

Constraining cosmological models with the Effective Field Theory of Large-Scale Structures

Théo SIMON

IAP - 04/06/2024



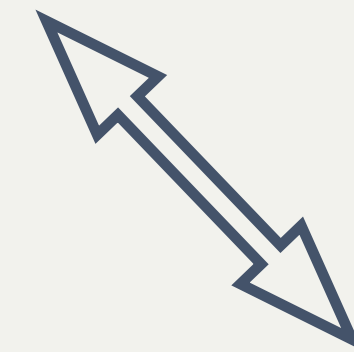
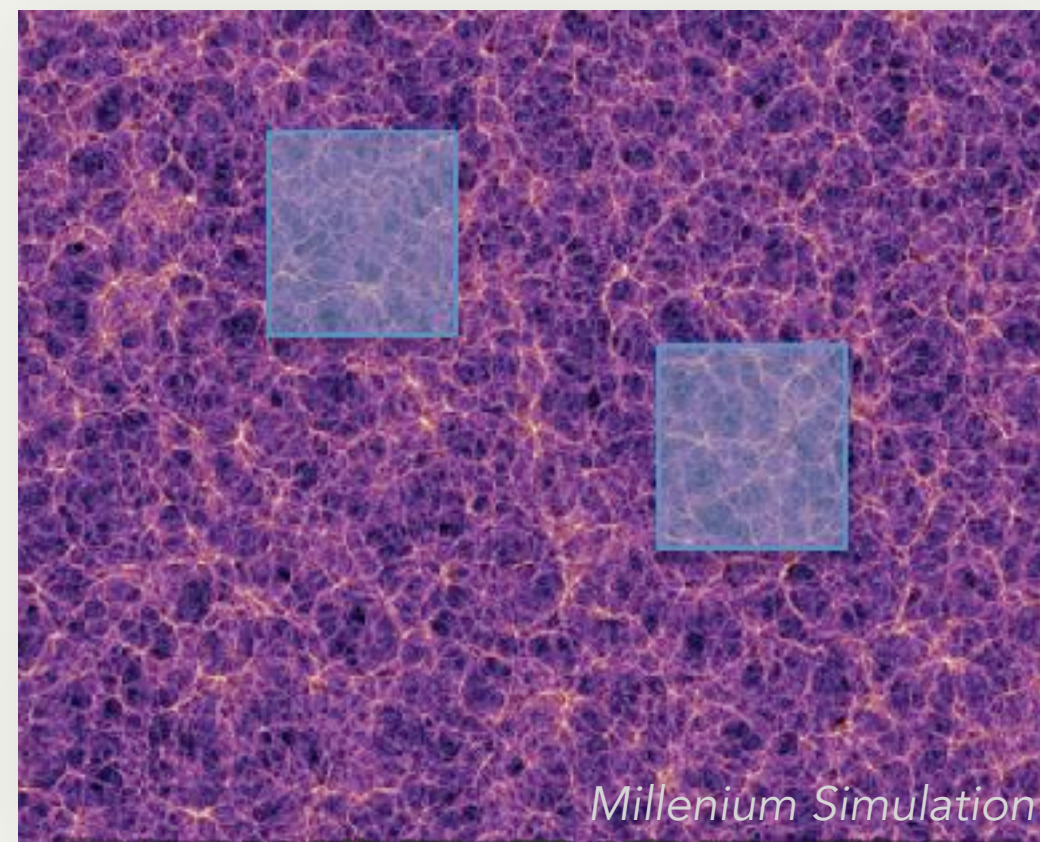
theo.simon@umontpellier.fr

Introduction: the general paradigm

The standard model of
Cosmology (Λ CDM)

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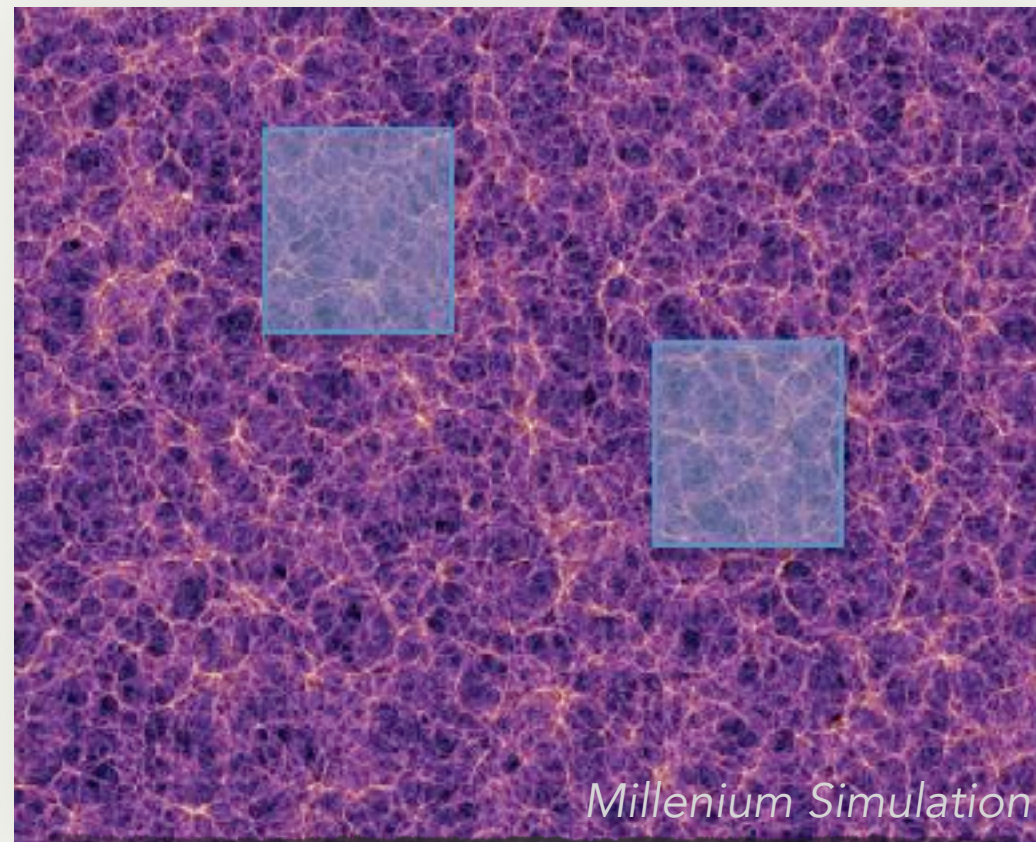
One conjecture: cosmological principle



The standard model of
Cosmology (Λ CDM)

Introduction: the general paradigm

One conjecture: cosmological principle



Two sets of fundamental equations:
Einstein's equations & Boltzmann's equations

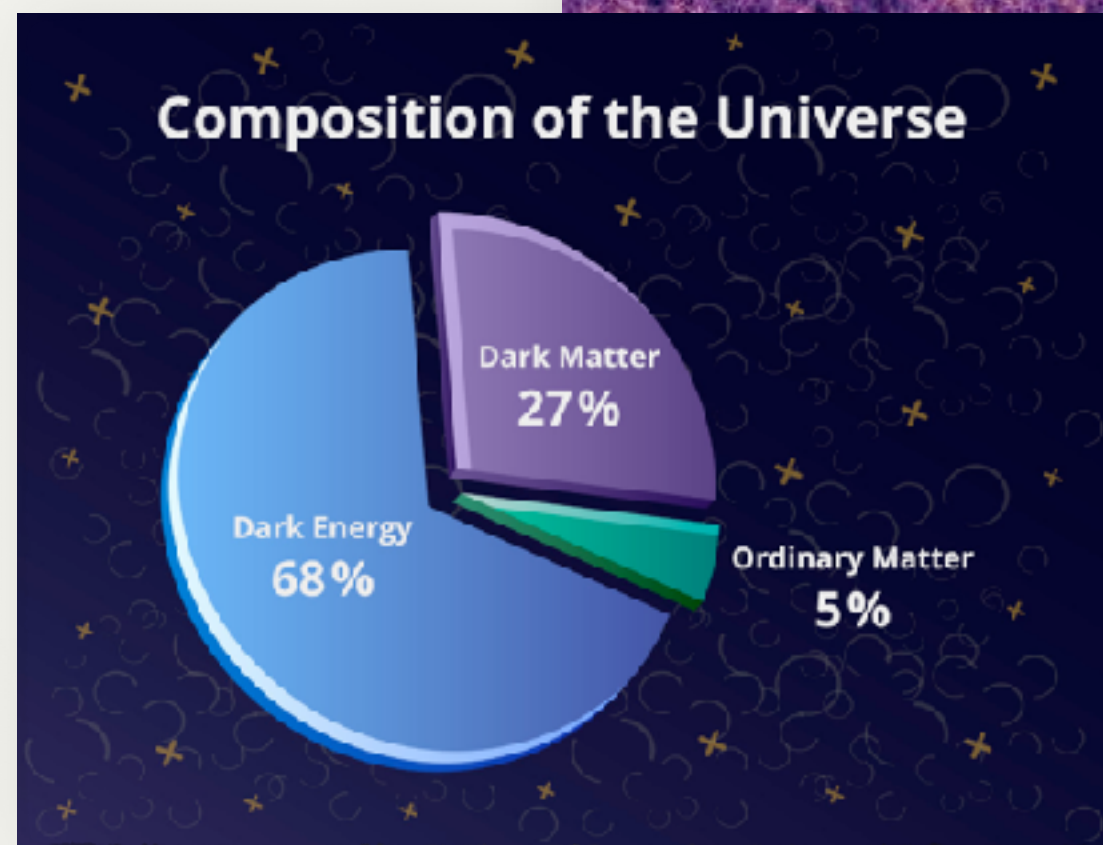
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The standard model of
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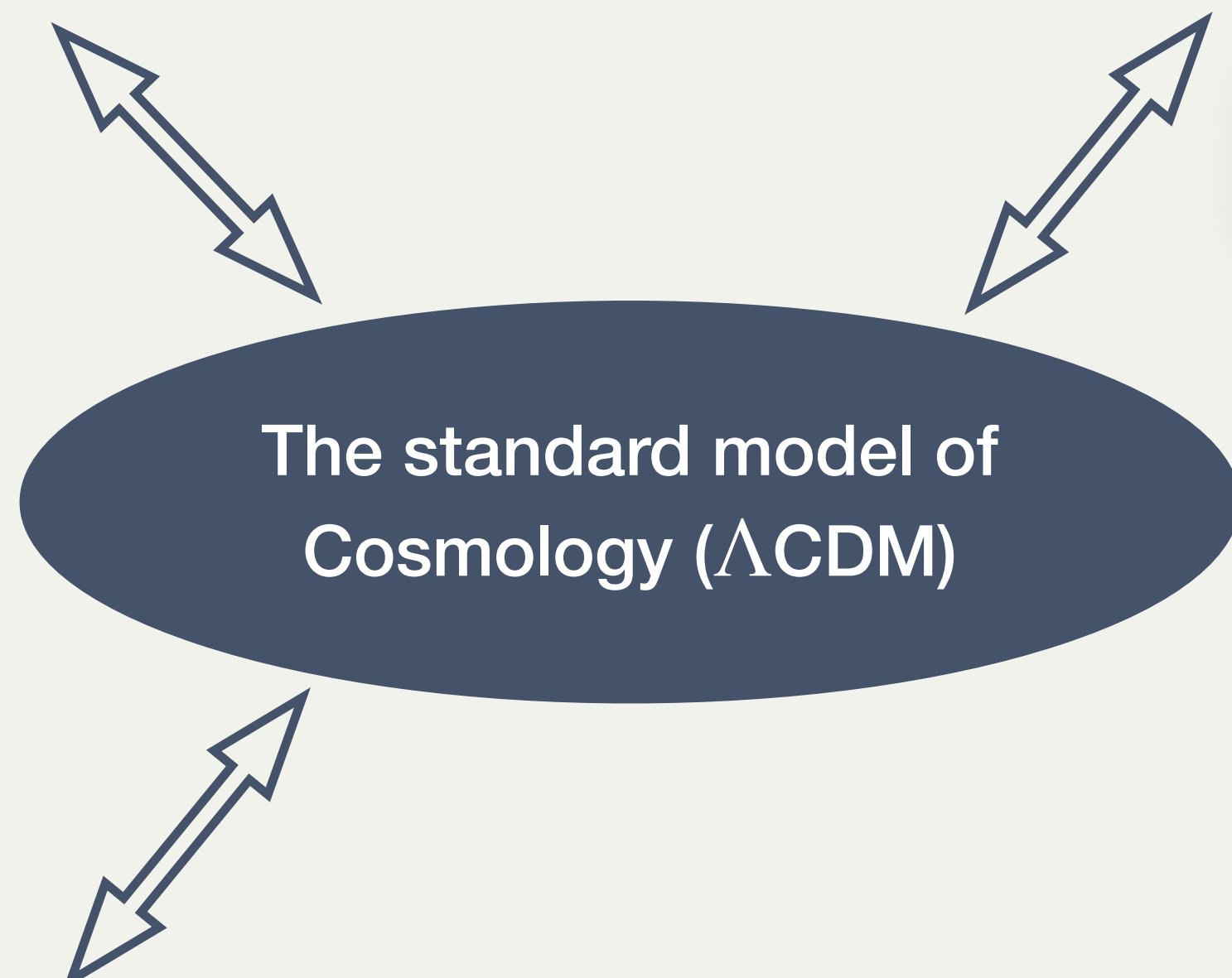
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One conjecture: cosmological principle

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Einstein's equations & Boltzmann's equations



Five main constituents: photons, baryons, neutrinos, dark matter and dark energy



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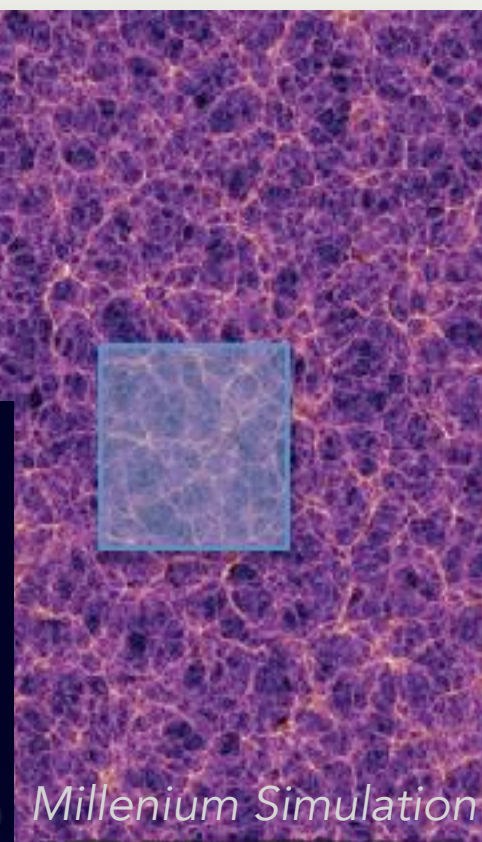
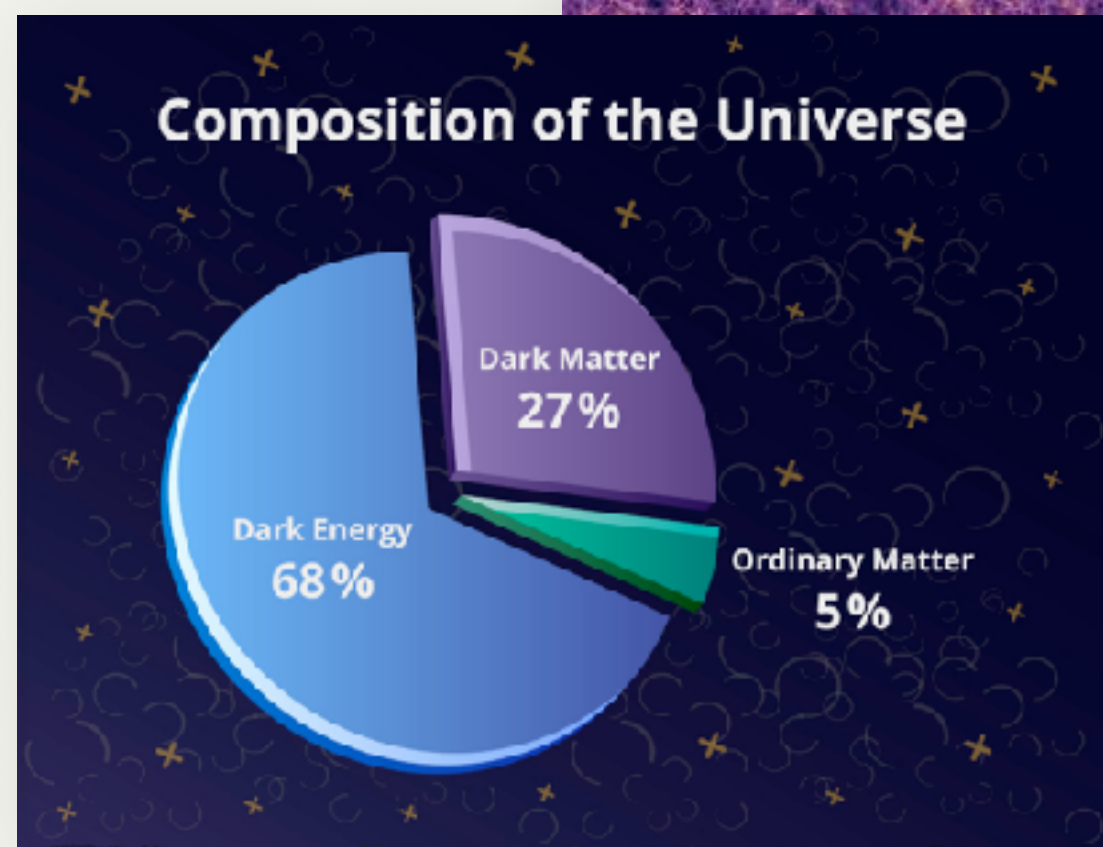
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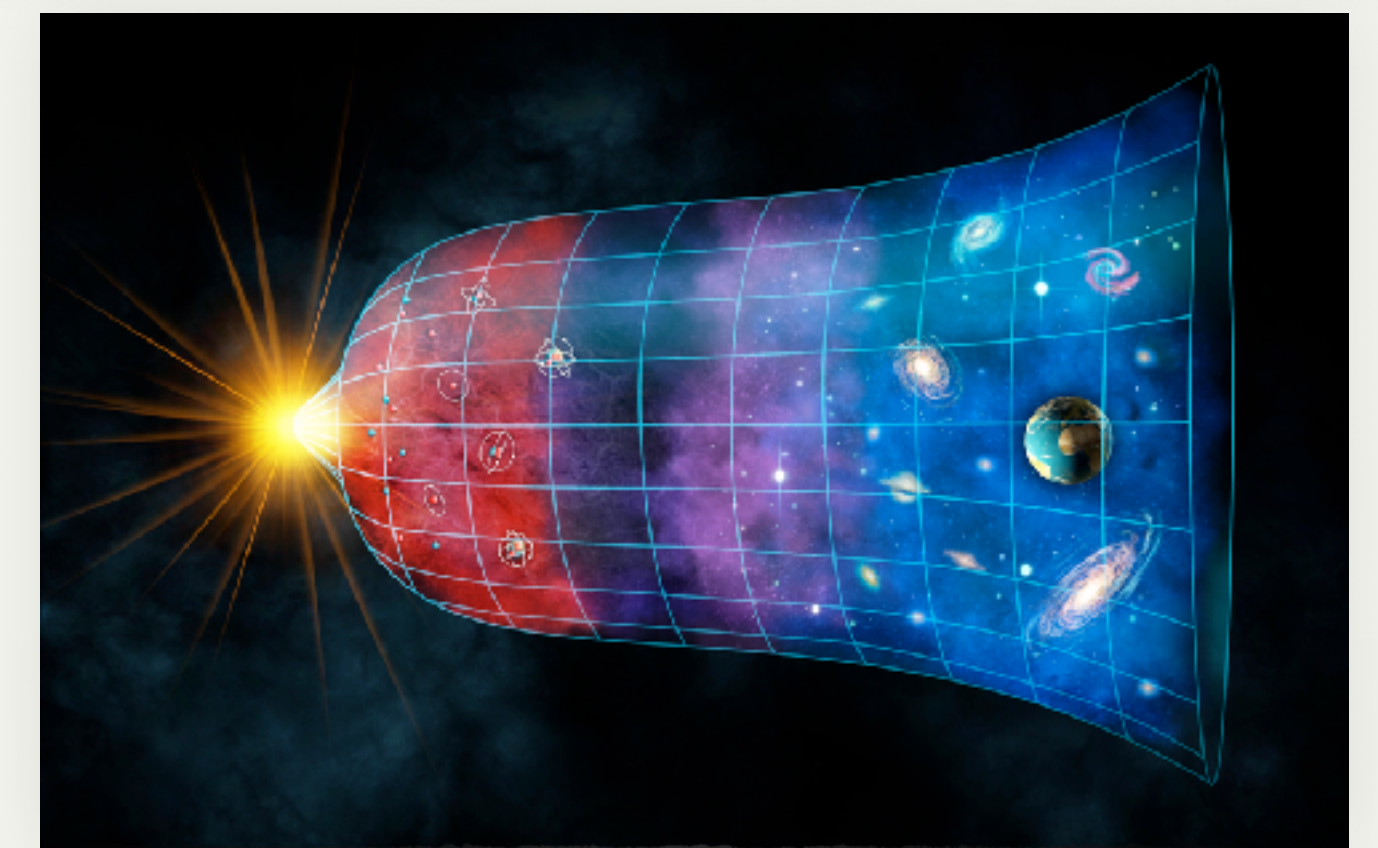
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The standard model of
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Millennium Simulation



