et al. 2006)
- However, still difficult to decorrelate variations of the different constants (correlation coefficients = **-0.53**, **-0.96**, **0.67**).
- More accurate, and more diverse measurements are required!!



# VARIATION OF CONSTANTS WITH GRAVITY

Annual modulation of the Sun gravitation potential at the Earth :



$$\frac{\Delta U(t)}{c^2} = - \underbrace{\frac{GM_S}{c^2 a}}_{\sim 1.6 \cdot 10^{-10}} \epsilon \cos \phi(t).$$

Correlations with varying potential are searched in atomic clock data. Sensitivity of atomic transitions to  $\alpha$ ,  $m_q/\Lambda_{\text{QCD}}$  and  $m_e/\Lambda_{\text{QCD}}$  are as before.

Putative sensitivities of  $\alpha$ ,  $m_q/\Lambda_{\text{QCD}}$  and  $m_e/\Lambda_{\text{QCD}}$  to  $\Delta U/c^2$  are defined:  $\frac{\delta\alpha}{\alpha} = k_\alpha \frac{\Delta U(t)}{c^2}$ ,

*PR A 76, 062104 (2007)*  
*Dy/Cs (Berkley)*  
*Hg<sup>+</sup>/Cs (NIST)*  
*Cs/H(hfs) (NIST, SYRTE, PTB)*

*arxiv:0801.1874 (2008)*  
*Sr/Cs (SYRTE, Tokyo, JILA)*  
*Hg<sup>+</sup>/Cs (NIST)*  
*Cs/H(hfs) (NIST, SYRTE, PTB)*

Parameter	Constraint
$k_\alpha + 0.17k_e$	$(-3.5 \pm 6) \times 10^{-7}$
$ k_\alpha + 0.13k_q $	$< 2.5 \times 10^{-5}$
$k_\alpha + 0.13k_q$	$(-1 \pm 17) \times 10^{-7}$
$k_\alpha$	$(-8.7 \pm 6.6) \times 10^{-6}$
$k_e$	$(4.9 \pm 3.9) \times 10^{-5}$
$k_q$	$(6.6 \pm 5.2) \times 10^{-5}$

$$k_\alpha = (-2.3 \pm 3.1) \times 10^{-6}$$

$$k_\mu = (1.1 \pm 1.7) \times 10^{-5}$$

$$k_q = (1.7 \pm 2.7) \times 10^{-5}$$

# OTHER FUNDAMENTAL TESTS USING CLOCKS AND ULTRA STABLE OSCILLATORS

- ✓ LLI in the photon sector using CSO vs H-maser (SYRTE)

*Wolf et al., Phys. Rev. Lett. 90, 060402 (2003)*

*Wolf et al., Gen. Rel. Grav. 36, 2351 (2004)*

*Wolf et al., Phys. Rev. D 70, 051902(R) (2004)*

Most stringent Kennedy-Thorndicke experiment to date

$$|\beta - \alpha - 1| < 4.7 \times 10^{-7}$$

- ✓ LLI in the photon sector using rotating CORE and rotating CSO (UWA and Berlin Univ., Düsseldorf Univ.)

*Antonini et al., Phys. Rev. A 71, 050101 (2005)*

*Stanwix et al., Phys. Rev. Lett. 95, 040404 (2005)*

*Müller et al., Phys. Rev. Lett. 99, 050401 (2007)*

Most stringent Michelson-Morley experiment to date

$$(\delta - \beta + 1/2) = 9.4(8.1) \times 10^{-11}$$

- ✓ LLI in the matter sector using Zeeman transitions in Cs-hfs in atomic fountains (SYRTE)

*P. Wolf et al., Phys. Rev. Lett. 96, 060801 (2006)*

## FUNDAMENTAL TESTS WITH PHARAO/ACES



- ✓ FM is being developed, yet with strong uncertainty on the development of the project
- ✓ If development continues, current launch schedule is 2014

- ✓ Measurement of the gravitational redshift

At  $H=450\text{km}$ , gravitational redshift is  $4.59 \cdot 10^{-11}$

With clock accuracy of  $10^{-16}$ , the red-shift can be measured at  $3 \times 10^{-6}$

Improvement by  $\sim 30$  over GPA, R. Vessot et al. (1976).

- ✓ Enhanced comparisons between ground clocks through common view comparisons with PHARAO/ACES, down to the  $10^{-17}$  level





# Acknowledgments

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Dawkins, R. Chicireanu,...