Five main constituents: photons, baryons, neutrinos, dark matter and dark energy

Initial conditions: inflation

Introduction: the cosmological data



The 6 Λ CDM parameters

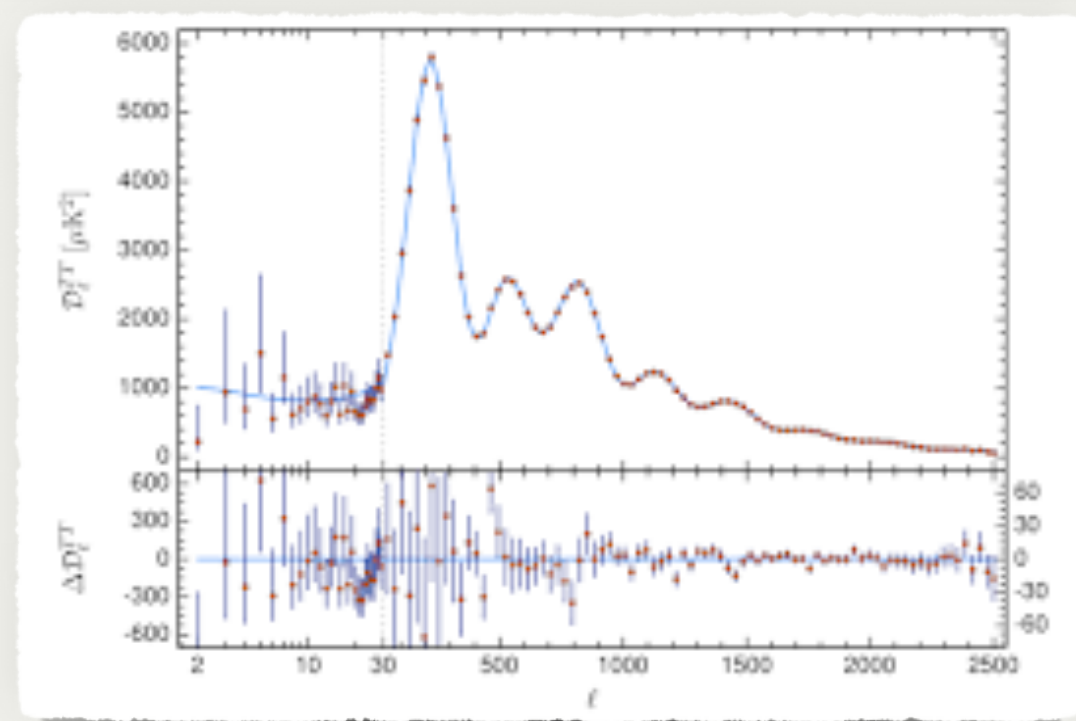
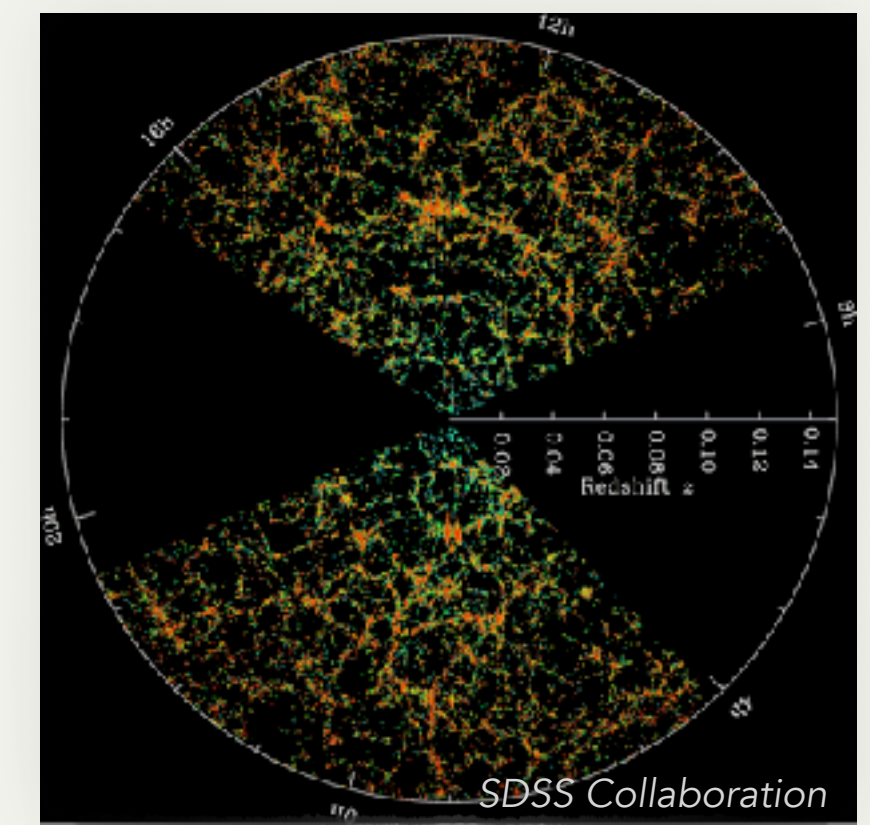
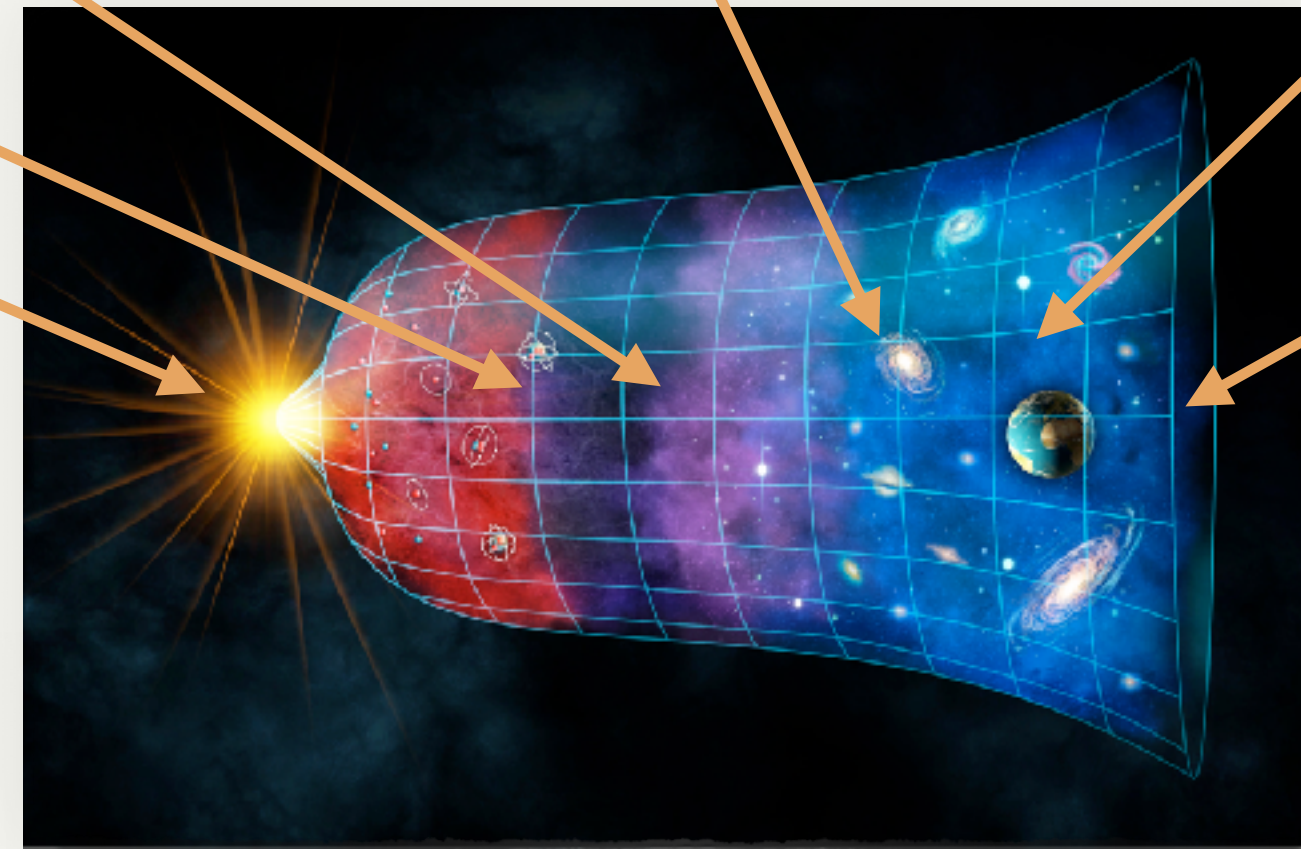
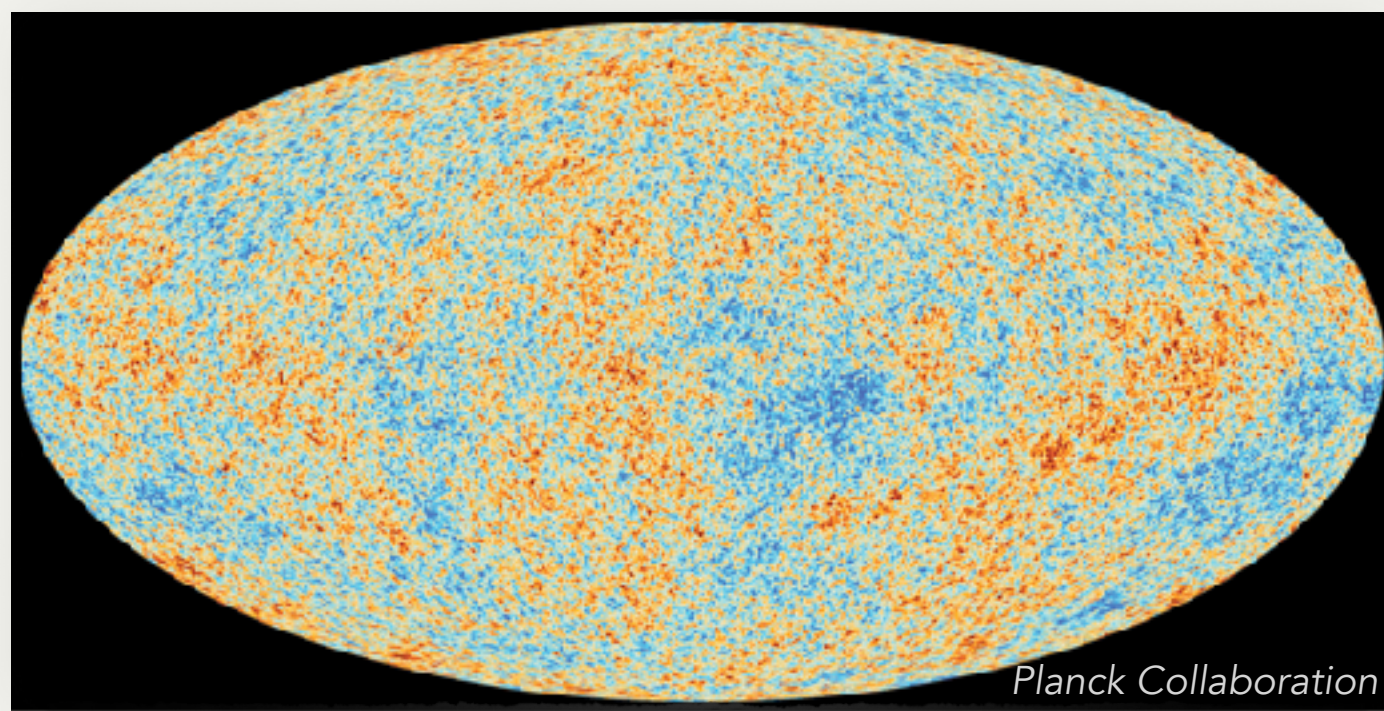
- Initial conditions: A_s, n_s
- Expansion history: H_0, Ω_b and Ω_{cdm}
- Reionization: τ_{reio}

Introduction: the cosmological data

BBN CMB 21cm (?) Galaxy clustering Weak lensing Supernovae

CMB data $z \sim 1000$

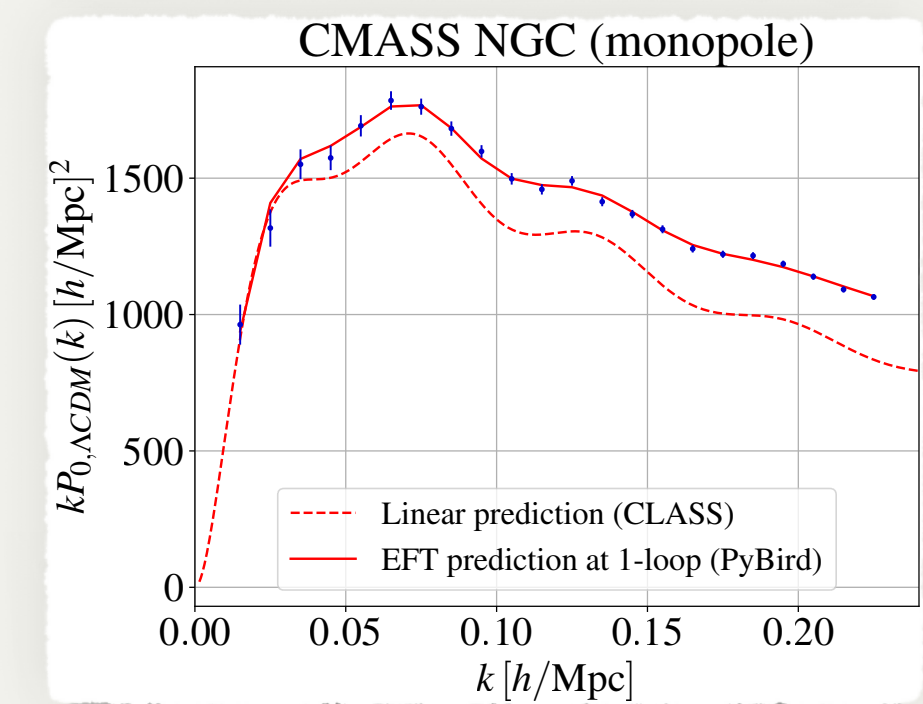
Galaxy clustering data $z \sim 1$



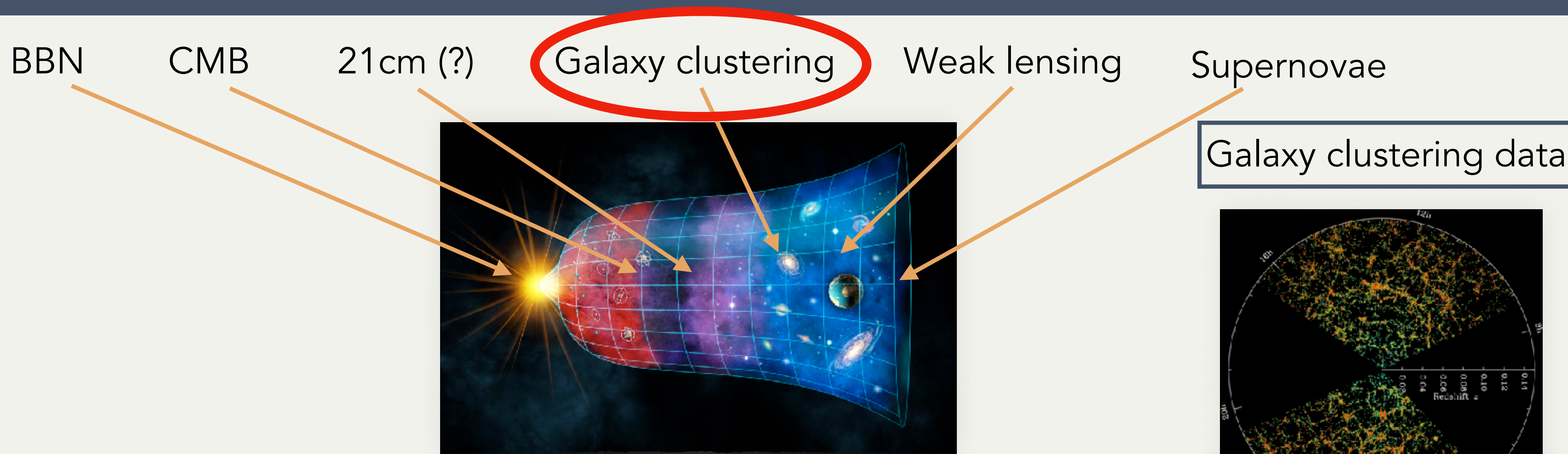
Planck Collaboration [arXiv:1807.06209]

The 6 Λ CDM parameters

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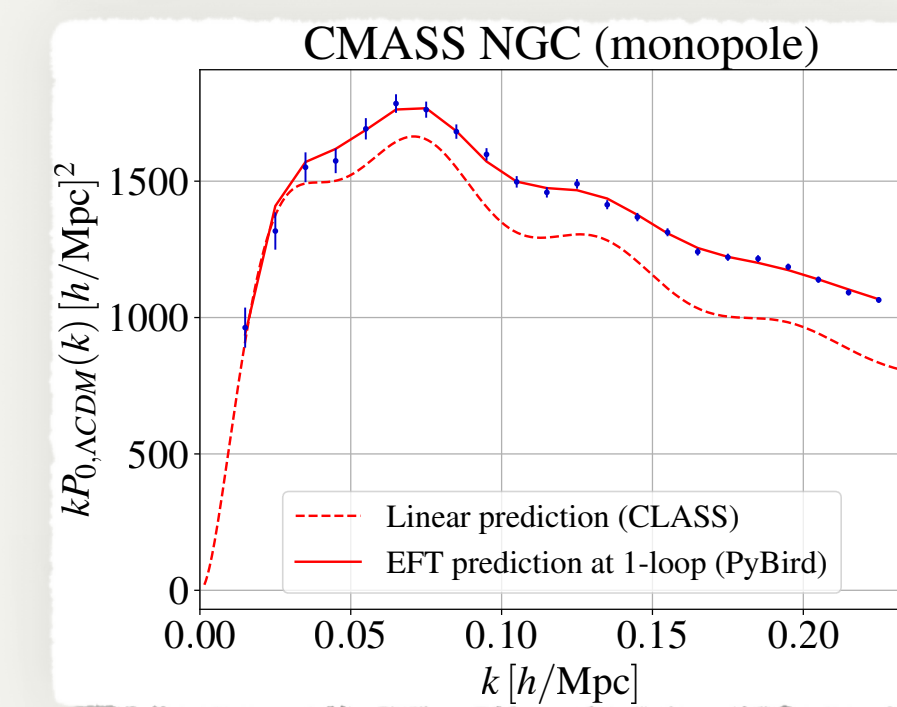
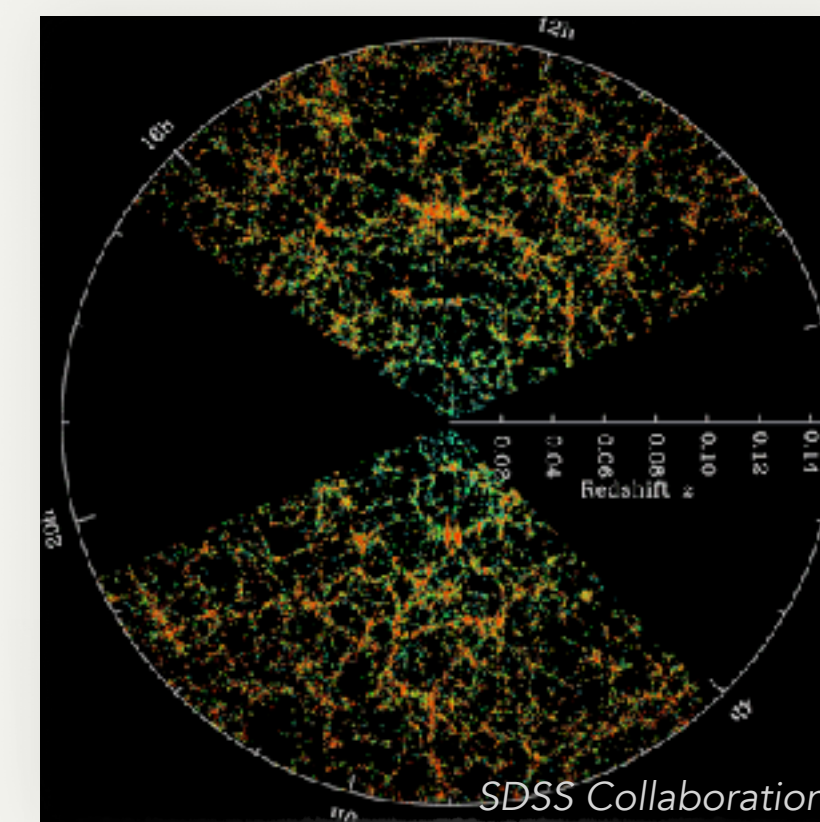


Introduction: the epistemic context



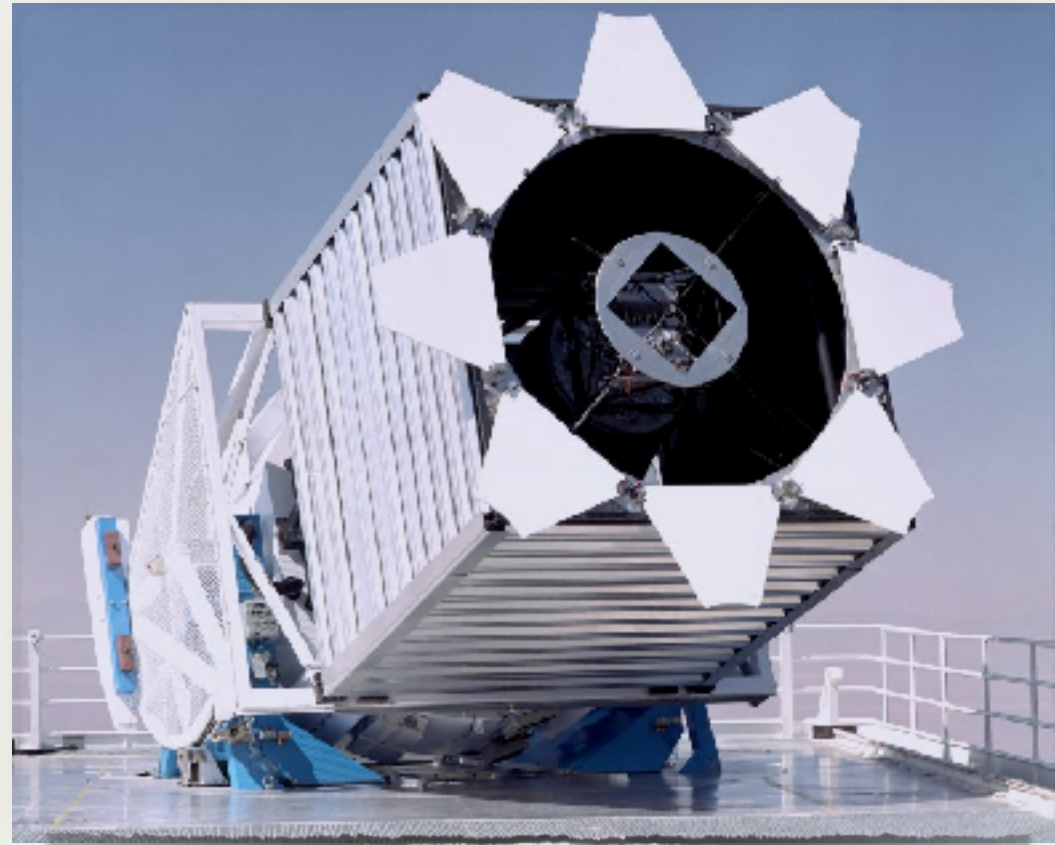
Part I: *How to obtain the most robust and optimised constraints on Λ CDM cosmological parameters from large-scale structure data?*

Part II: *What is the impact of large-scale structure data on the constraints on models beyond Λ CDM?*

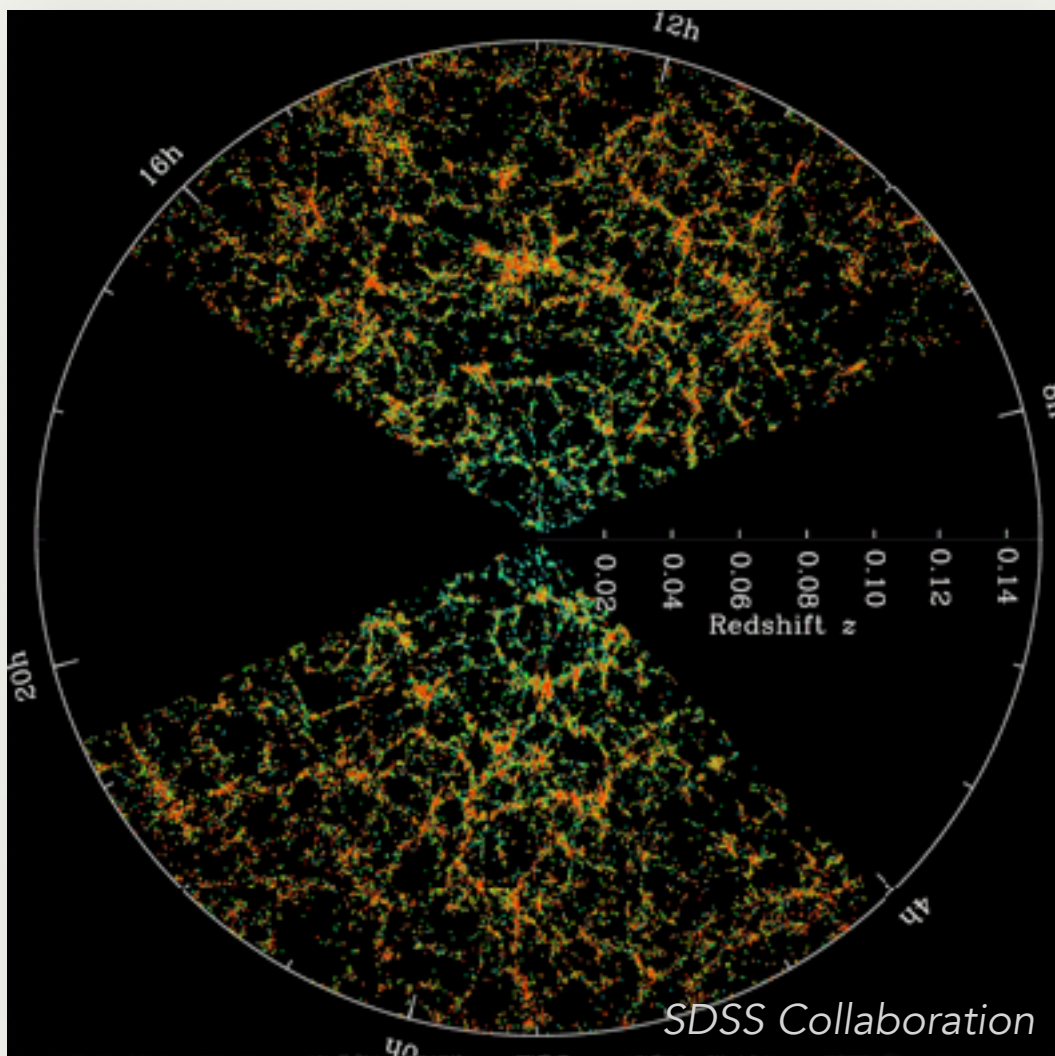


Introduction: large-scale structure data

The Sloan Digital Sky Survey (SDSS)



www.sdss.org



BOSS DR12 LRG (Luminous Red Galaxies)

Galaxies (~ 1.5 million) selected in two redshift ranges:

→ LOWZ (SGC/NGC): $0.2 < z < 0.43$ ($z_{\text{eff}} = 0.32$)

→ CMASS (SGC/NGC): $0.43 < z < 0.7$ ($z_{\text{eff}} = 0.57$)

BOSS Collaboration [arXiv:1607.03155]

eBOSS DR16 QSO

Quasars ($\sim 300\,000$) selected in one redshift range:

$0.8 < z < 2.2$ ($z_{\text{eff}} = 1.5$)

eBOSS Collaboration [arXiv:2007.08991]

I- EFTofLSS applied to (e)BOSS data and its consistency within the Λ CDM model

Based on:

- **TS** *et al.*, JCAP [arXiv:2210.14931]
Cosmological inference from the EFTofLSS: the eBOSS QSO full-shape analysis
- **TS** *et al.*, Phys. Rev. D [arXiv:2208.05929]
Consistency of effective field theory analyses of the BOSS power spectrum
- Emil B. Holm, Laura Herold, **TS**, *et al.*, Phys. Rev. D [arXiv:2208.05929]
Bayesian and frequentist investigation of prior effects in EFTofLSS analyses of full-shape BOSS and eBOSS data

The Effective Field Theory of Large-Scale Structures (EFTofLSS)

Before EFTofLSS...

There are **two main** historical ways of using LSS data:

1. From the linear perturbation theory:

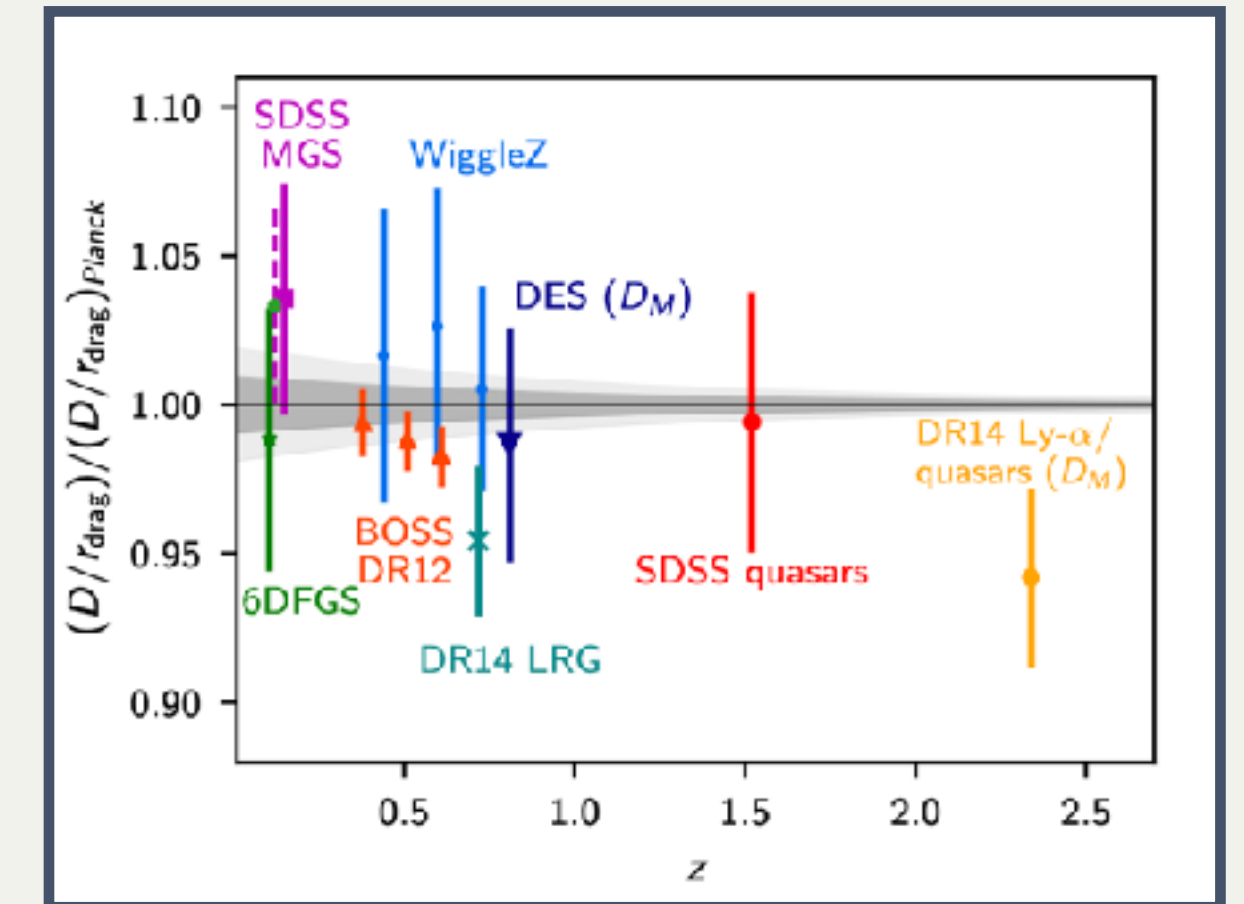
$$P_g(z, k, \mu) \simeq [b_1(z) + f\mu^2]^2 P_m(z, k)$$

Kaiser '87

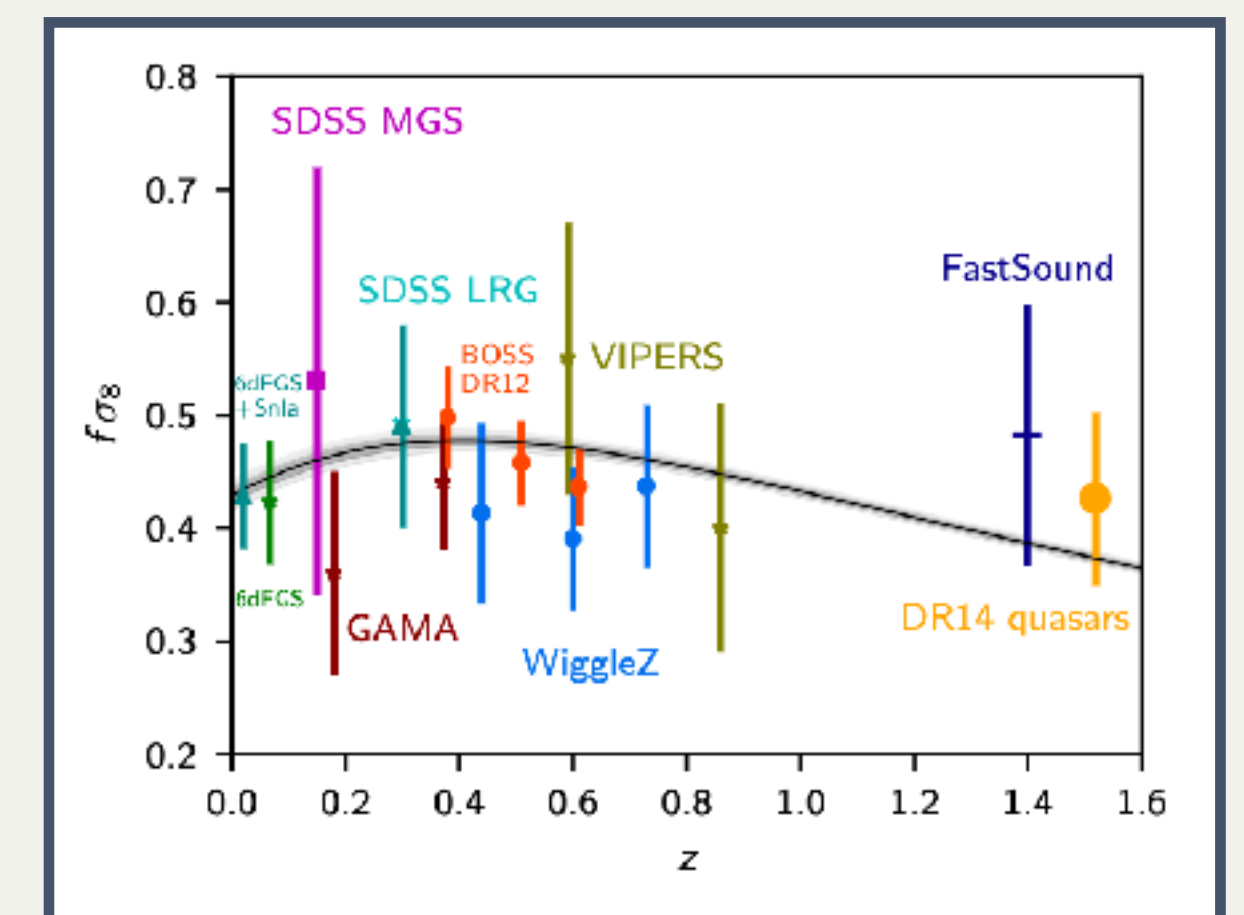
b_1 : bias parameter, f : growth factor and $\mu = \hat{z} \cdot \hat{k}$

2. BAO angles + Redshift Space Distortion (RSD) information: $BAO/f\sigma_8$

LSS collaborations conventionally use the second method



Planck Collaboration [arXiv:1807.06209]



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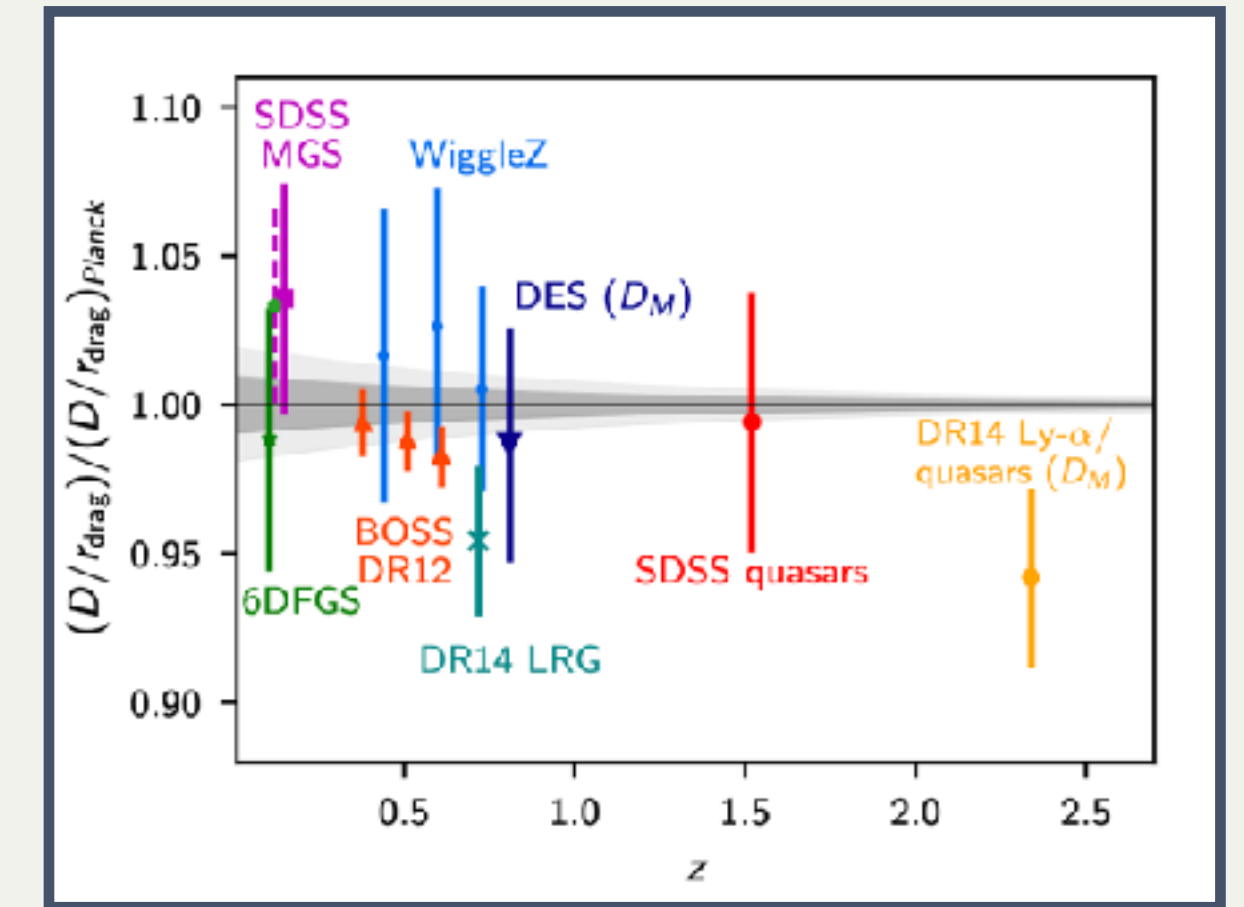
Lack of precision

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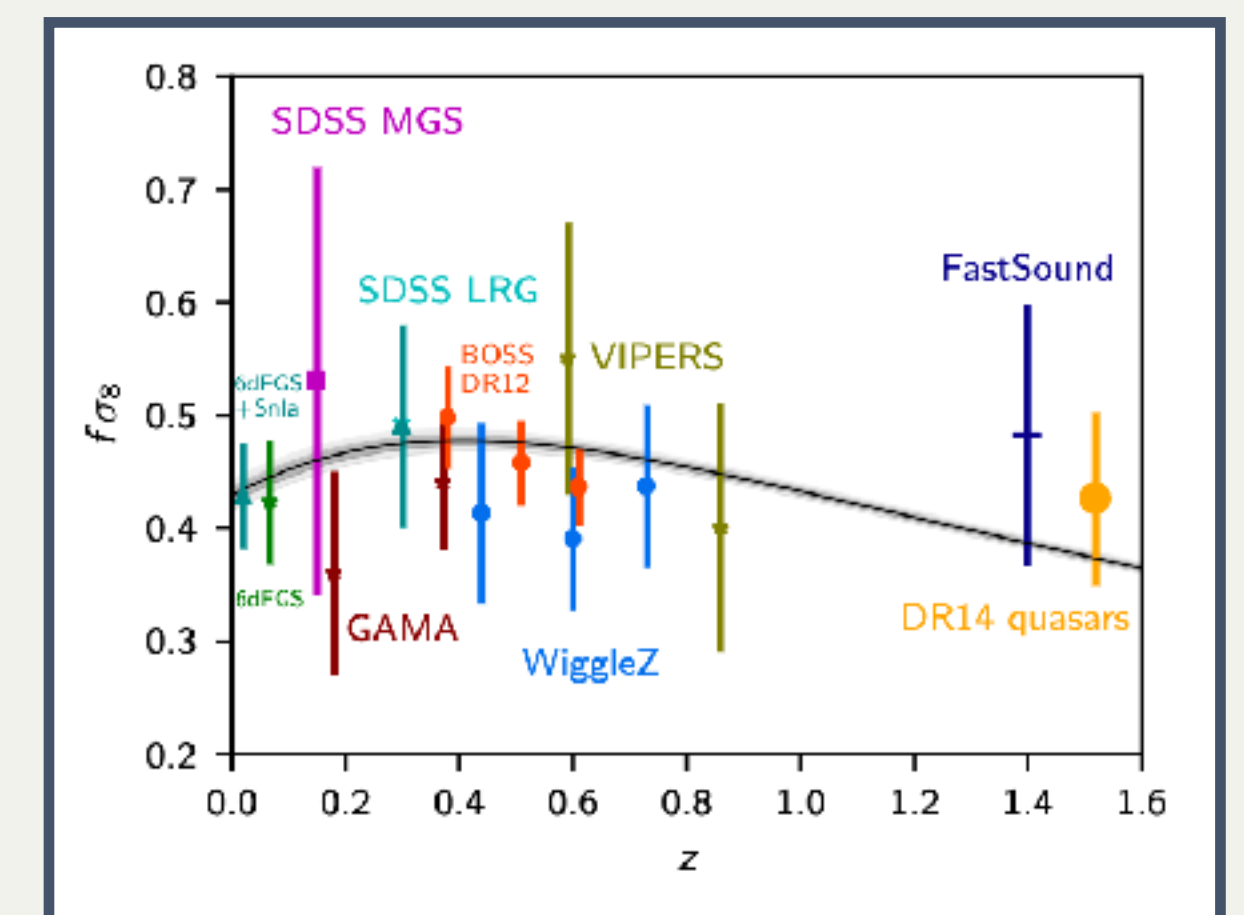
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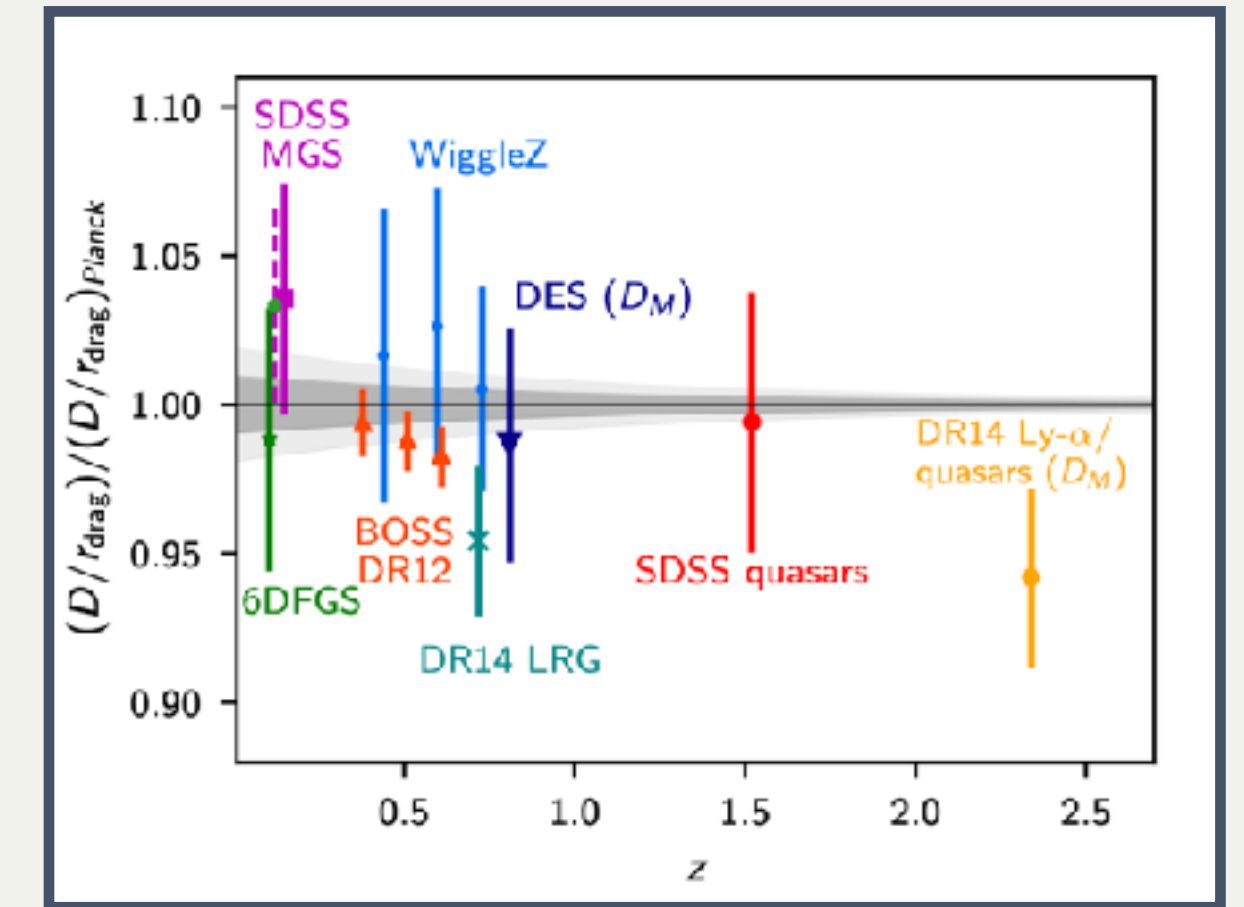
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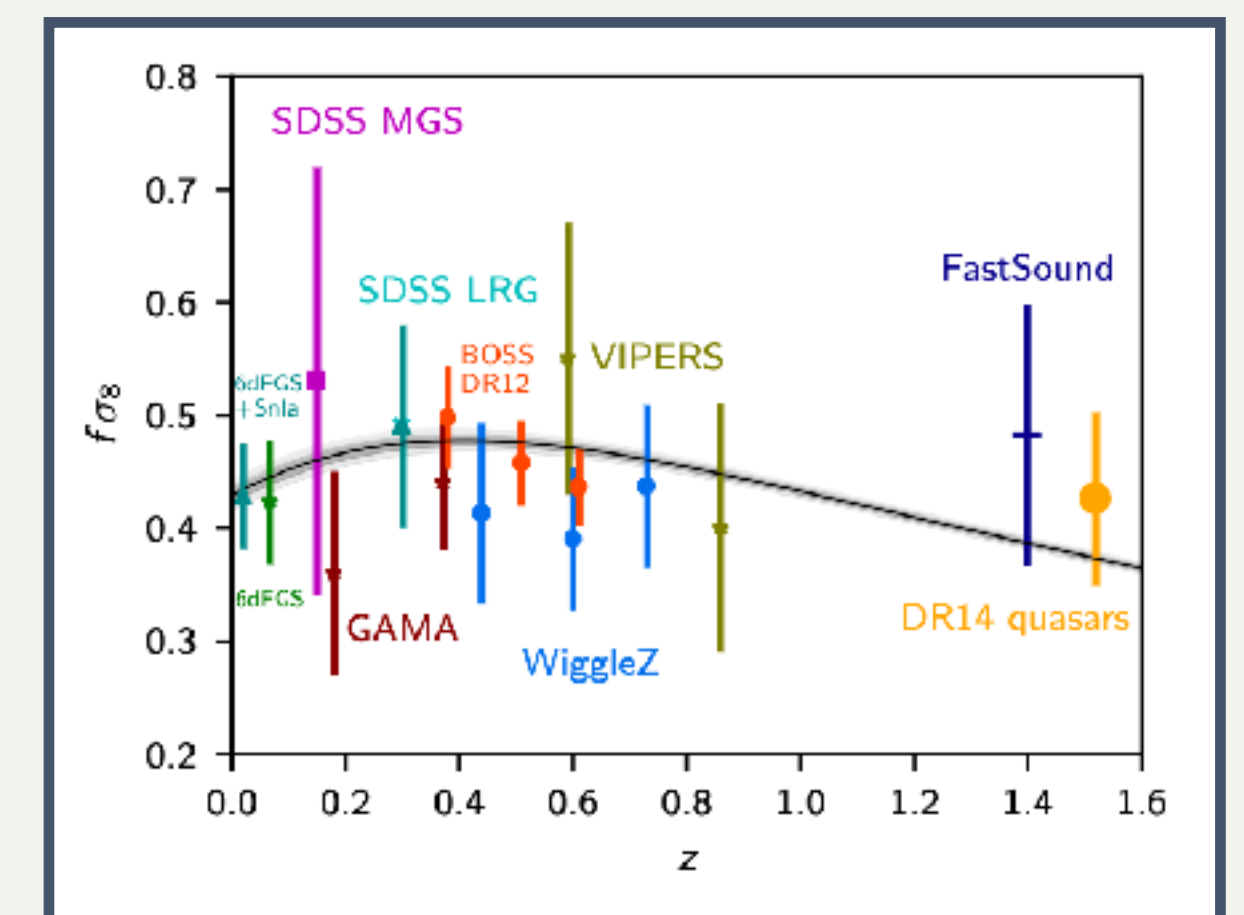
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The effective field theory of large-scale structures (EFTofLSS)

Main ingredients



Carrasco++ [arXiv:1206.2926]

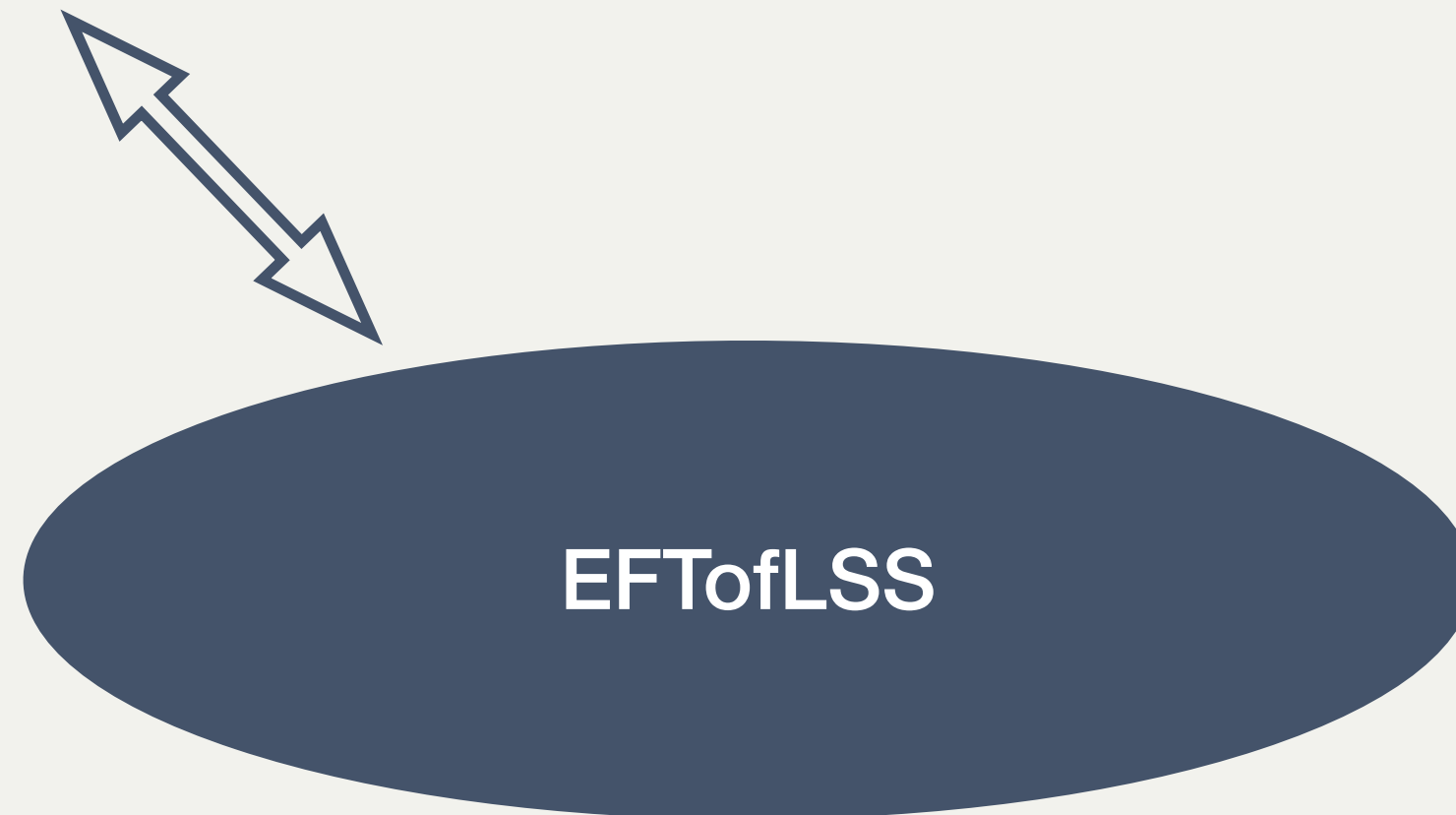
Baumann++ [arXiv:1004.2488]

The effective field theory of large-scale structures (EFTofLSS)

Main ingredients

Solve cosmological equations only for **large-scale physics**

$$\begin{aligned}\delta(\mathbf{k}) &= \delta_l(\mathbf{k}) + \delta_s(\mathbf{k}), \\ \delta_l(\mathbf{k}) &= \delta(\|\mathbf{k}\| < \Lambda^{-1})\end{aligned}$$



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Dark matter: the Vlasov system

$$\begin{aligned}\dot{\delta}_l + \theta_l &= -\delta\theta_l - v^j \partial_j \delta_l, \\ \dot{\theta}_l + aH\theta_l + \nabla^2 \psi_l &= -v^j \partial_j \theta_l - \partial_i v_l^j \cdot \partial_j v_l^i - \partial_j \left(\frac{1}{\rho_l} \partial_i [\tau^{ij}]_\Lambda \right)\end{aligned}$$

EFTofLSS

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EFTofLSS

$$\nabla^2 \psi_l = \frac{3}{2} \Omega_m(a) (aH)^2 \delta_l$$

Gravity: the Poisson equation

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EFTofLSS

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Gravity: the Poisson equation

$$\vec{x} \rightarrow \vec{x} + \vec{a}$$

$$\vec{v} \rightarrow \vec{v} + \partial_t \vec{a}$$

Symmetries: Galilean invariance

Carrasco++ [arXiv:1206.2926]

Baumann++ [arXiv:1004.2488]

The effective field theory of large-scale structures (EFTofLSS)

Main steps

Step by Step...

1- **Solve** dark matter equations **perturbatively**: $\delta_l(\vec{x}, t) = \delta_l^{(1)}(\vec{x}, t) + \delta_l^{(2)}(\vec{x}, t) + \dots + \delta_l^{(n)}(\vec{x}, t)$ *Bernardeau++ '01*

2- Obtain the **mildly non-linear matter power spectrum**:

Carrasco++ [arXiv:1206.2926]

$$P_m(k, \tau) = P_{11}(k, \tau) + P_{22}^\Lambda(k, \tau) + 2P_{13}^\Lambda(k, \tau) + 2P_{e_{\text{comb}}}^\Lambda(k, \tau)$$

Senatore [arXiv:1406.7843]

Mirbabayi++ [arXiv:1412.5169]

3- Write down **all possible operators in the galaxy bias expansion**: $\delta_g = b_1 \delta_l^{(1)} + b_2 \delta_l^{(2)} + R_*^2 \partial^2 \delta_l^{(1)} + \dots$

4- Take into account the **redshift-space distortion** (RSD) effect (to subtract the contribution of the peculiar velocity of the galaxies to the cosmological redshift) *Senatore++ [arXiv:1409.1225]*

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Tree-level

One-loop level

Counterterm

Senatore [arXiv:1406.7843]

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The effective field theory of large-scale structures (EFTofLSS)

Galaxy power spectrum

The **galaxy power spectrum** in the EFTofLSS framework:

$$P_g(k, \mu) \simeq [b_1 + f\mu^2]^2 P_m(k) = Z_1(\mu)^2 P_m(k)$$



We go from 1 to 10 free parameters

$$\begin{aligned} P_g(k, \mu) = & Z_1(\mu)^2 P_{11}(k) + 2Z_1(\mu) P_{11}(k) \left(c_{\text{ct}} \frac{k^2}{k_M^2} + c_{r,1} \mu^2 \frac{k^2}{k_R^2} + c_{r,2} \mu^4 \frac{k^2}{k_R^2} \right) \\ & + 2 \int \frac{d^3 q}{(2\pi)^3} Z_2(\mathbf{q}, \mathbf{k} - \mathbf{q}, \mu)^2 P_{11}(|\mathbf{k} - \mathbf{q}|) P_{11}(q) + 6Z_1(\mu) P_{11}(k) \int \frac{d^3 q}{(2\pi)^3} Z_3(\mathbf{q}, -\mathbf{q}, \mathbf{k}, \mu) P_{11}(q) \\ & + \frac{1}{\bar{n}_g} \left(c_{\epsilon,0} + c_{\epsilon}^{\text{mono}} \frac{k^2}{k_M^2} + 3c_{\epsilon}^{\text{quad}} \left(\mu^2 - \frac{1}{3} \right) \frac{k^2}{k_M^2} \right), \end{aligned}$$

Carrasco++ [arXiv:1206.2926] ; Baumann++ [arXiv:1004.2488]

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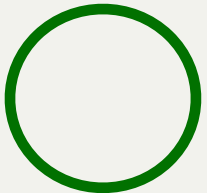
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$$P_g(k, \mu) = Z_1(\mu)^2 P_{11}(k) + 2Z_1(\mu)P_{11}(k) \left(c_{\text{cl}} \frac{k^2}{k_M^2} + c_{r,1} \mu^2 \frac{k^2}{k_R^2} + c_{r,2} \mu^4 \frac{k^2}{k_R^2} \right) \\ + 2 \int \frac{d^3q}{(2\pi)^3} Z_2(\mathbf{q}, \mathbf{k} - \mathbf{q}, \mu)^2 P_{11}(|\mathbf{k} - \mathbf{q}|) P_{11}(q) + 6Z_1(\mu)P_{11}(k) \int \frac{d^3q}{(2\pi)^3} Z_3(\mathbf{q}, -\mathbf{q}, \mathbf{k}, \mu) P_{11}(q) \\ + \frac{1}{\bar{n}_g} \left(c_{\epsilon,0} + c_{\epsilon}^{\text{mono}} \frac{k^2}{k_M^2} + 3c_{\epsilon}^{\text{quad}} \left(\mu^2 - \frac{1}{3} \right) \frac{k^2}{k_M^2} \right),$$

Carrasco++ [arXiv:1206.2926] ; Baumann++ [arXiv:1004.2488]

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2 renormalization scales

 Renormalization scale controlling the **spatial derivative expansion**, given by the typical size of a **virialized halo**

 Renormalization scale of the **velocity products** appearing in the redshift-space expansion

The effective field theory of large-scale structures (EFTofLSS)

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Carrasco++ [arXiv:1206.2926] ; Baumann++ [arXiv:1004.2488]

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$P_g(k, \mu)$ can be determined directly
from $P_{11}(k) = P_m^{\text{lin}}(k)$

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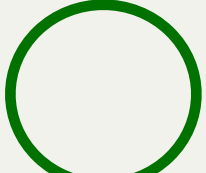
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10 EFT parameters

 **4 parameters** b_i ($i = 1, 2, 3, 4$) to describe the **galaxy bias** which arises from the one-loop contributions

 **3 parameters** corresponding to **counterterms** (c_{ct} linear combination of a higher derivative bias and the dark matter sound speed, while $c_{r,1}$ and $c_{r,2}$ are the redshift-space counterterms)

 **3 parameters** which describe **stochastic terms**

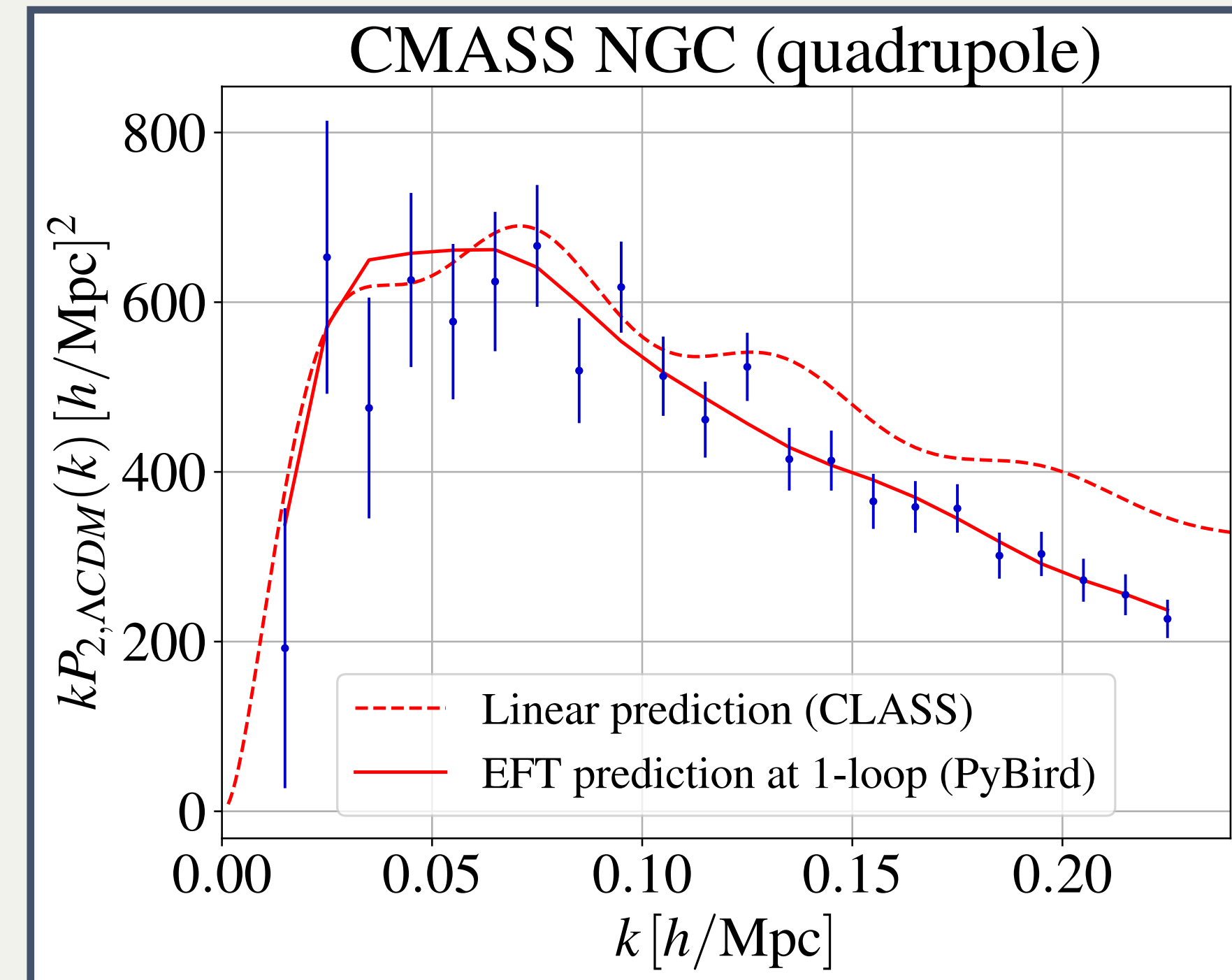
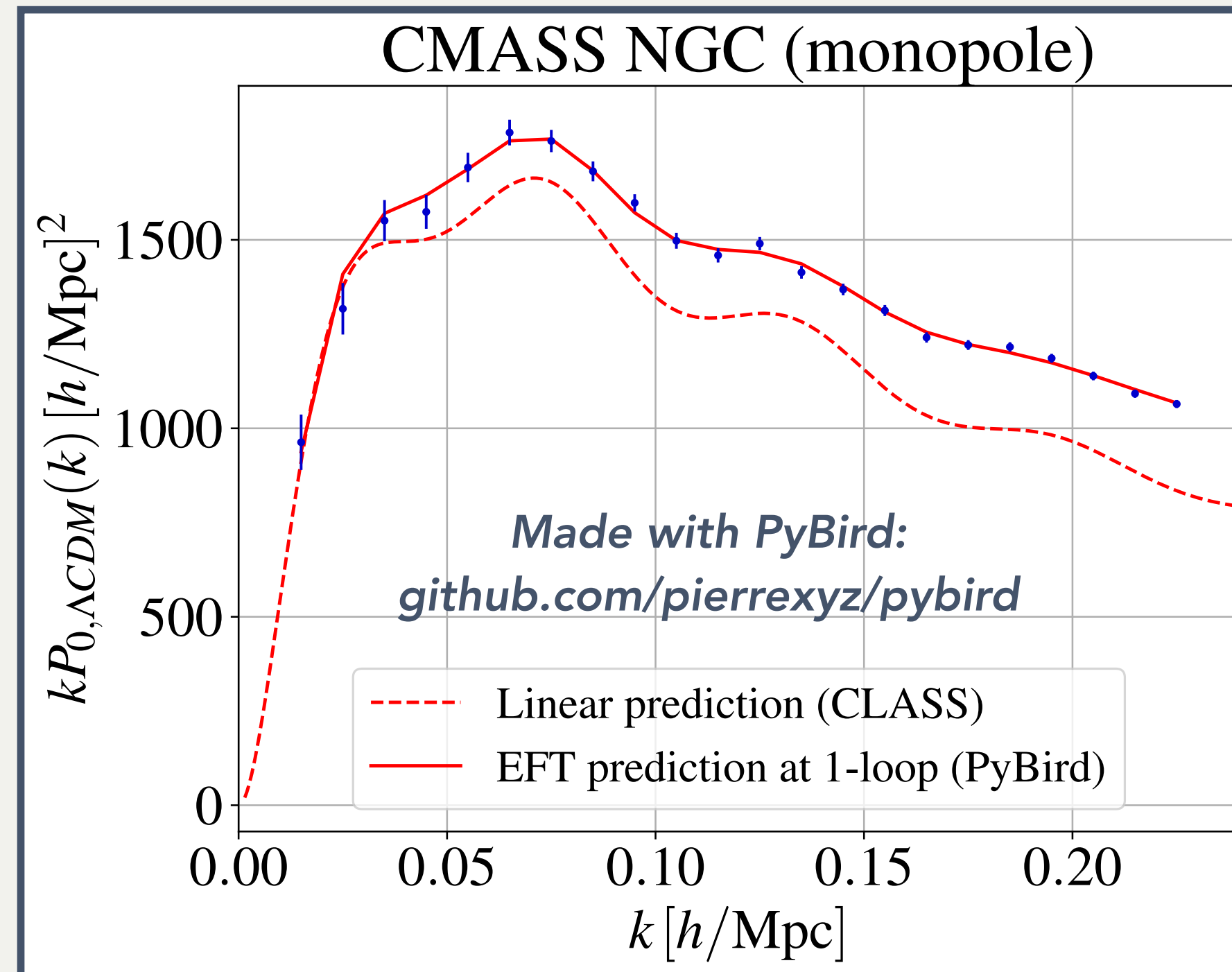
The effective field theory of large-scale structures (EFTofLSS)

Application to BOSS data

Multipoles of the galaxy power spectrum, obtained through a **Legendre** polynomials (\mathcal{L}_ℓ) decomposition:

$$P_g(z, k, \mu) = \sum_{\ell \text{ even}} \mathcal{L}_\ell(\mu) P_\ell(z, k)$$

→ the two main contributions to $P_g(z, k, \mu)$ are the **monopole** ($\ell = 0$) and the **quadrupole** ($\ell = 2$)



D'Amico++ [arXiv:1909.05271] ; Colas++ [arXiv:1909.07951]
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The effective field theory of large-scale structures (EFTofLSS)

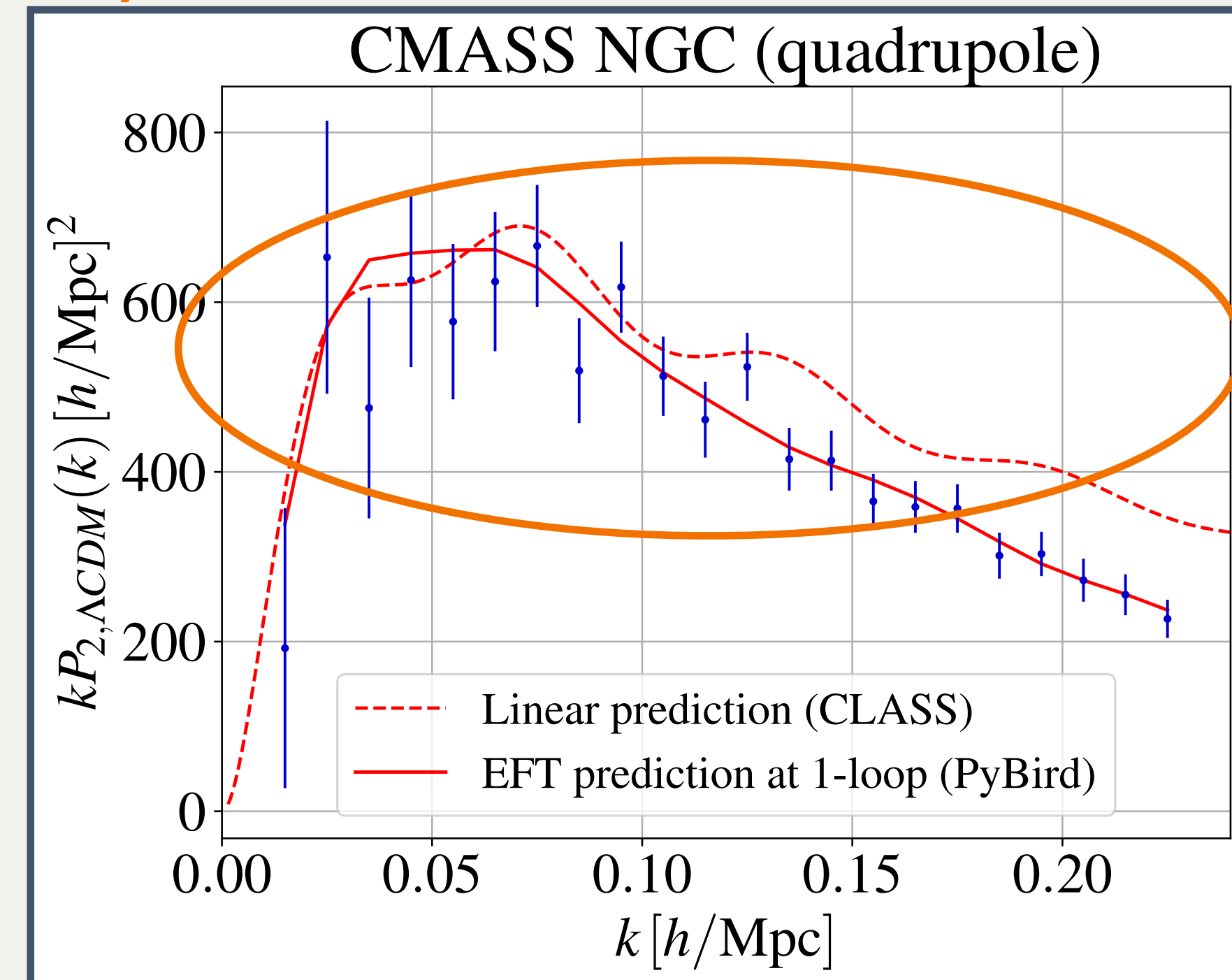
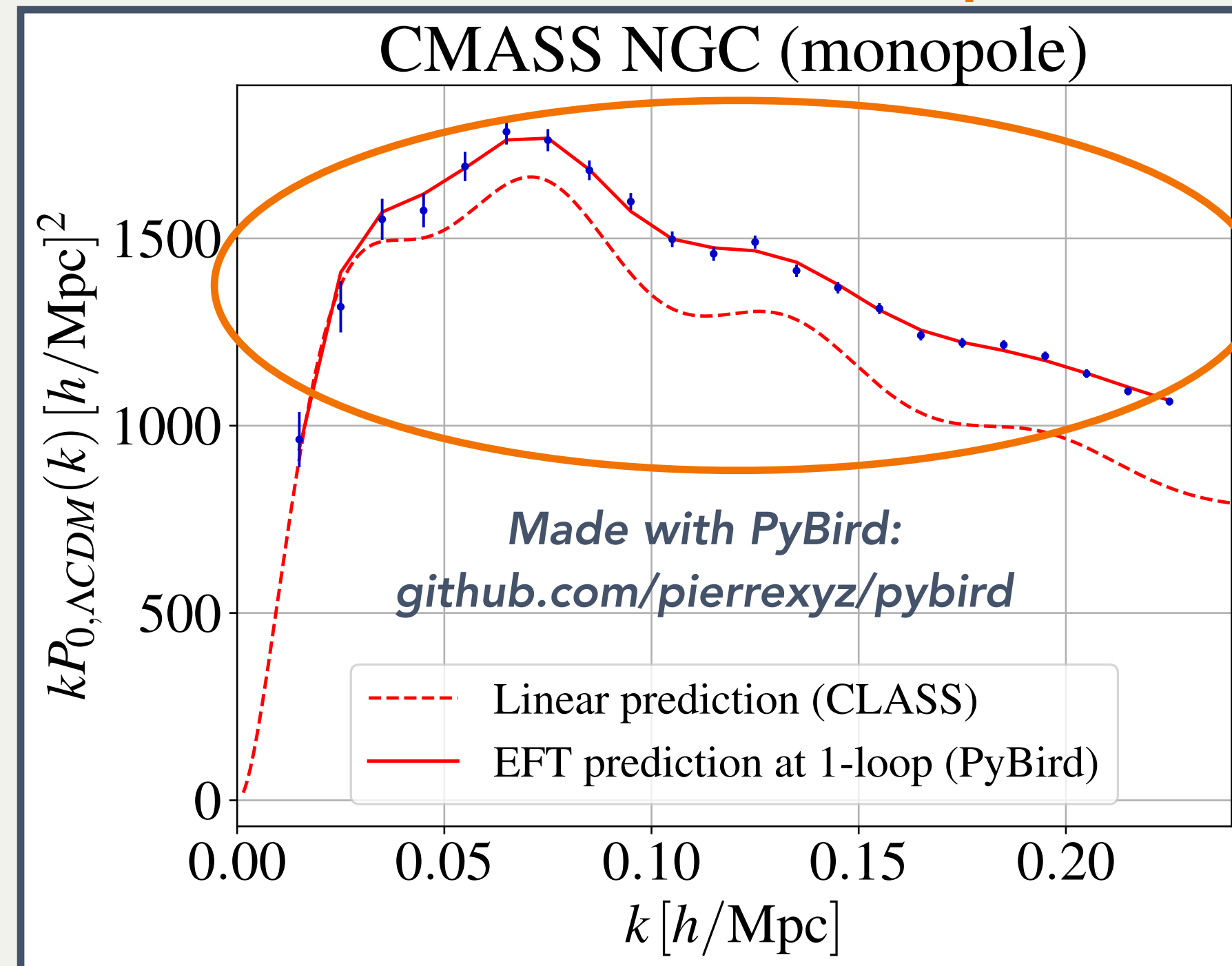
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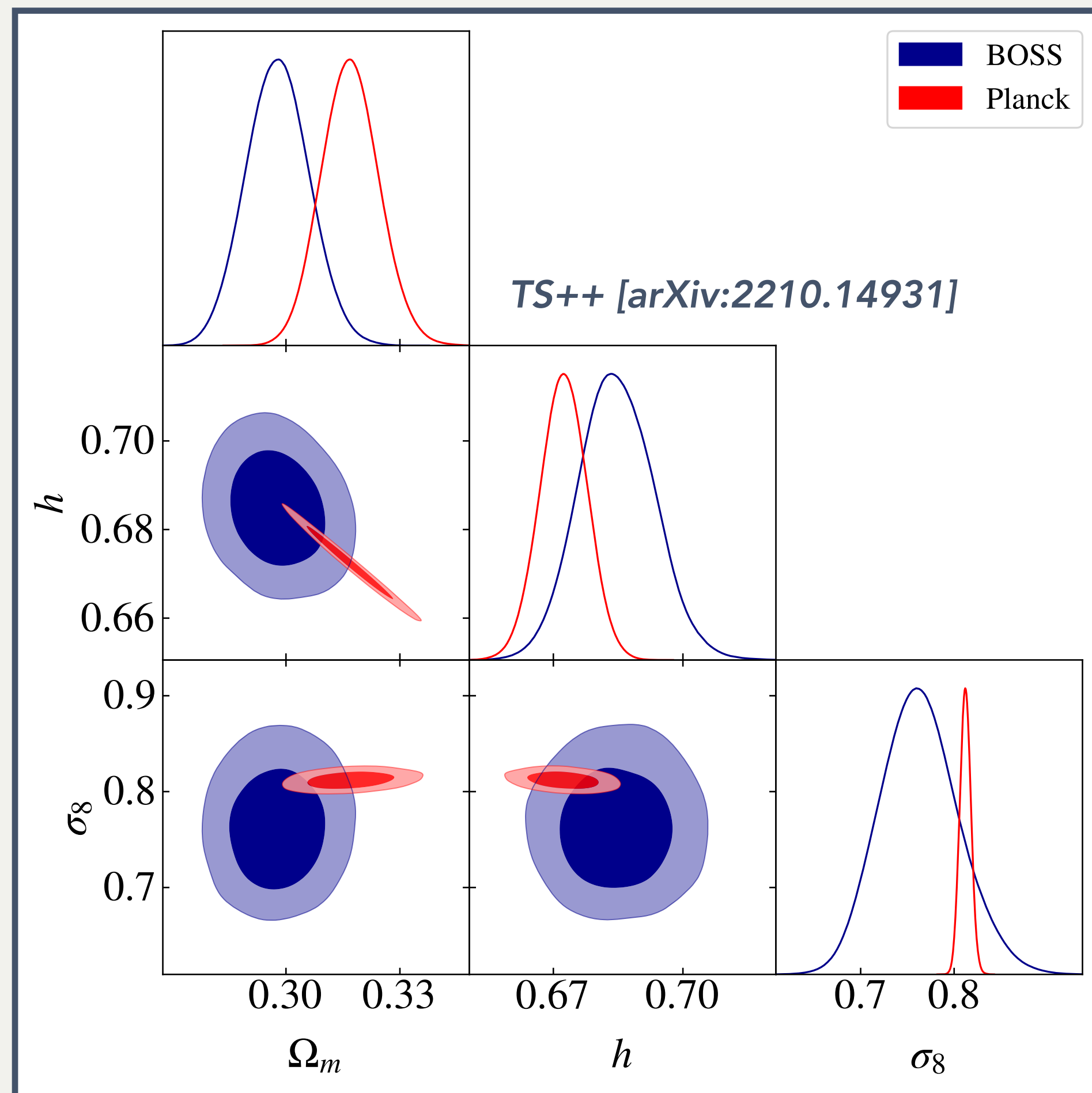
Improvement in precision!



D'Amico++ [arXiv:1909.05271] ; Colas++ [arXiv:1909.07951]
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The effective field theory of large-scale structures (EFTofLSS)

Application to BOSS data



The EFTofLSS analysis of BOSS data allows to determine Ω_m and h at a **precision only 10 % and 60 %** worse than Planck

This is ~ 5.4 (for Ω_m) and ~ 3.2 (for h) times better than the BAO/ $f\sigma_8$ analysis

See also D'Amico++ [arXiv:1909.05271] ; Philcox++ [arXiv:2002.04035]

On the consistency of EFTofLSS

Presentation of the problem

There are **several codes** in the literature with **different parametrizations**:

→ **CLASS PT** with the **EC** parametrization *Chudaykin++ [arXiv:2004.10607]*

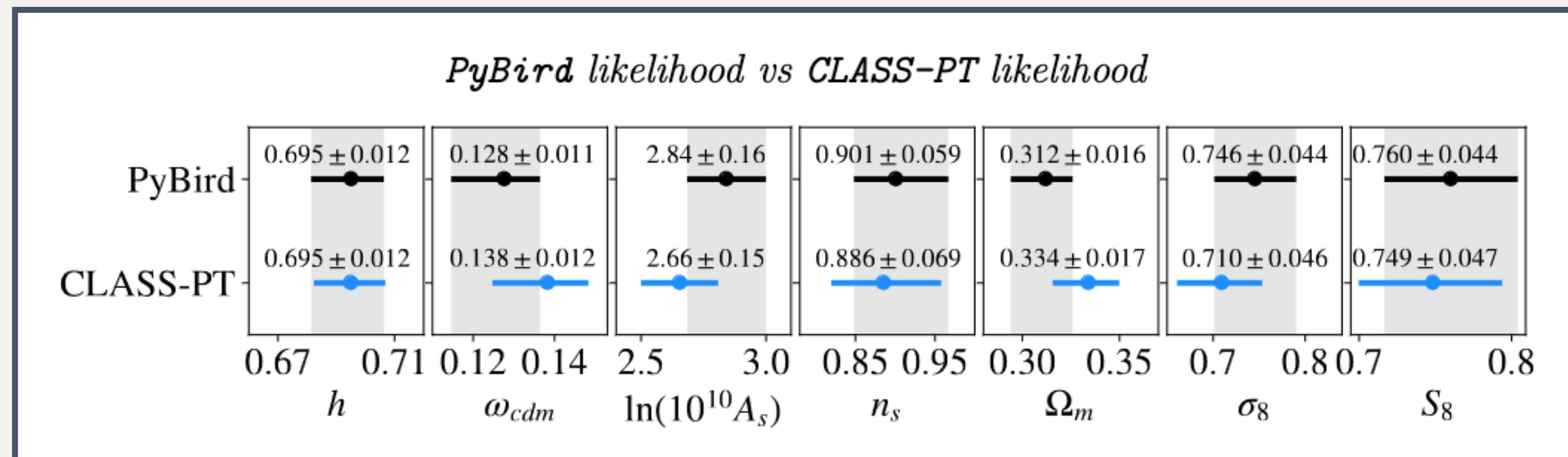
→ **PyBird** with the **WC** parametrization *D'Amico++ [arXiv:2003.07956]*

(+ **Velocileptors** + **CLASS-OneLoop**)

Chen++ [arXiv:2005.00523] ; Linde++ [arXiv:2402.09778]

→ these codes use **different sets of priors** on EFT parameters

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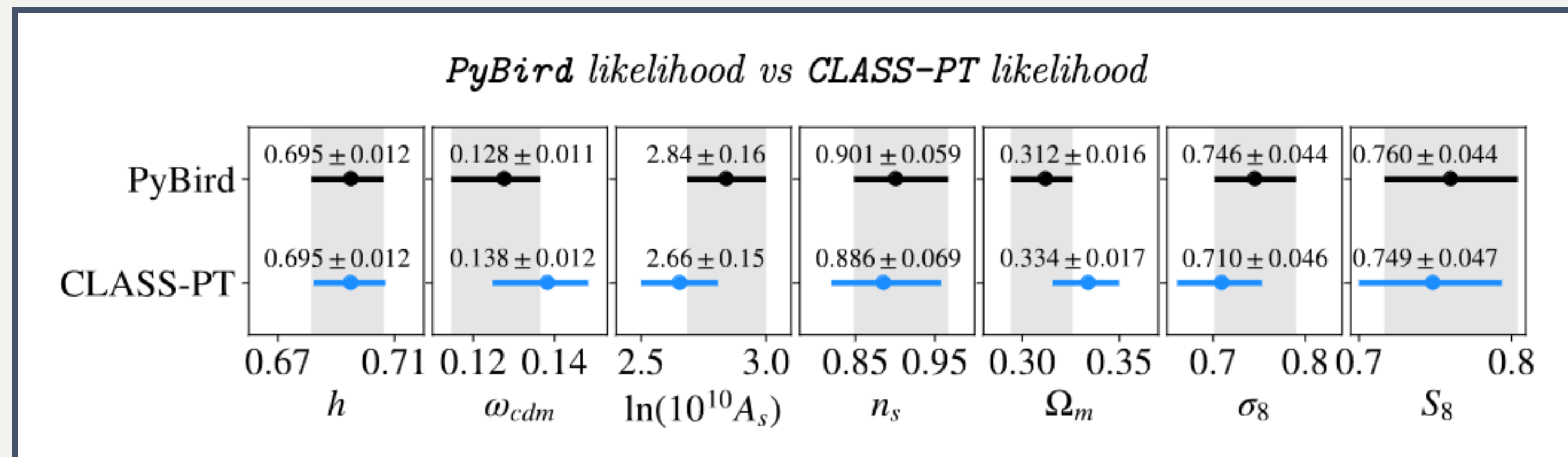
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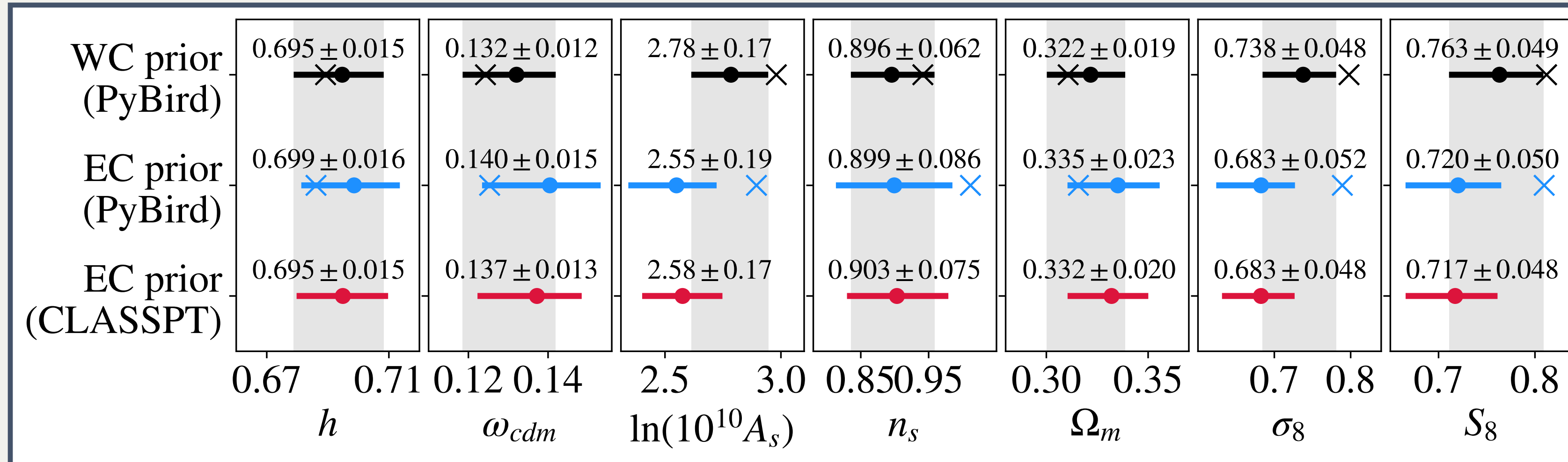


TS++ [arXiv:2208.05929]

Data, theoretical parametrizations and codes are supposed to be equivalent: what is going on?

On the consistency of EFTofLSS

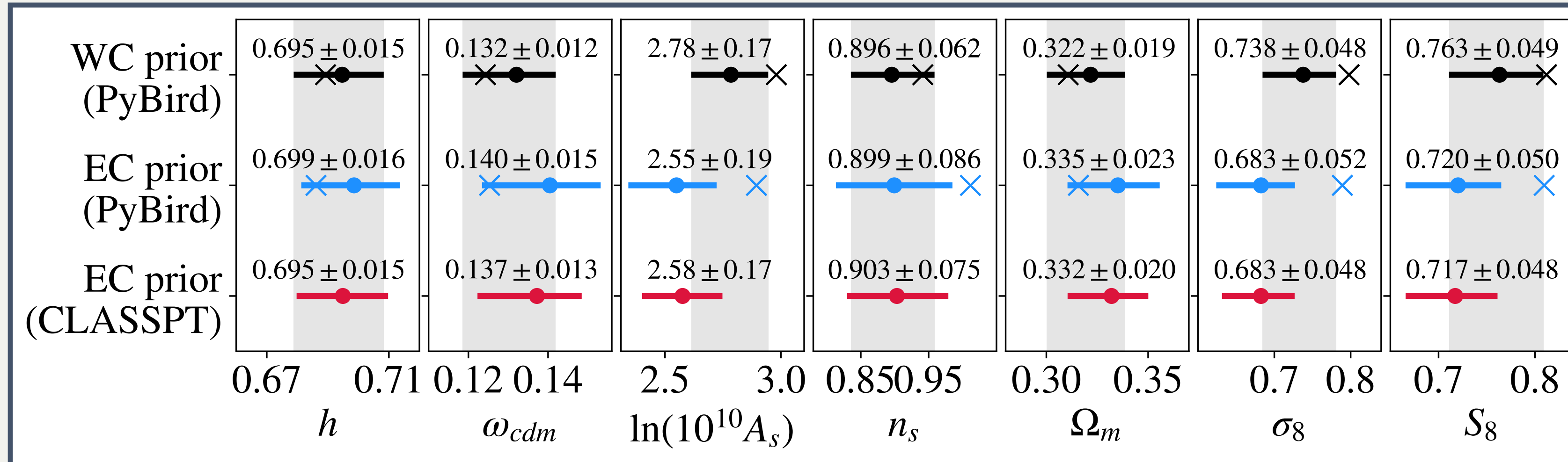
The EFT prior issue



TS++ [arXiv:2208.05929]

On the consistency of EFTofLSS

The EFT prior issue



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Prior effects

- **The prior weight effect:** if the region allowed by the prior is far from the true value of a parameter, then the posterior and bestfit will be shifted from the true value
- **The prior volume effect:** a posterior depends on the volume enclosed by the priors \implies large parameter regions are emphasized compared to smaller regions

Bayes' theorem:

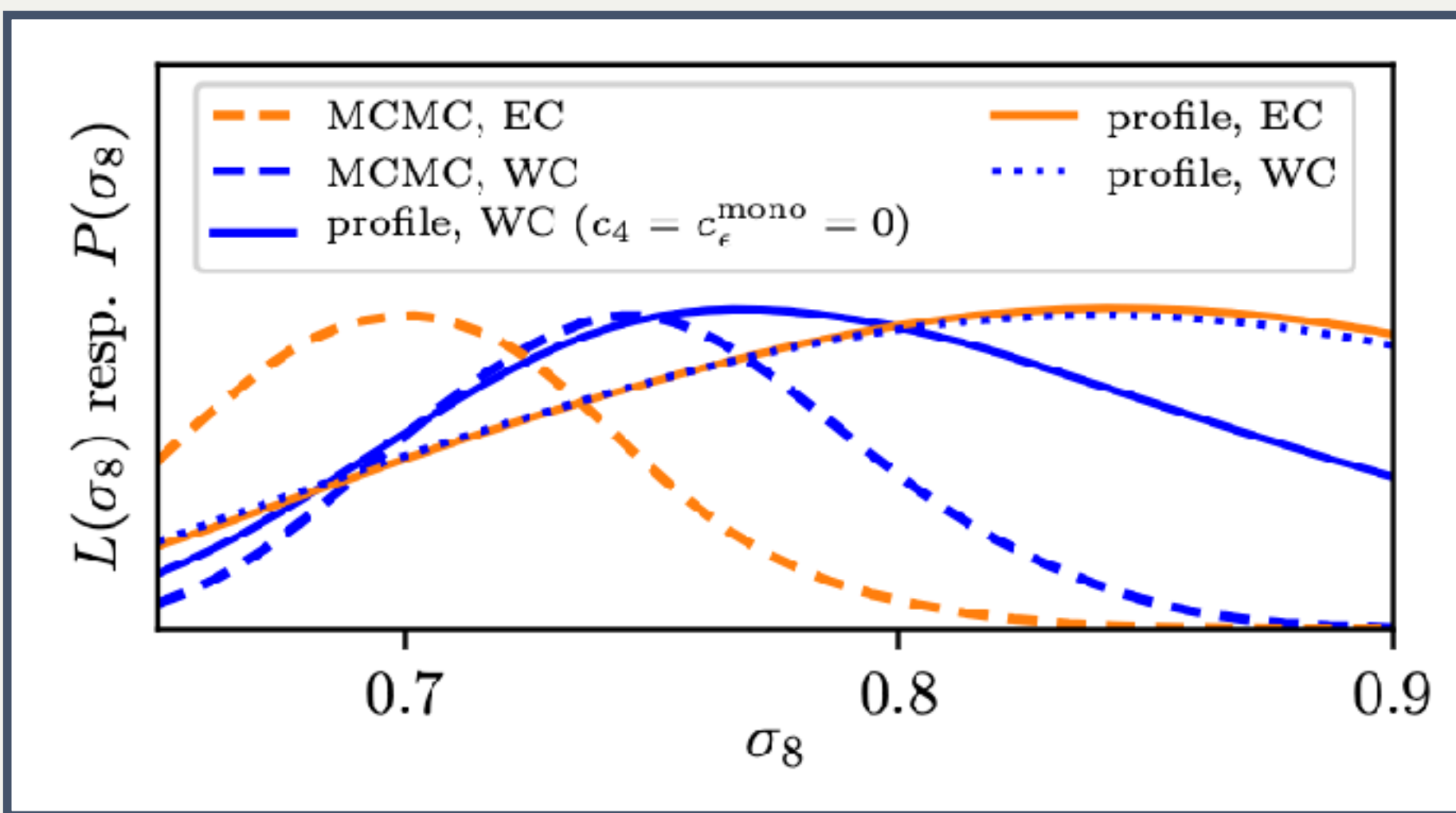
$$P \propto \mathcal{L} \times p$$

On the consistency of EFTofLSS

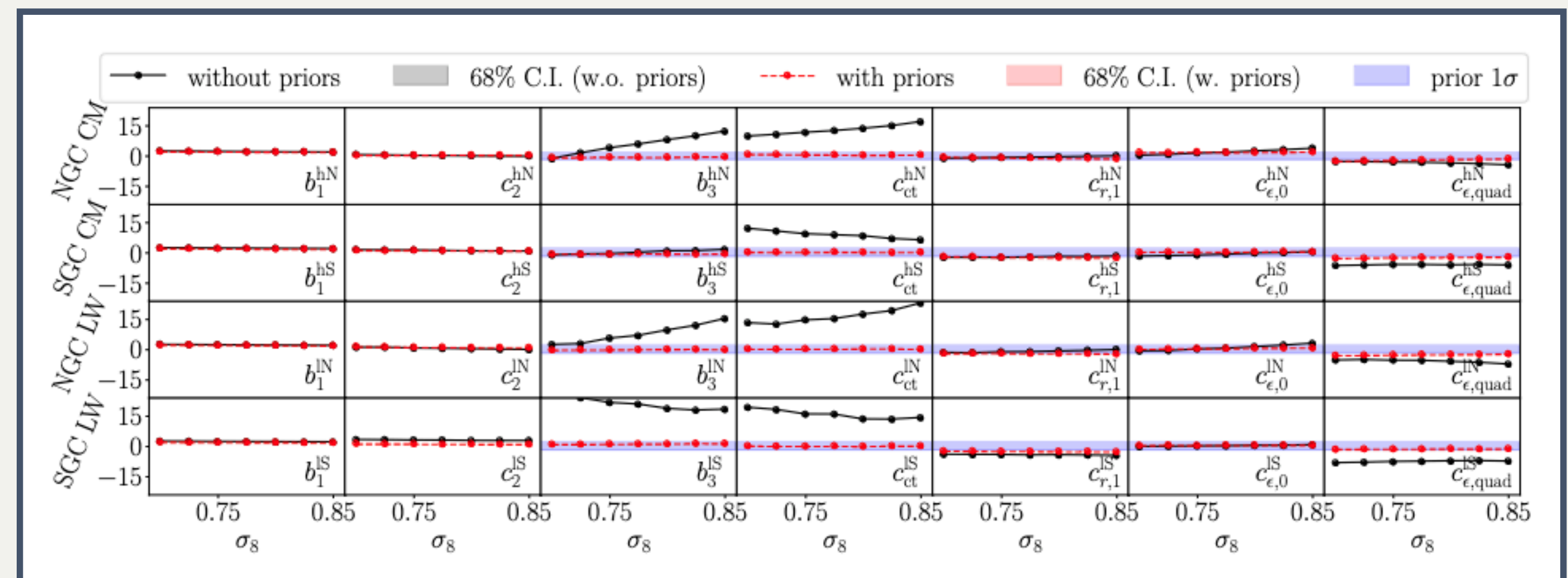
Profile likelihood

Advantage: frequentist analysis is **independent of priors** and therefore of projection effects

Disadvantage: the data prefers several EFT parameters to take on **extreme values**, possibly breaking the perturbative nature of the theory



Brinch, Herold, TS++ [arXiv:2309.04468]

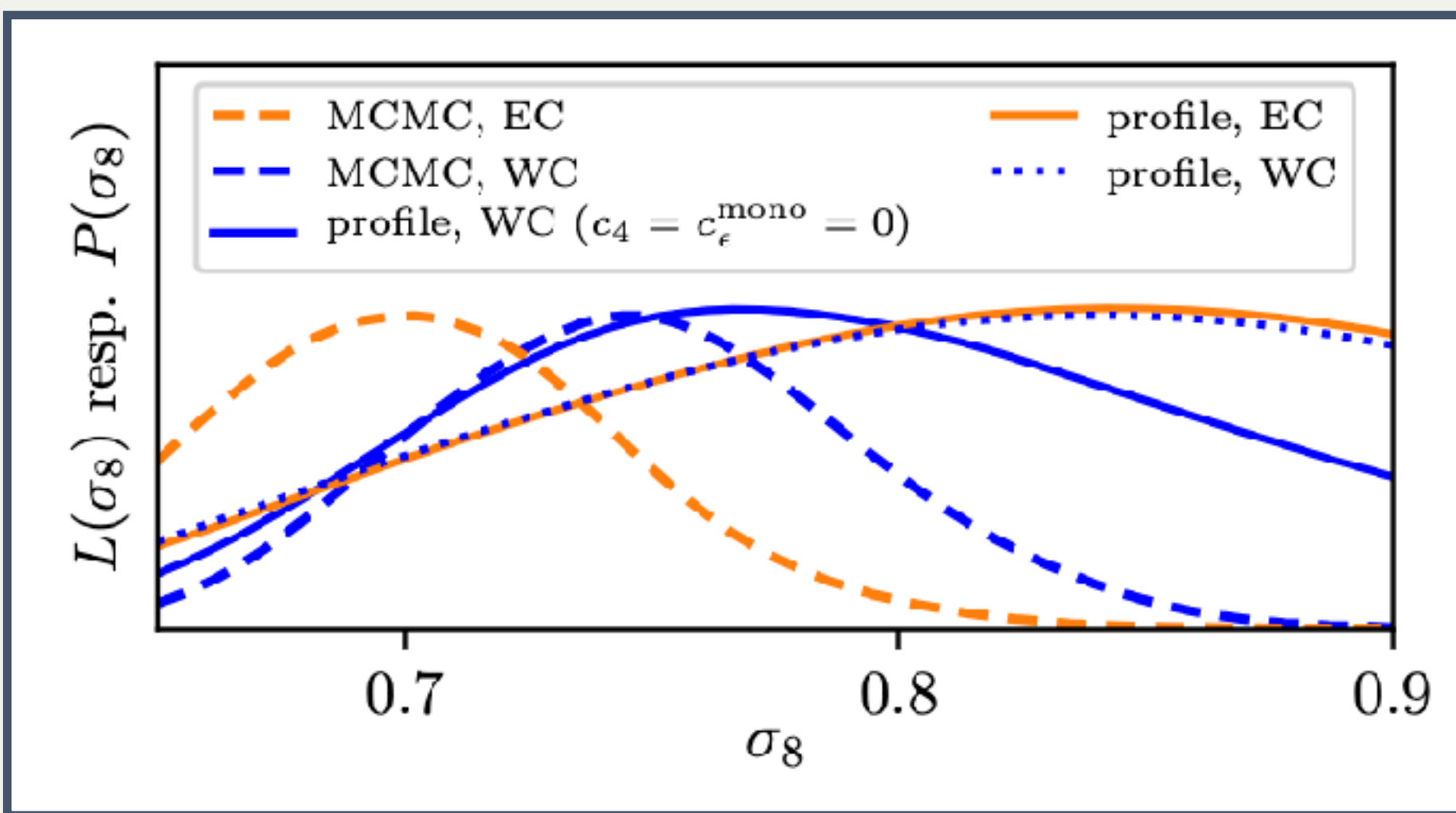


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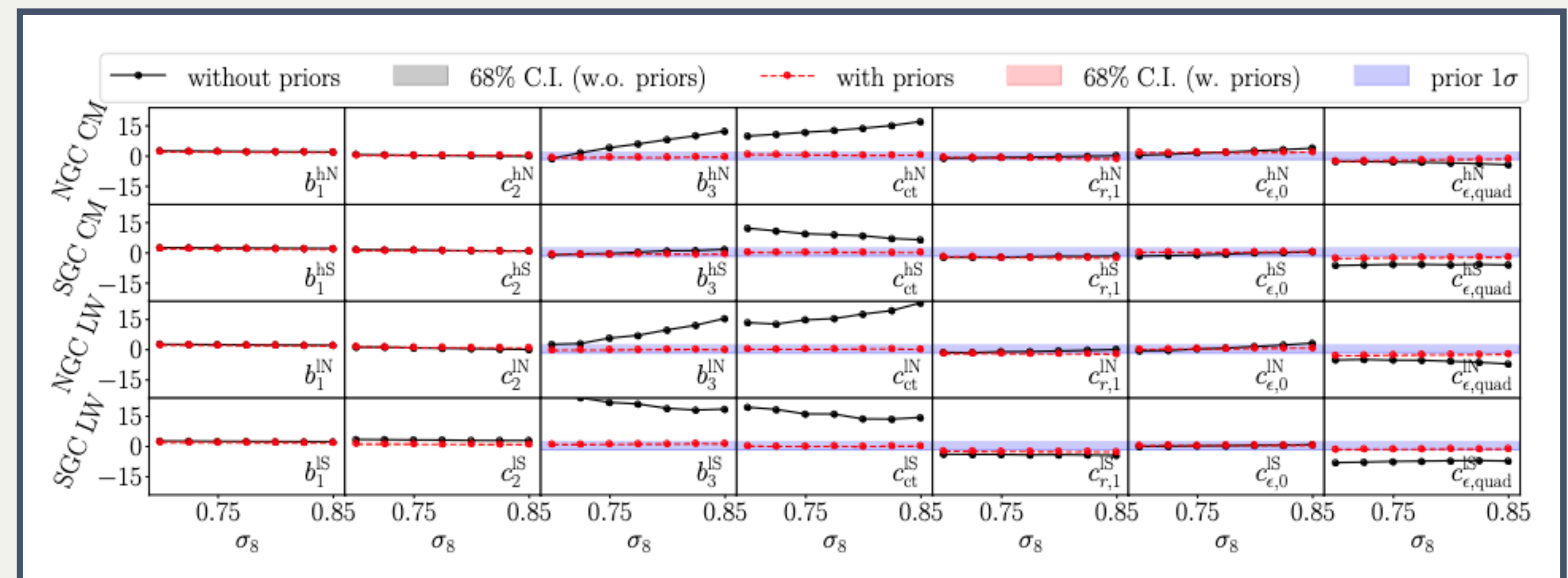
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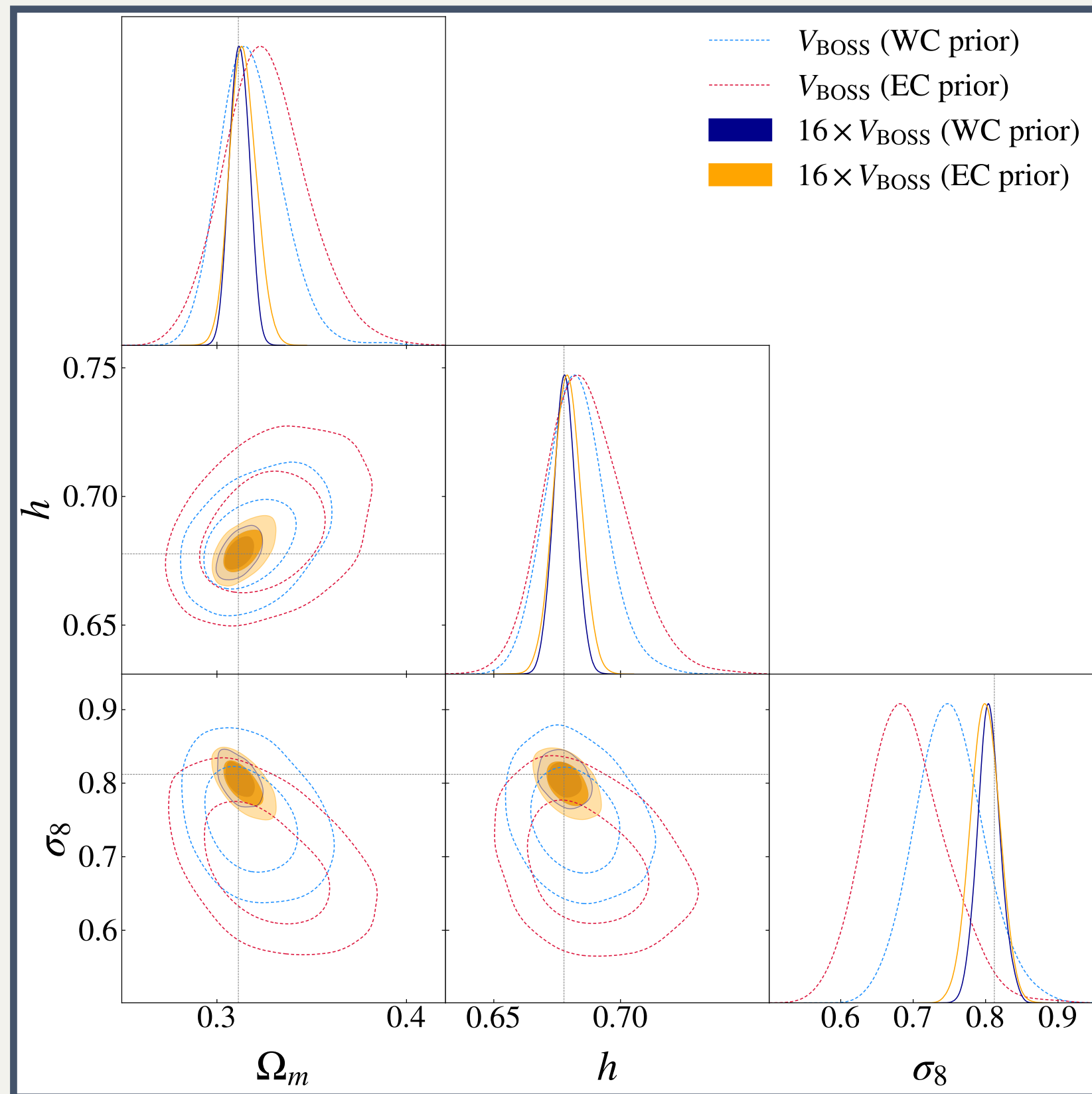
The low value of σ_8 is due to prior effects!

Brinch, Herold, TS++ [arXiv:2309.04468]



On the consistency of EFTofLSS

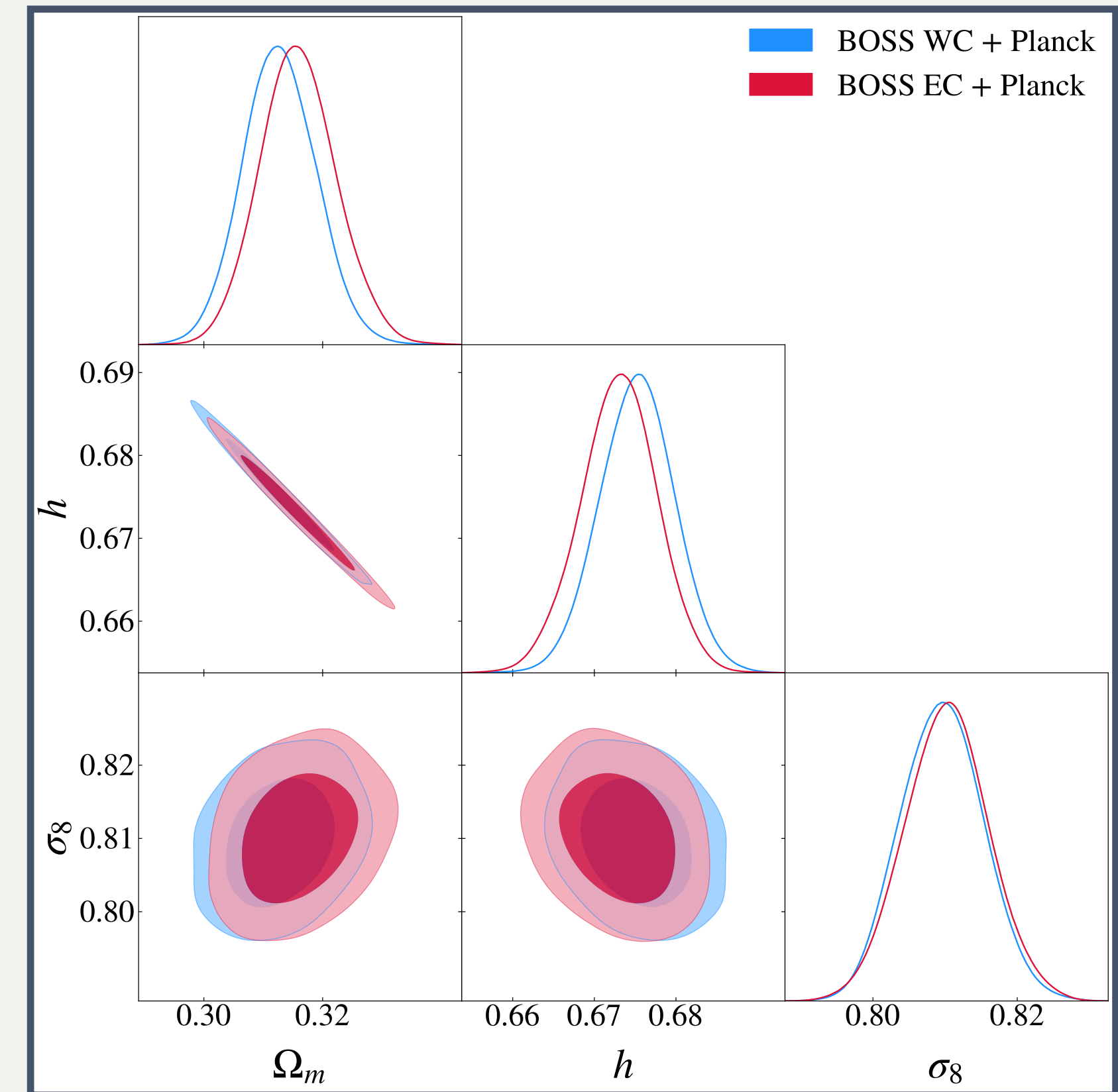
How to overcome this problem?



We find good consistency for:

- a larger volume of data (future experiments like DESI or EUCLID)
- a combination with Planck data

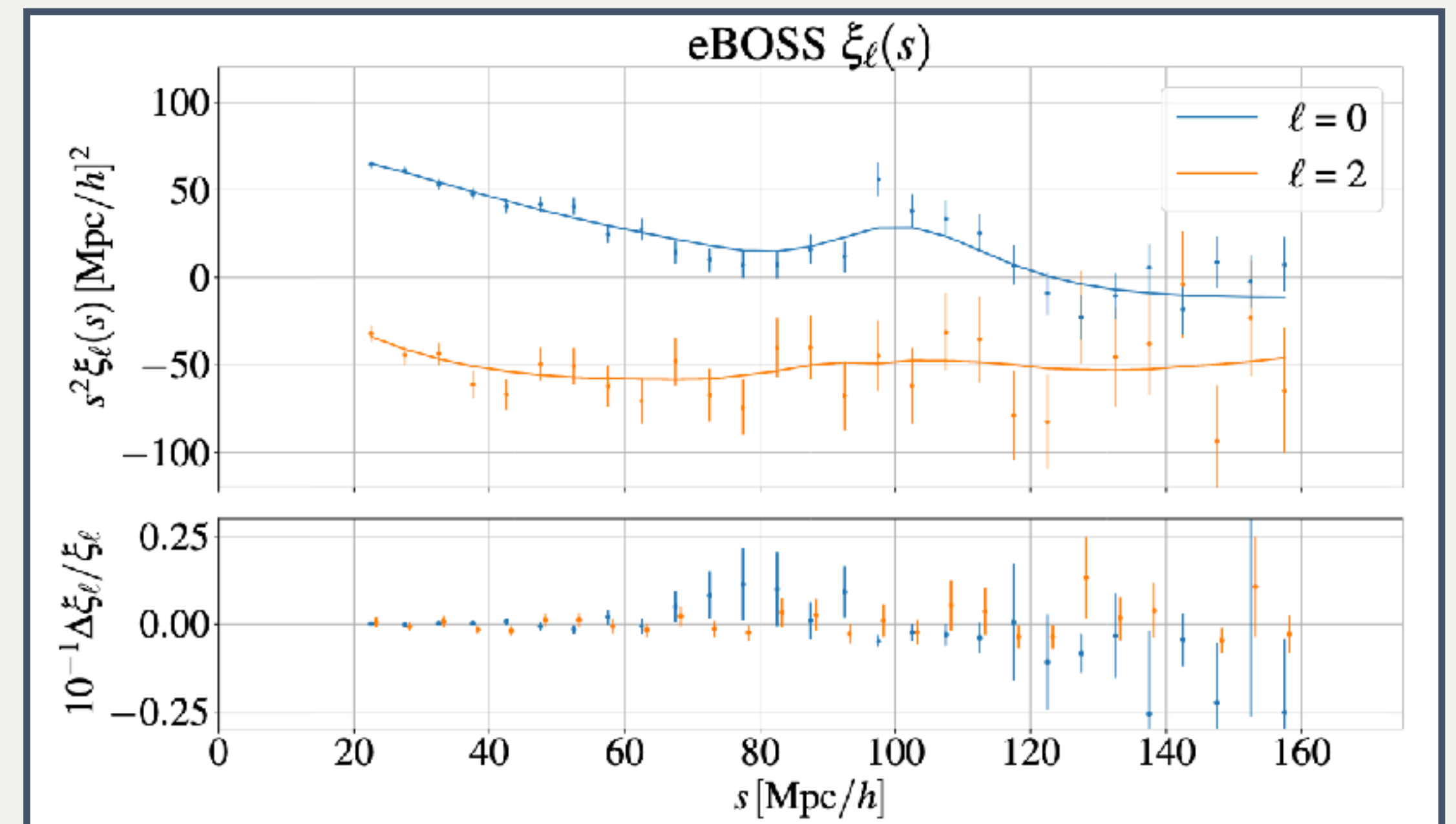
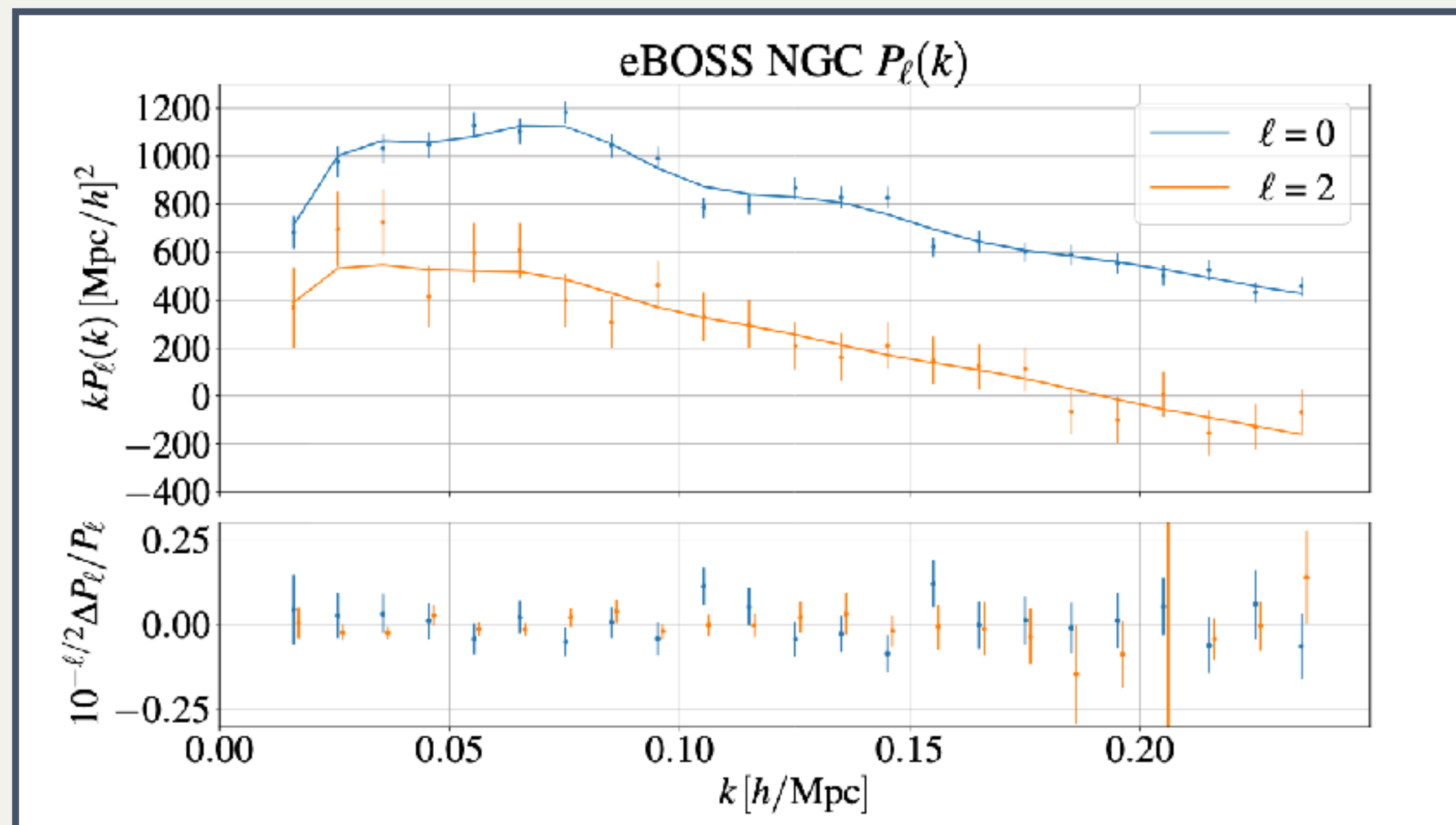
TS++ [arXiv:2208.05929]



EFTofLSS applied to eBOSS QSO data

- **343 708 quasars** selected in the redshift range $0.8 < z < 2.2$
- $z_{\text{eff}} = 1.5$
- 2 skycuts: NGC and SGC

eBOSS Collaboration
[arXiv:2007.08991]

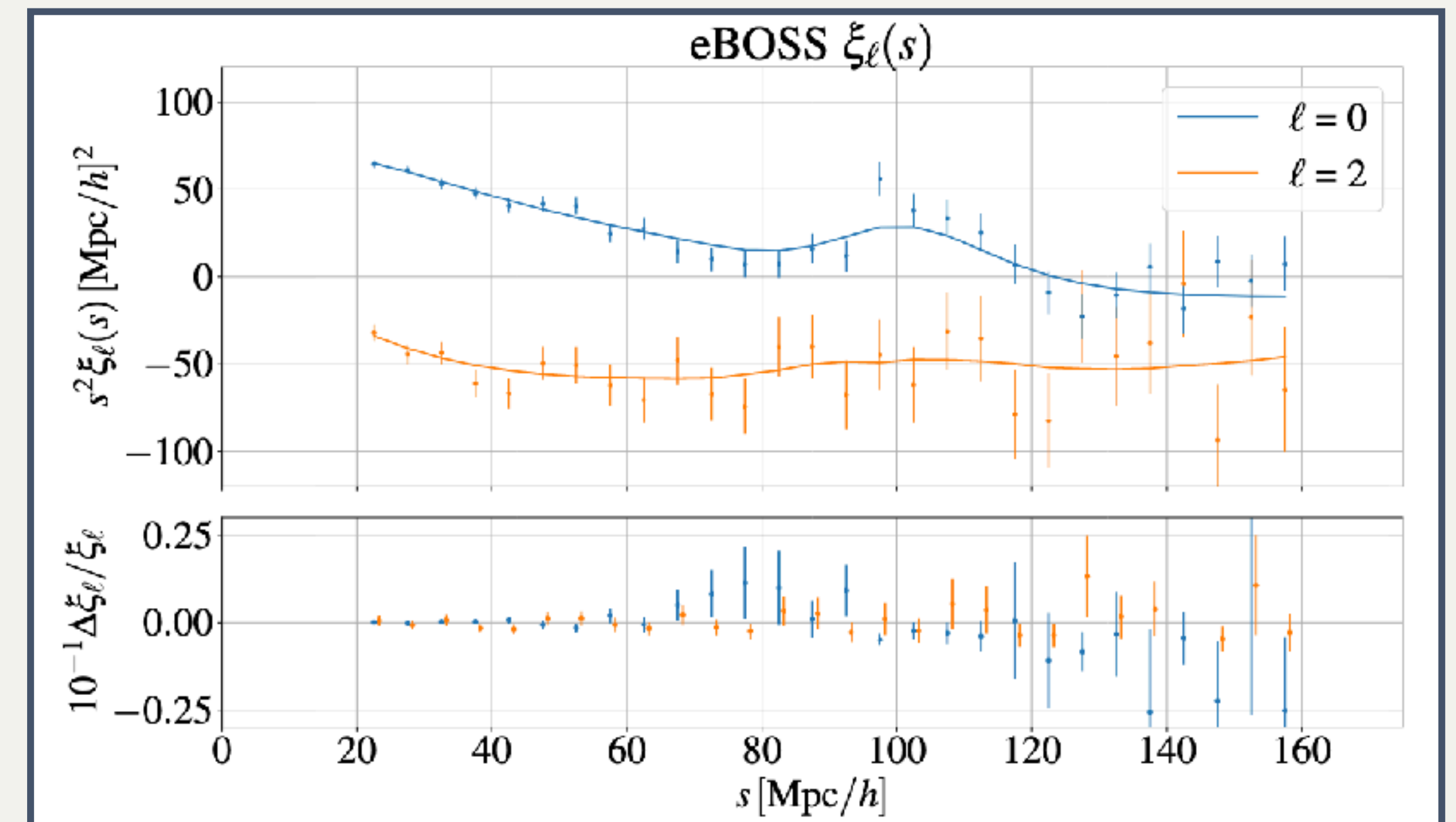
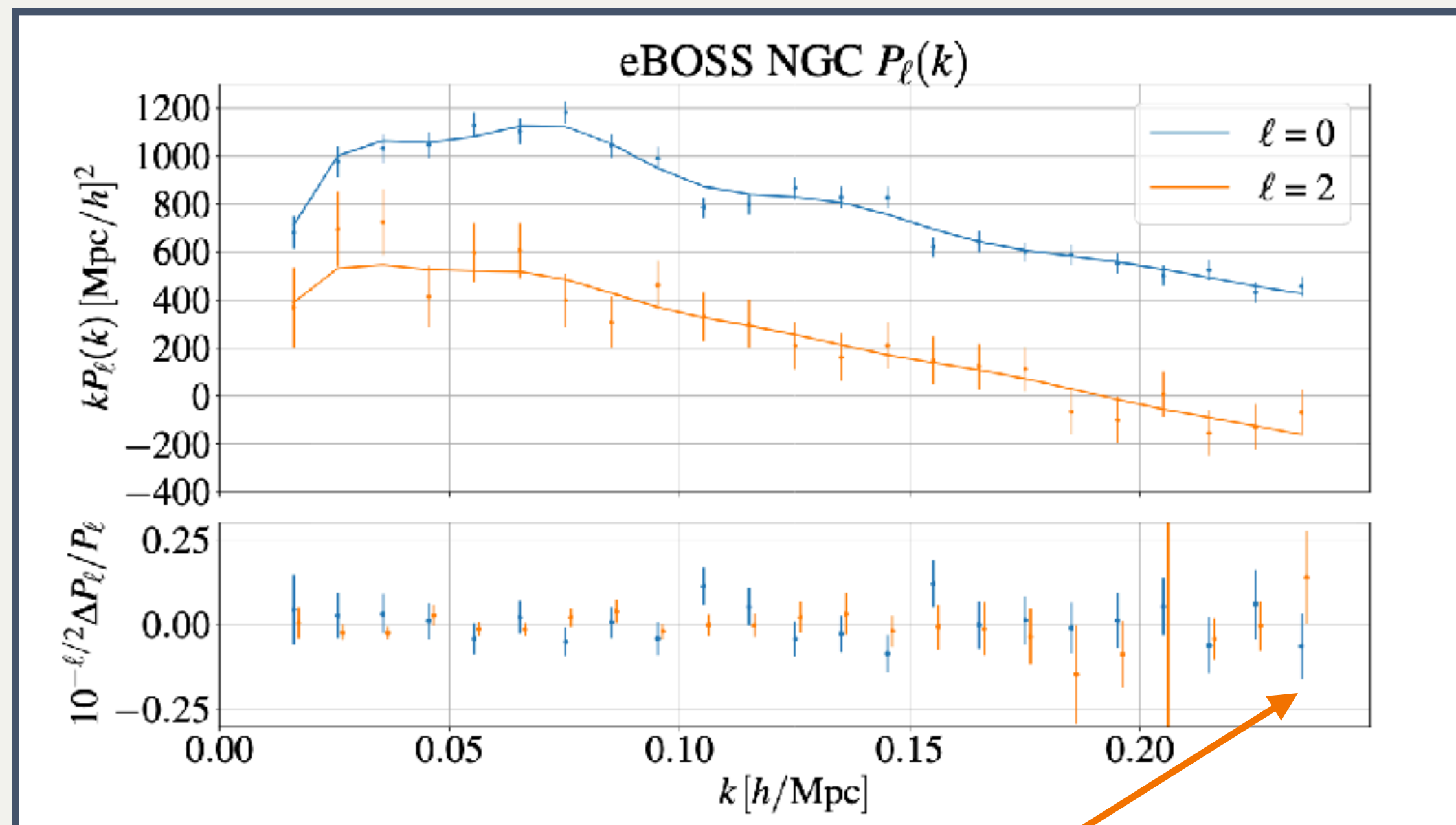


TS++ [arXiv:2210.14931]

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cut-off scale $k_{\text{max}} = 0.24 h \cdot \text{Mpc}^{-1}$

TS++ [arXiv:2210.14931]

Determination of the cut-off scale k_{\max} of the one-loop prediction

The next-to-next-to-leading order (NNLO) terms

At **one-loop order**, the galaxy power spectrum reads:

$$P_g(k, \mu) = Z_1(\mu)^2 P_{11}(k) + 2Z_1(\mu) P_{11}(k) \left(c_{\text{ct}} \frac{k^2}{k_M^2} + c_{r,1} \mu^2 \frac{k^2}{k_M^2} + c_{r,2} \mu^4 \frac{k^2}{k_M^2} \right) \\ + 2 \int \frac{d^3 q}{(2\pi)^3} Z_2(\mathbf{q}, \mathbf{k} - \mathbf{q}, \mu)^2 P_{11}(|\mathbf{k} - \mathbf{q}|) P_{11}(q) + 6Z_1(\mu) P_{11}(k) \int \frac{d^3 q}{(2\pi)^3} Z_3(\mathbf{q}, -\mathbf{q}, \mathbf{k}, \mu) P_{11}(q) \\ + \frac{1}{\bar{n}_g} \left(c_{\epsilon,0} + c_{\epsilon,1} \frac{k^2}{k_M^2} + c_{\epsilon,2} f \mu^2 \frac{k^2}{k_M^2} \right),$$

One can add the **NNLO terms** (*i.e.*, the dominant two-loop terms):

$$P_{\text{NNLO}}(k, \mu) = \frac{1}{4} b_1 (c_{r,4} b_1 + c_{r,6} \mu^2) \mu^4 \frac{k^4}{k_R^4} P_{11}(k)$$

Zhang++ [arXiv:2110.07539]

If the contribution of $P_{\text{NNLO}}(k, \mu)$ becomes **too large**, the one-loop prediction is **not accurate enough** → this determines the **cut-off scale** k_{\max} of the prediction

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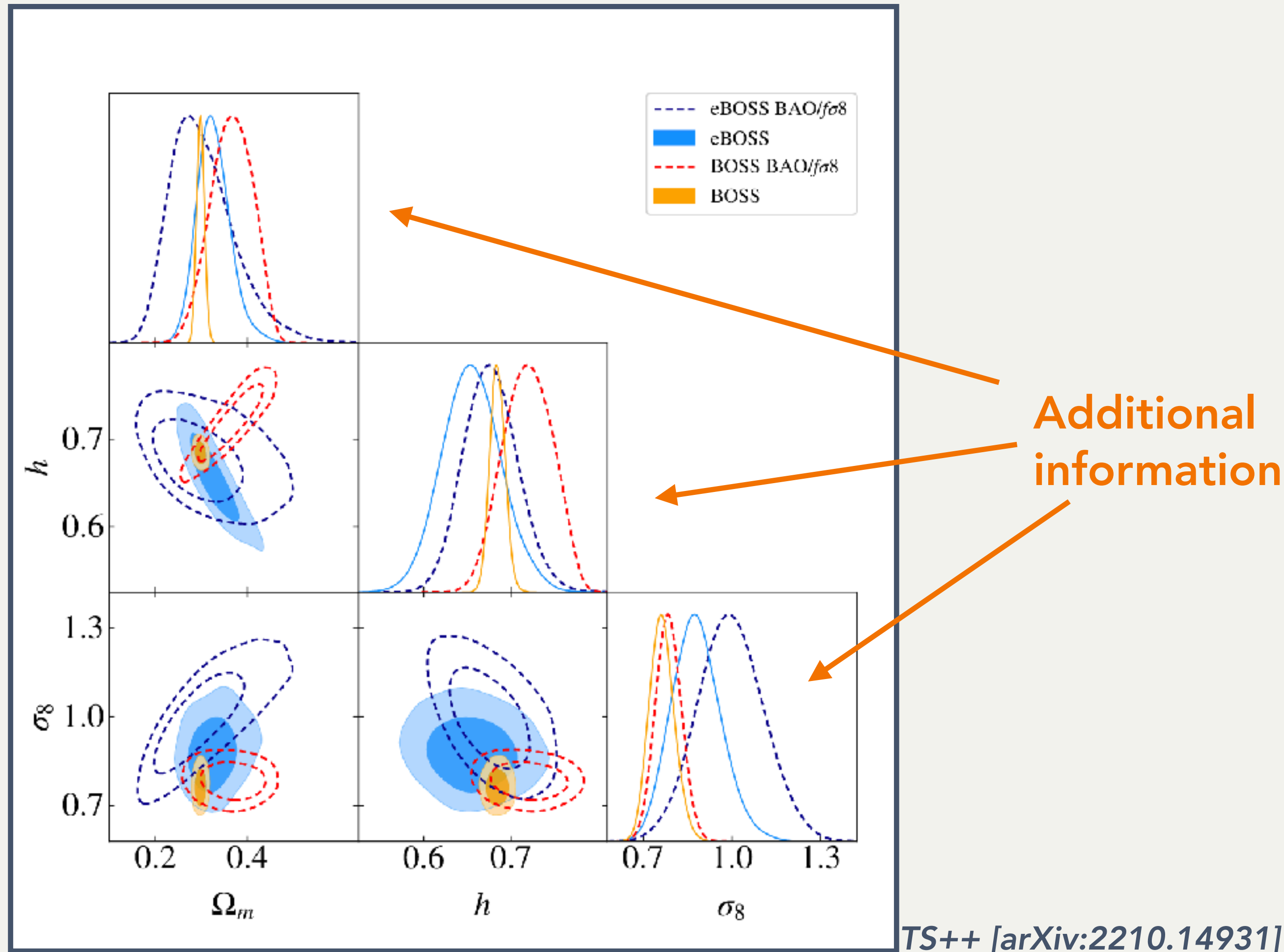
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2 new EFT parameters

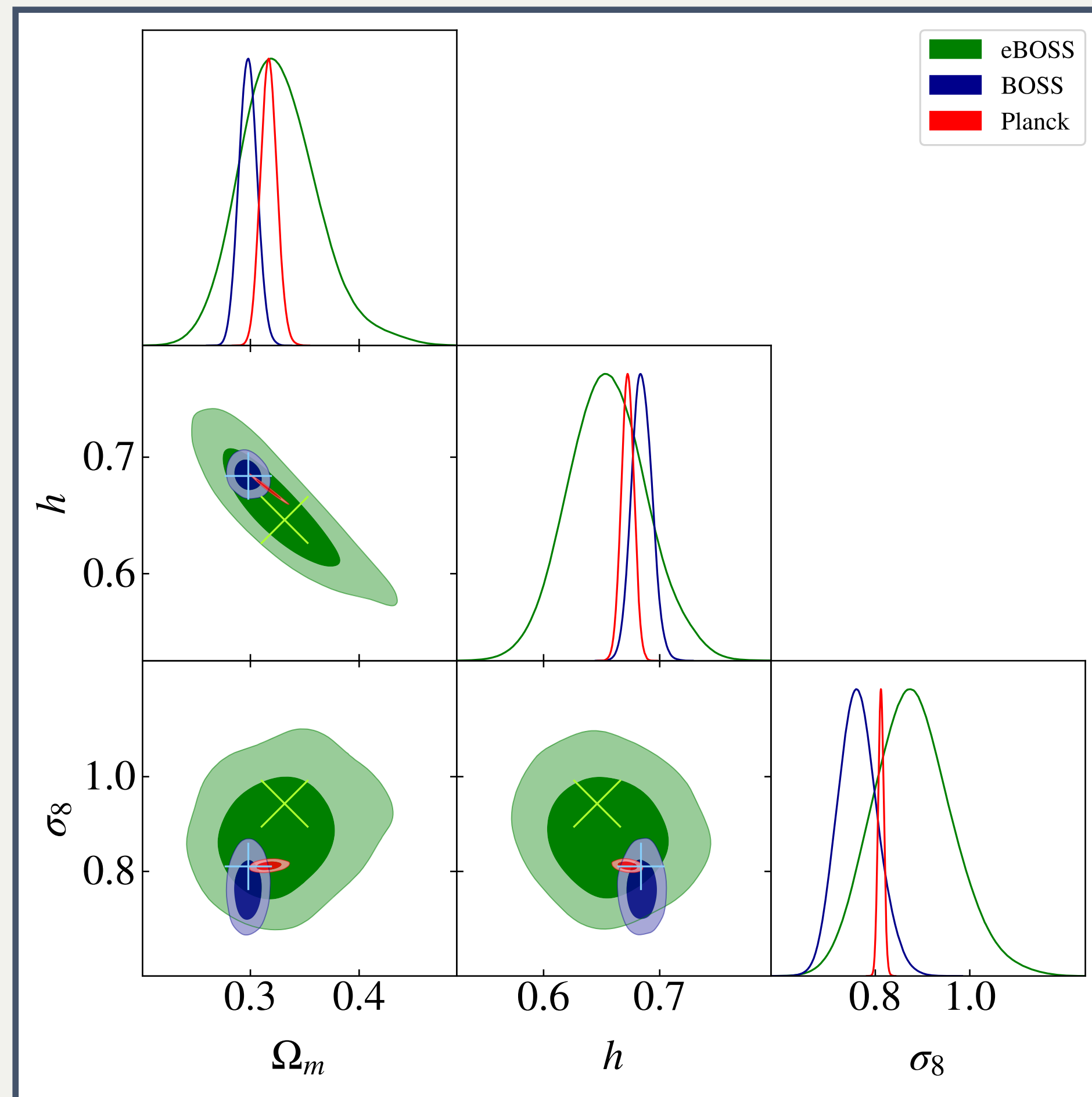
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BAO/ $f\sigma_8$ vs EFTofLSS



- For **eBOSS**, the error bars of Ω_m and σ_8 are reduced by a factor ~ 2.0 and ~ 1.3
- For **BOSS**, the error bars of Ω_m and h are reduced by a factor ~ 5.4 and ~ 3.2

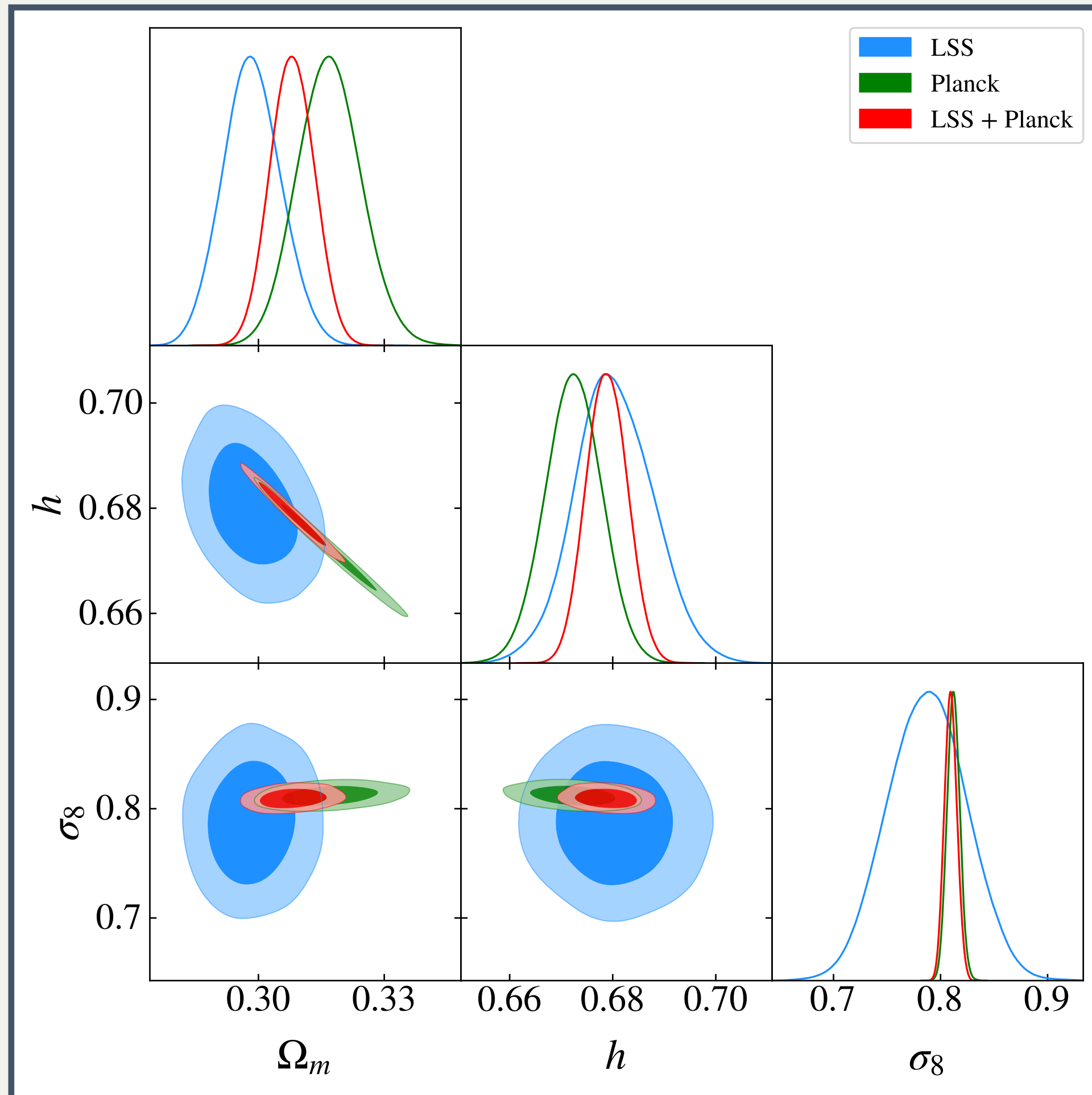
LSS data vs Planck



- **eBOSS, BOSS and Planck are consistent at $\lesssim 1.8\sigma$ on all cosmological parameters**
 - The h and σ_8 Planck values are **in-between** those of BOSS and eBOSS
- **there is no tension between Planck and BOSS/eBOSS**

TS++ [arXiv:2210.14931]

LSS data combined with Planck



LSS: eBOSS + BOSS + ext-BAO + Pantheon

(Uncalibrated Supernovae)

- **Compared to Planck alone**, the constraints on Ω_m and h are **improved by $\sim 30\%$**
- σ_8 and A_s are not significantly impacted

TS++ [arXiv:2210.14931]

Extensions to Λ CDM: total neutrino mass $\sum m_\nu$

- The LSS constraint derived in this work is **only** $\sim 10\%$ **weaker than the Planck constraint** ($\sum m_\nu < 0.241eV$)
- The EFT analysis **significantly improves the constraints** on $\sum m_\nu$ (by a factor of ~ 18) over the conventional BAO/ $f\sigma_8$ analysis ($\sum m_\nu < 4.84eV$)
- This analysis **disfavors the inverse hierarchy** at $\sim 2.2\sigma$ & is **competitive to the Lyman- α constraints**

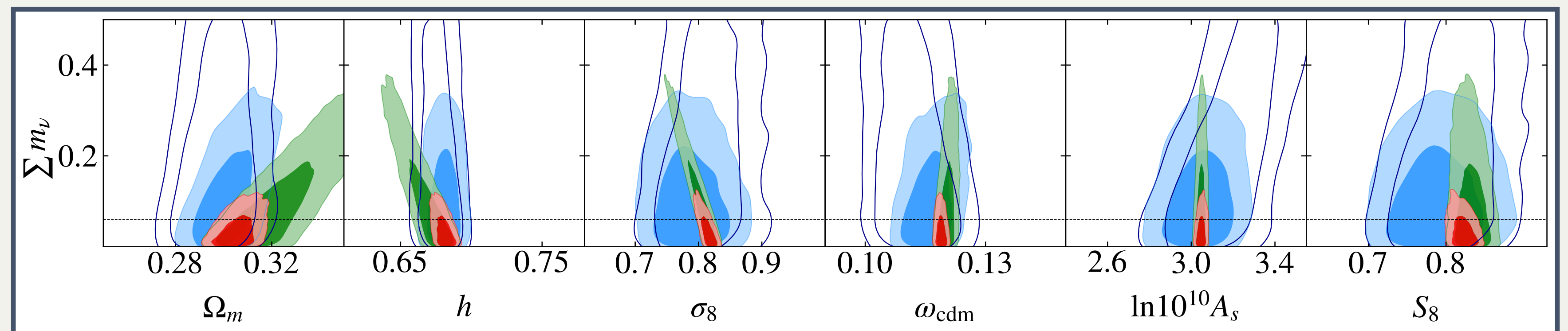
Palanque-Delabrouille++ [arXiv:1911.09073]



TS++ [arXiv:2210.14931]

LSS:
 $\sum m_\nu < 0.274eV$

LSS+Planck:
 $\sum m_\nu < 0.093eV$



Conclusion

- The EFTofLSS is a novel method that provides an **accurate description of LSS data (up to mildly non linear scales) at a controlled precision**
- Constraints from LSS data are **competitive with CMB data** and their combination **improves over Planck alone**
- EFTofLSS allows to highlight that **there is no tension** between current BOSS/eBOSS data and Planck data (but not in tension with weak lensing neither)
- Data are consistent with Λ CDM at $\lesssim 1.3\sigma \rightarrow$ Strong constraints on canonical extensions to Λ CDM
e.g. LSS+Planck: $\sum m_\nu < 0.093eV$
- Does EFTofLSS provide **interesting constraints on non-canonical extensions** to the Λ CDM model ?